



eHeroes 2nd General Meeting  
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# RESIK MEASUREMENTS OF PARTICLE BACKGROUND RADIATION & NON- SOLAR EMISSION

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# outline

- ▣ Construction & placement of RESIK detectors
- ▣ Signals useful to study particle environment
- ▣ Calibration of RESIK particle signal against particle spectrometer MKL
- ▣ Conclusions and further work-
- ▣ Possible collaborations

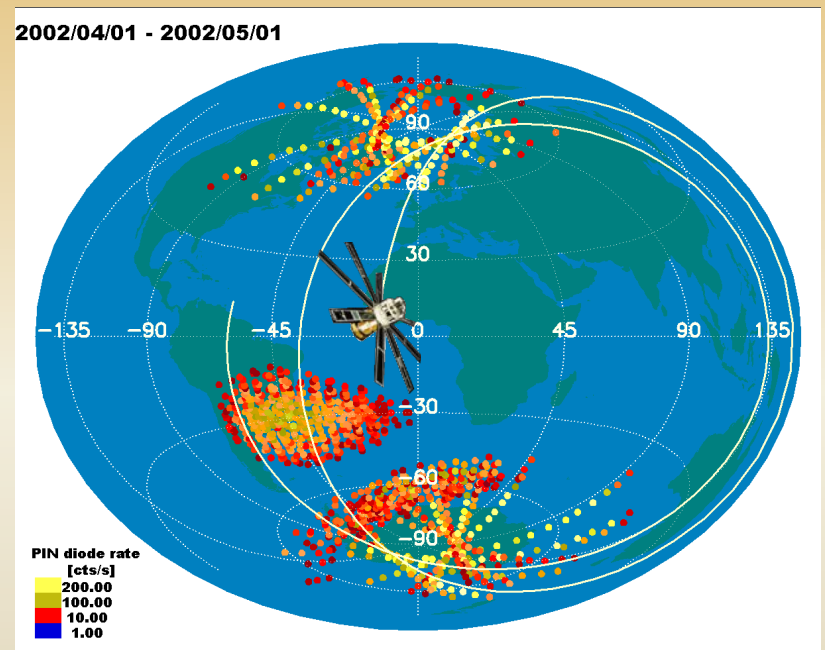
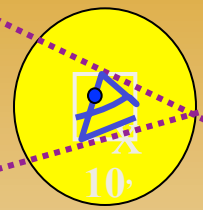


# CORONAS-F operation

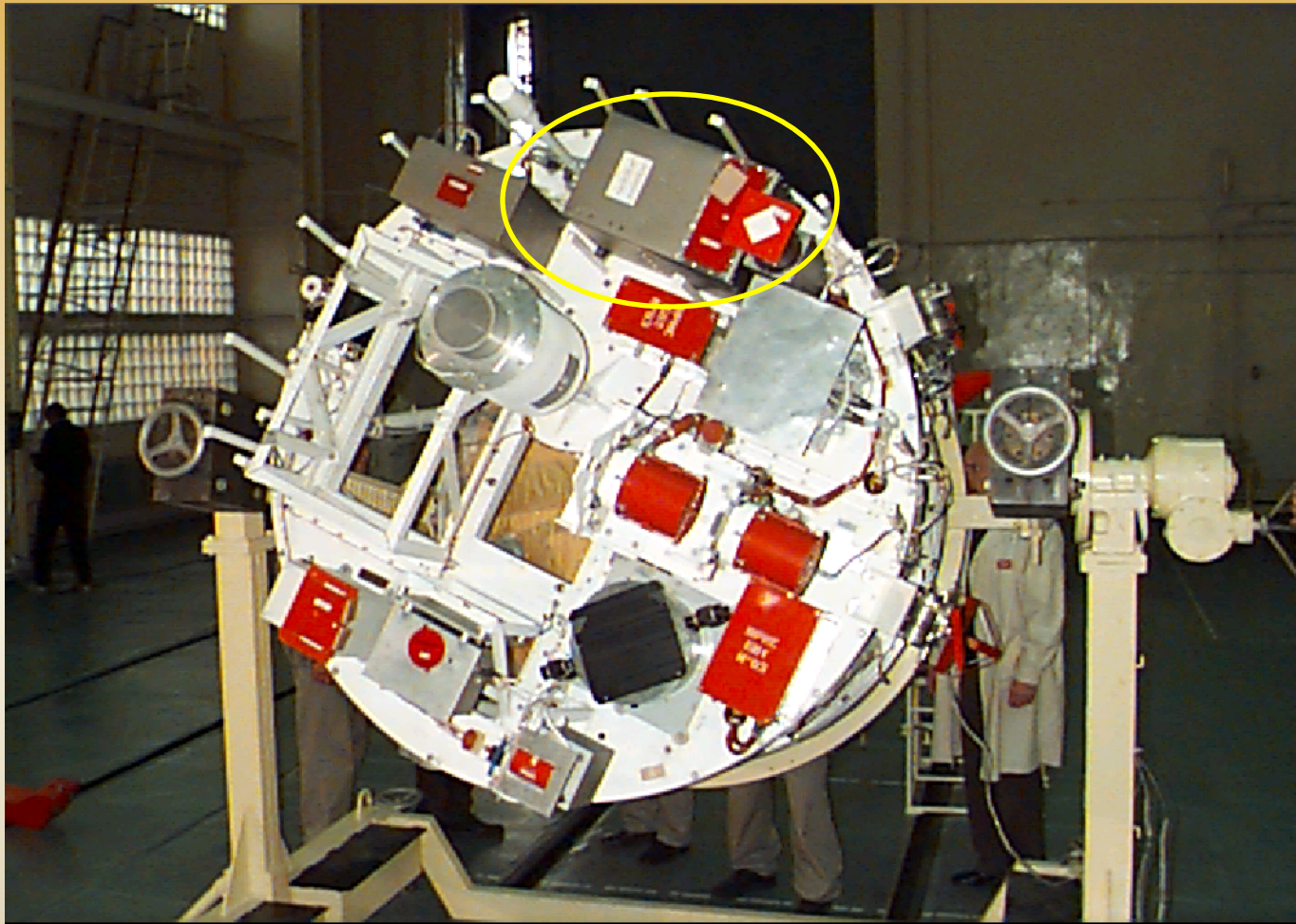


SS-14 Cyclone

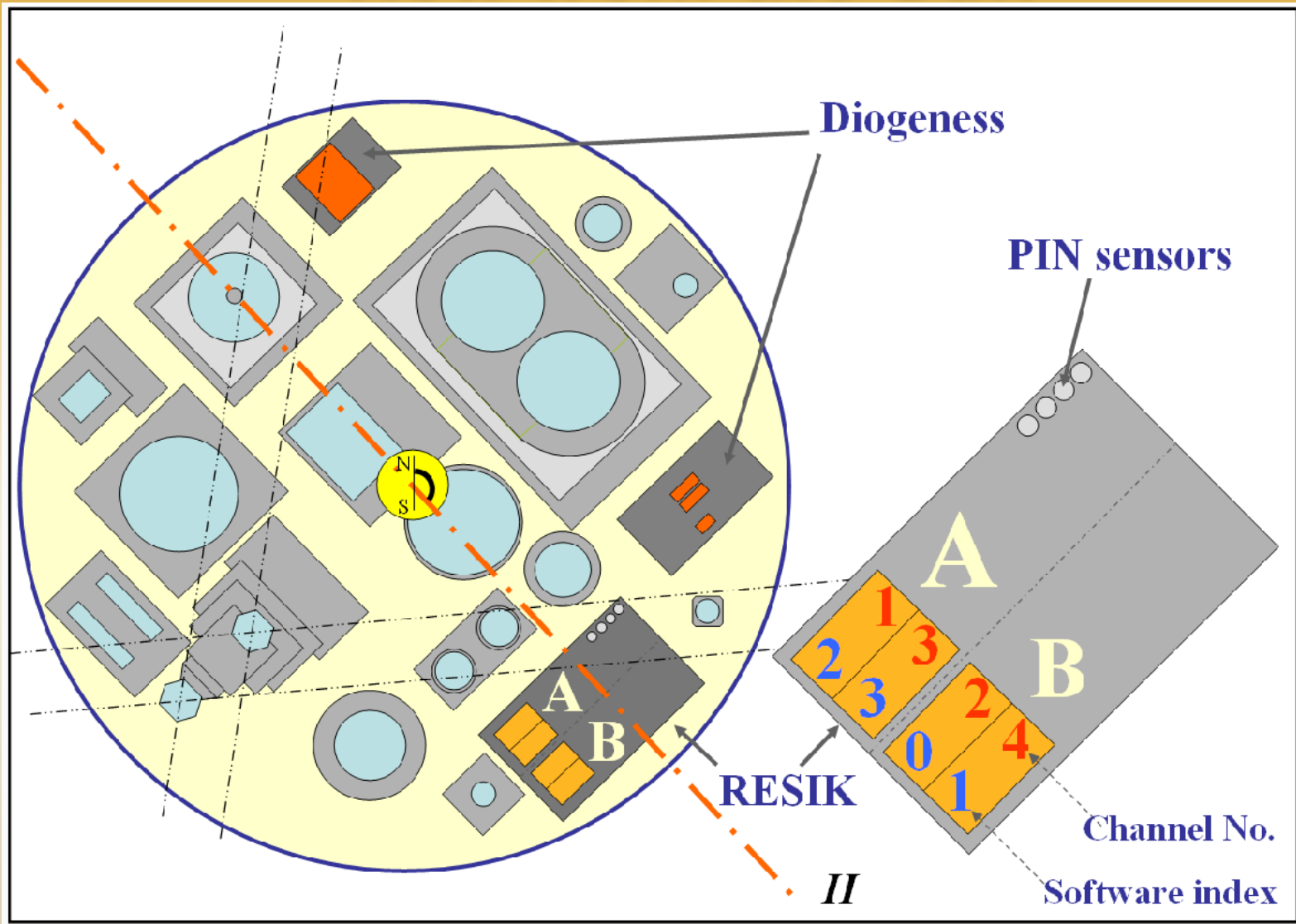
31 July 2001-Dec. 2006  
polar orbit, 95min,  
~500 km  
semi-Sun-synchronous



# RESIK on CORONAS-F



# CORONAS-F payload view from the Sun



# RESIK detectors

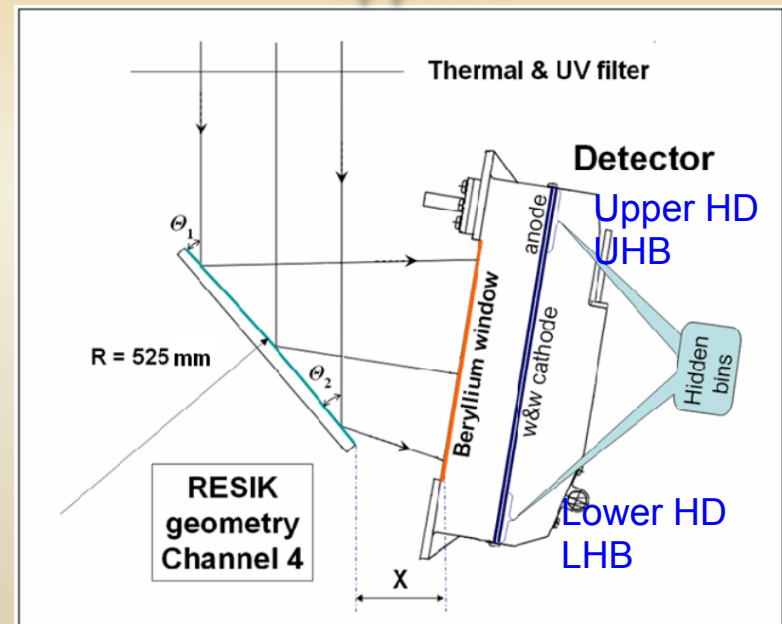


This was a great idea to put 4 PIN detectors inside RESIK the detectors were of the type NTE3033 Infrared Photodiode 7 x 8 mm behind 2 x 80 micron Al foils diameter 14 mm. This part was covered by MLI

The main spectroscopy 4 detectors contained parts of anodes (hidden bins HB) upper (better screened and lower). Every signal recorded on anode has been pulse –height analysed into 32 bin spectrum. Its energy range depends on HV applied



Titanium cover with MLI removed



# RESIK signals with particle signatures



- ▣ PIN detectors
  - 4 independent „identical” detectors (4 x20 mln data)
- ▣ RESIK PHA
  - 2 independent detectors with regulable HV
  - 4 independent PHA 32 bin data (4 x32 x 20mln/4 since commuted)
- ▣ Satellite conditions
  - Sat day/night, longest night 25 min →0 for long days
  - SAA, polar ovals transition, geomagnetic storms
- ▣ RESIK spectral hidden bins
  - UHB,LHB x 2 mln spectral measurements
- ▣ Total of 40 mln points, up to 40 at the same time



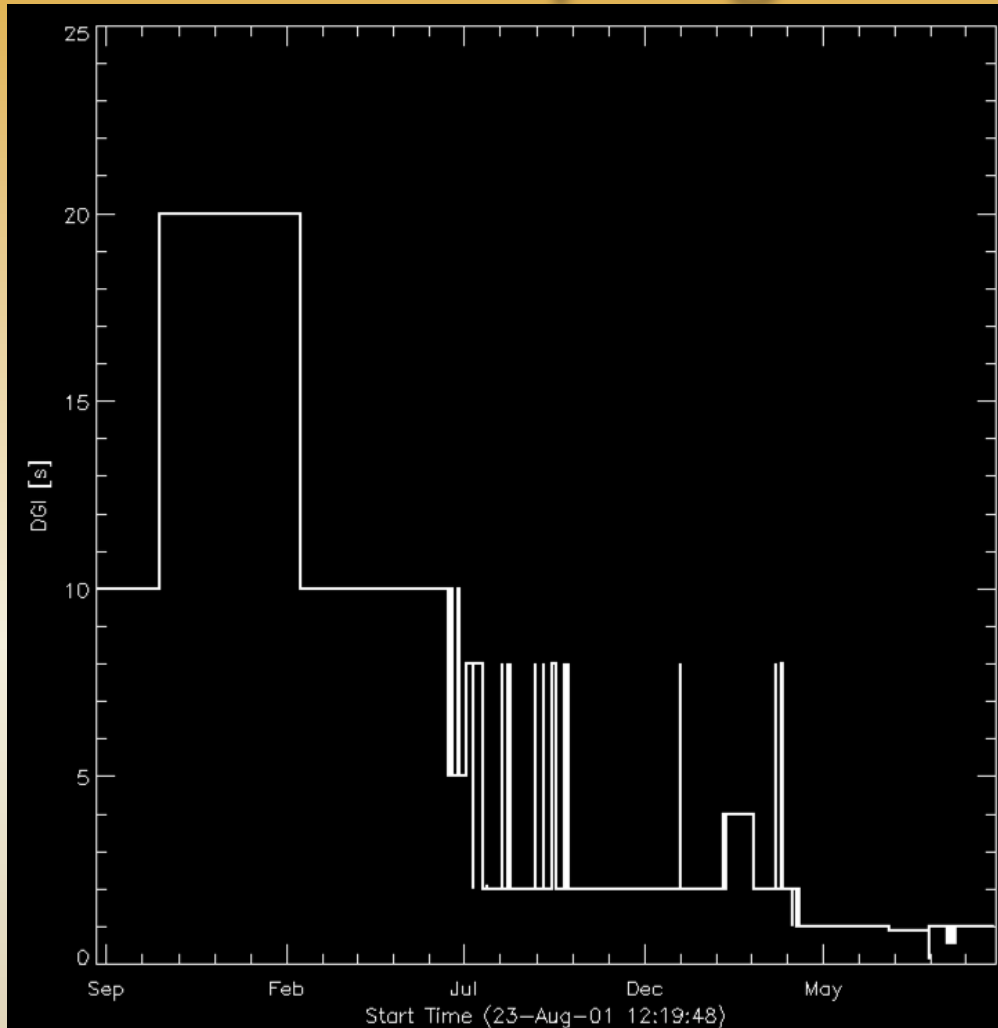
# RESIK important periods of operation



- Early „learning“ operations between 2001-Aug-24 12:19:48 & 2001-Oct-07 00:59:34
  - Const HV ON
  - Const DGI=10s
- Const DGI=20 s 2001-Oct. & 2002 Feb.
- ADS Tweaking 2002-Feb. & 2002 Jul.
  - Constant DGI = 10s
- DGI =8 or 2s 2002-Jul. & 2003 Apr.
- DGI =1 s 2003-Apr & 2003 Sep.

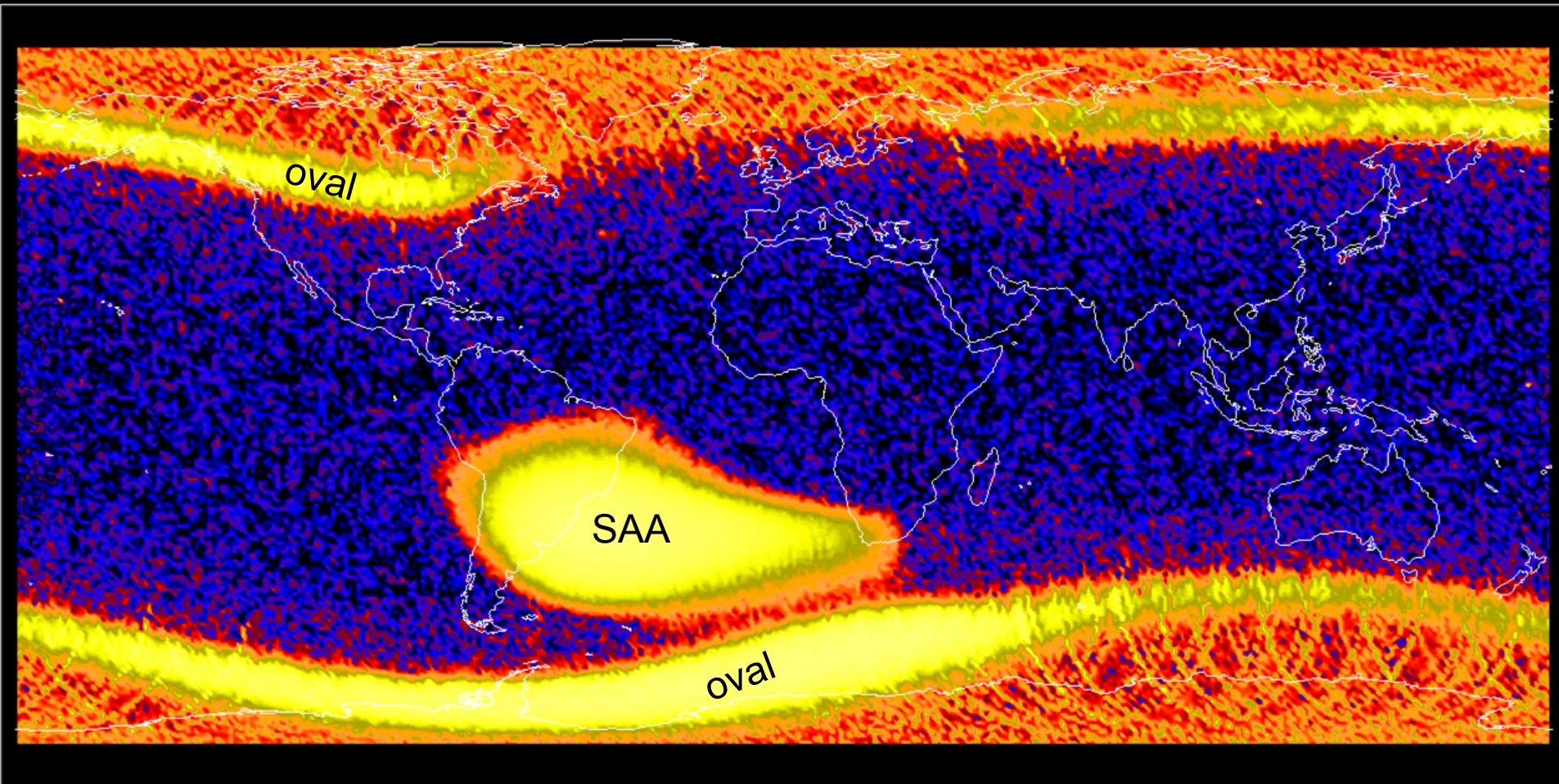


# RESIK DGI- sampling time plot



- Sampling of conditions along the orbit are  $\sim$ DGI
- 8km for 1s
- 160 km for 20s

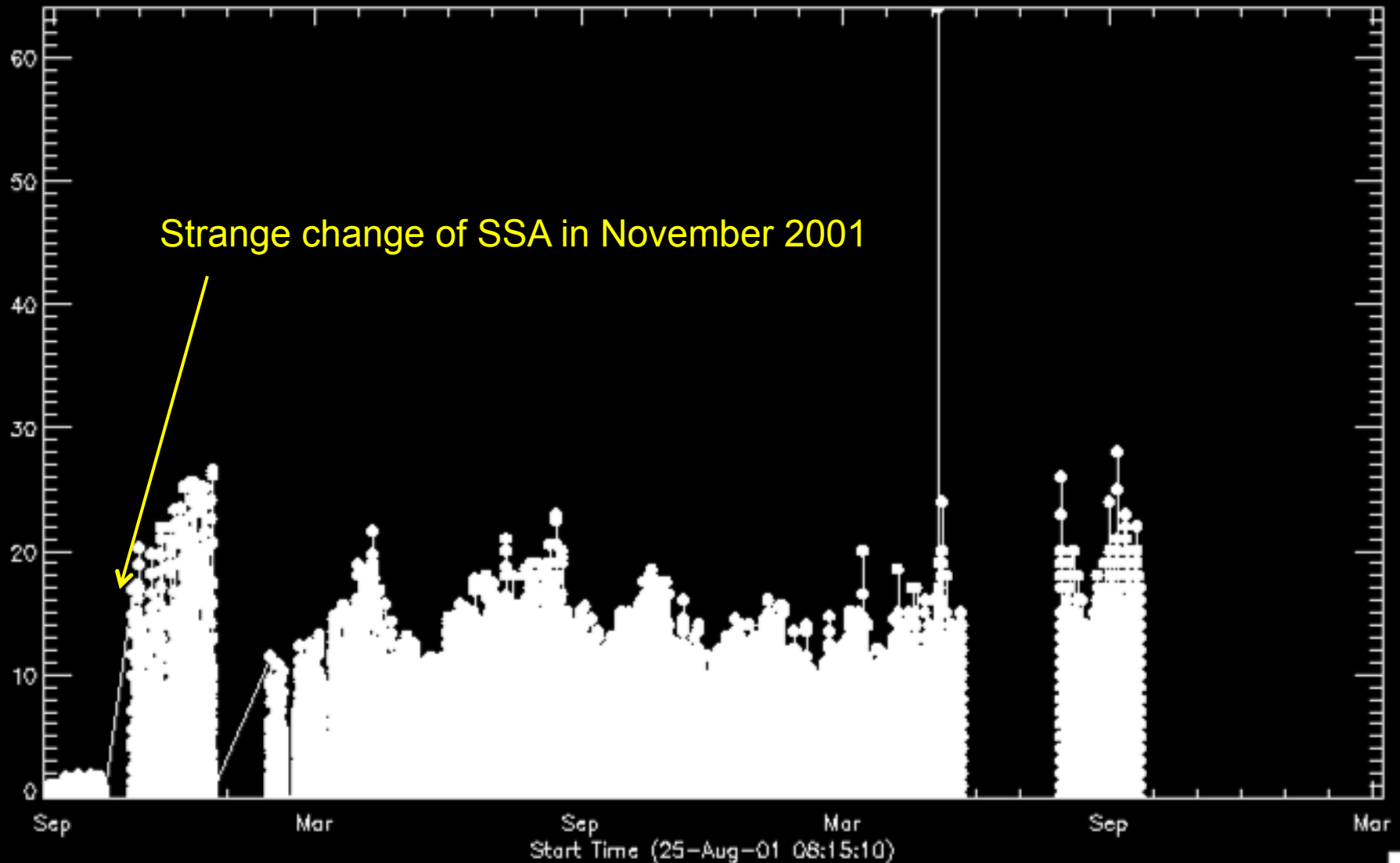
PIN sum:  $3e7$  events  
mission long 1 x1 deg



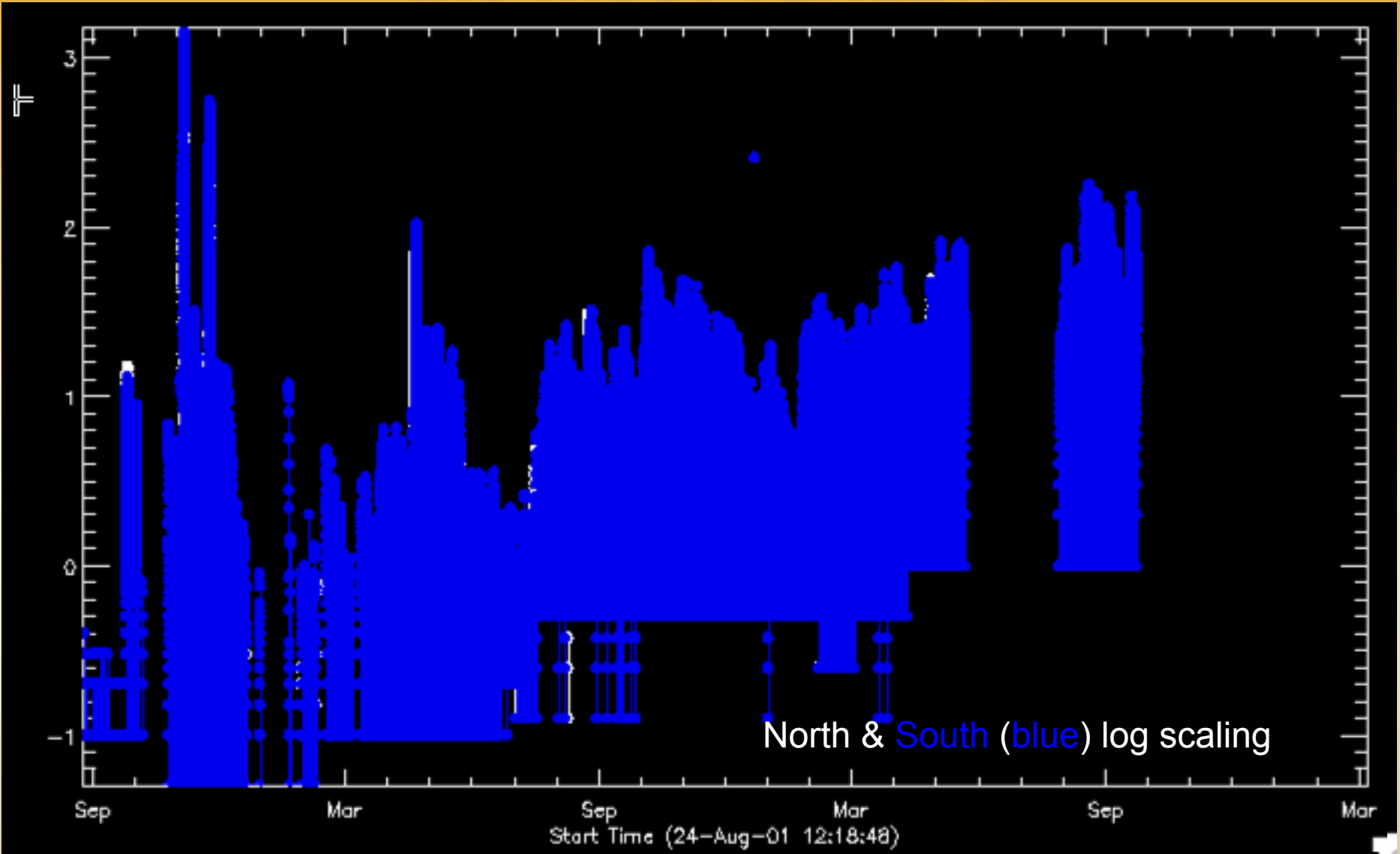
# Rates in SAA



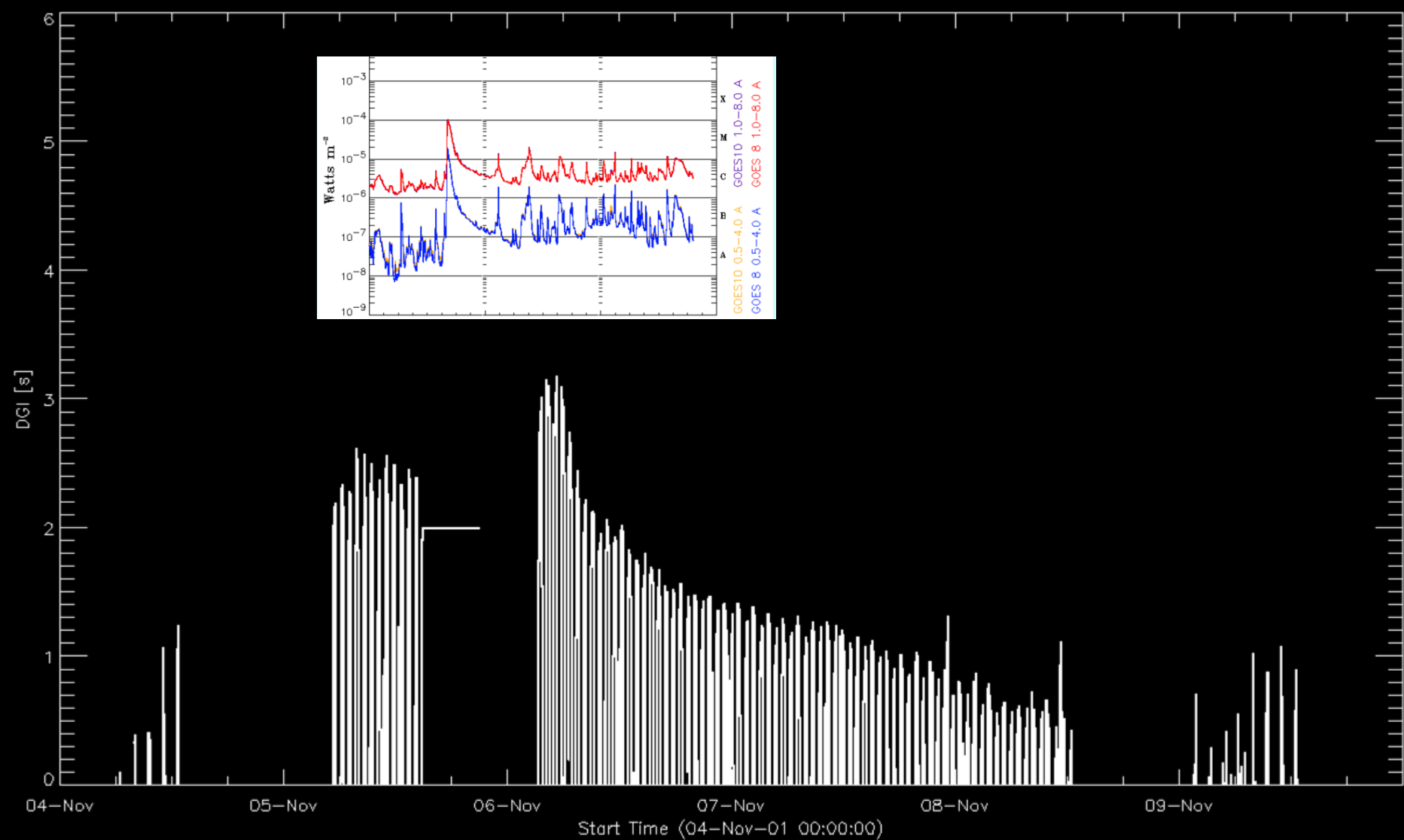
Strange change of SSA in November 2001



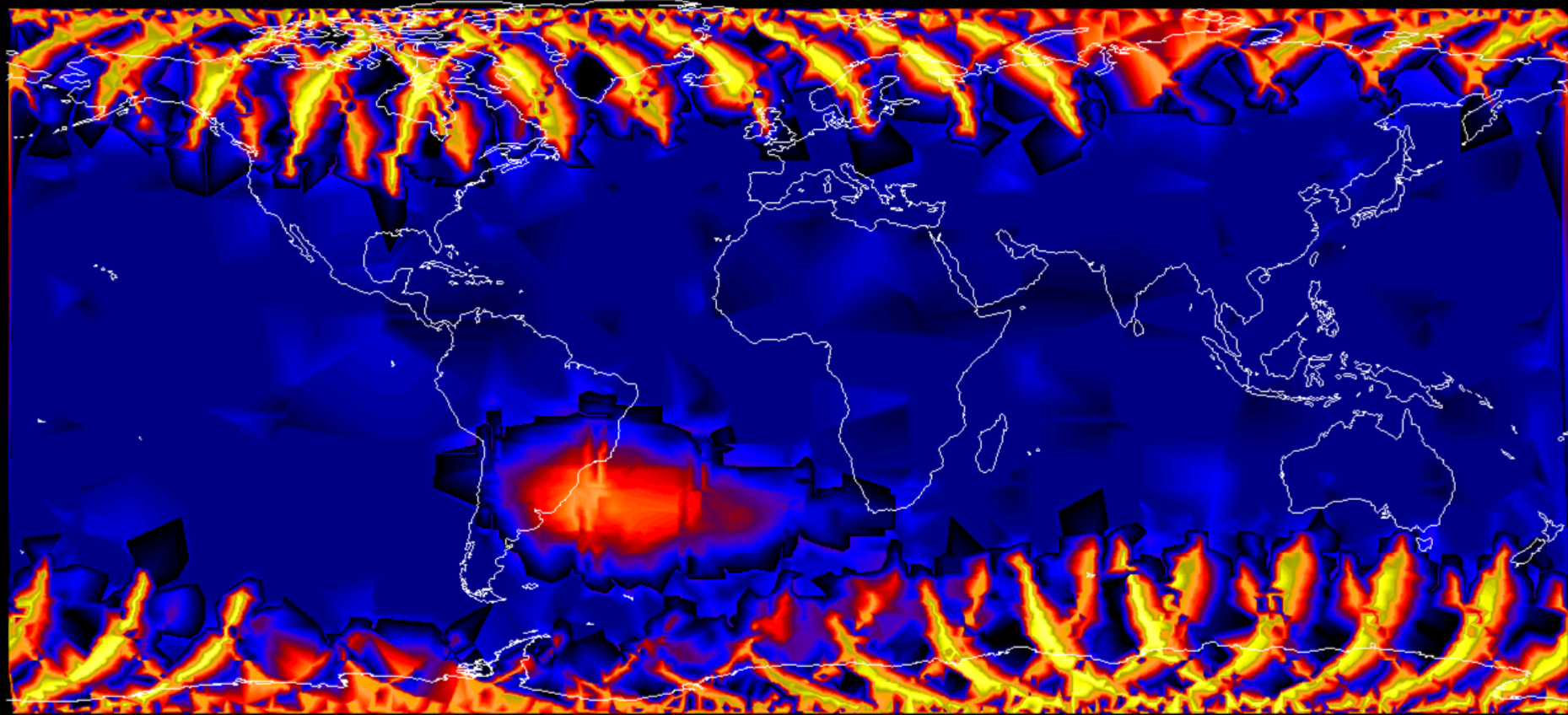
# Polar regions the most variable



# What is seen during a medium size storm?



# RESIK within CORONAS-F



# RESIK and MKL on CORONAS-F

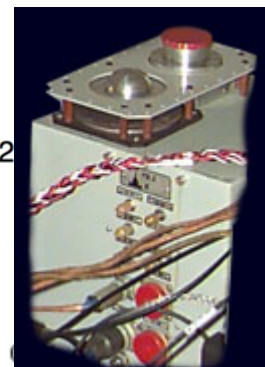
<http://smdc.sinp.msu.ru/index.py?nav=coronasf>

- Electron fluxes:
  - e.3-.6 - 0.3-0.6MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - e.6-1.5 - 0.6-1.5MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - e1.5-3 - 1.5-3.0MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - e3-6 - 3 - 6MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - e6-12 - 6 - 12MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
- Proton fluxes:
  - p1-5 - 1 - 5MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - p14-26 - 14 - 26MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - p23-26 - 23 - 26MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - p26-50 - 26 - 50MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - p50-90 - 50 - 90MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
- Mixed channels:
  - 0.5p>9 - electrons  $E > 0.5\text{MeV}$ , protons  $E > 9\text{MeV}$ ;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$  e>1p>9 -
  - electrons  $E > 1\text{MeV}$ , protons  $E > 9\text{MeV}$ ;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$  e1.6p23 -
  - electrons  $E > 1.6\text{MeV}$ , protons  $E > 23\text{MeV}$ ;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$  e1.6p90 -
  - electrons  $E > 1.6\text{MeV}$ , protons  $E > 90\text{MeV}$ ;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$
  - Alpha (He ions)a23-33 - alpha 23-33MeV;  $1/\text{cm}^2 \cdot \text{sr} \cdot \text{s}$



<http://www.springer.com/978-3-642-39267-2>

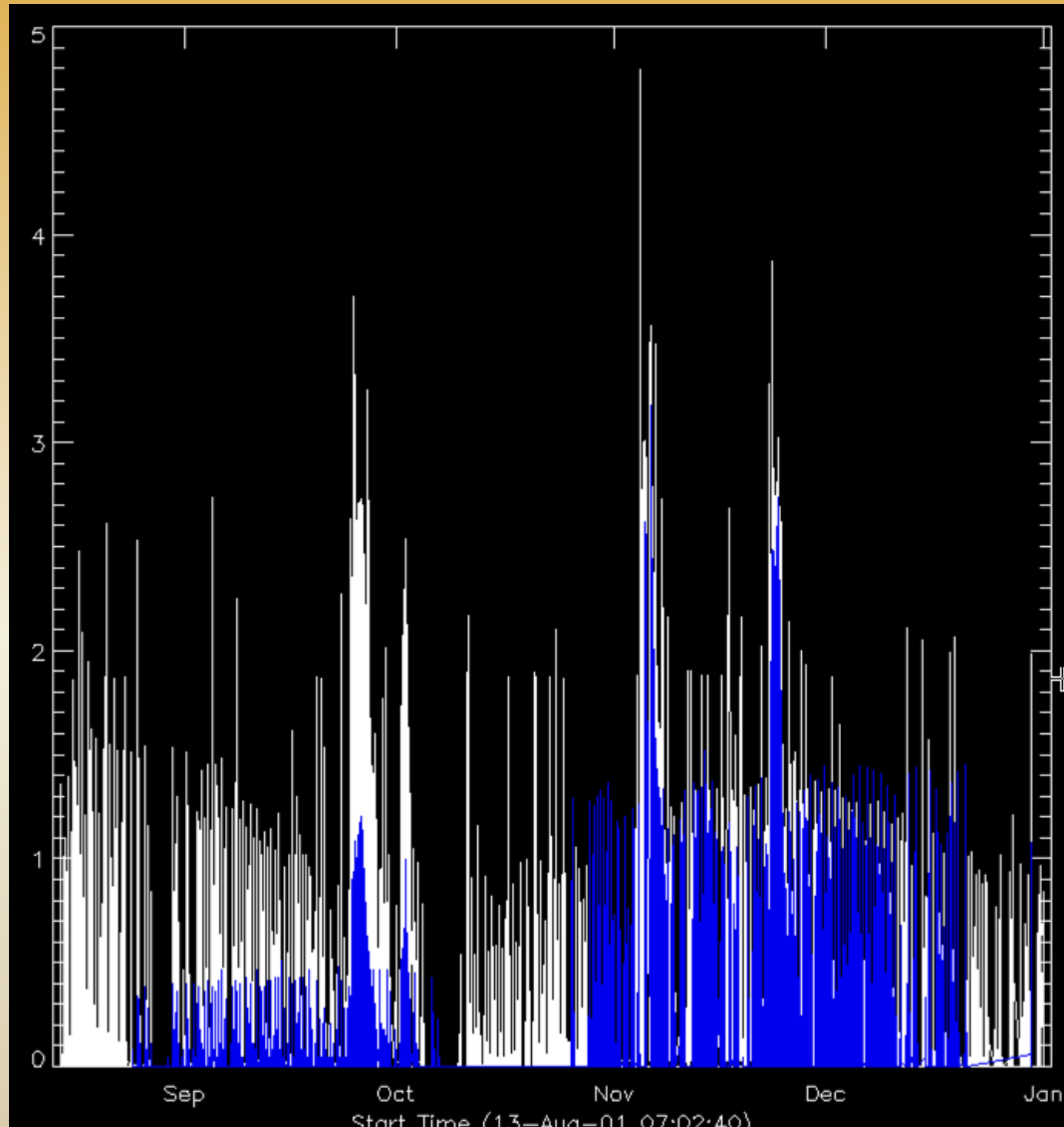
The Coronas-F Space Mission  
 Key Results for Solar Terrestrial Physics  
 Kuznetsov, V. (Ed.)  
 2014, XII, 483 p. 319 illus., 48 illus. in color  
 ISBN: 978-3-642-39267-2



<http://www.springer.com/astronomy/extraterrestrial+physics%2C+space+sciences/book/978-3-642-39267-2>

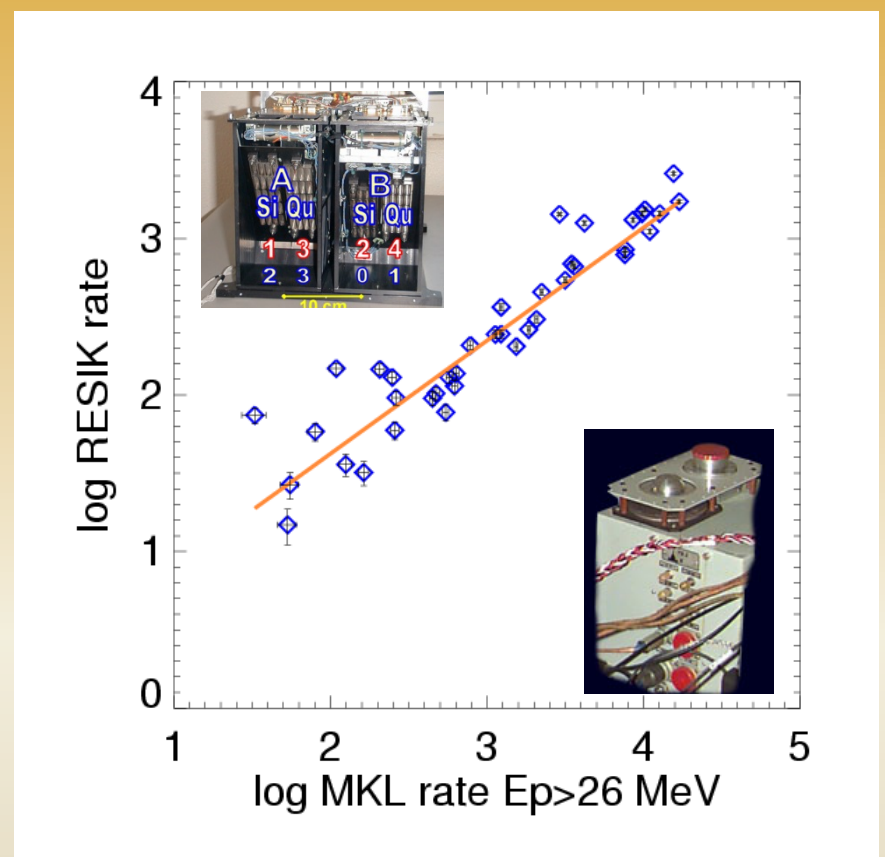
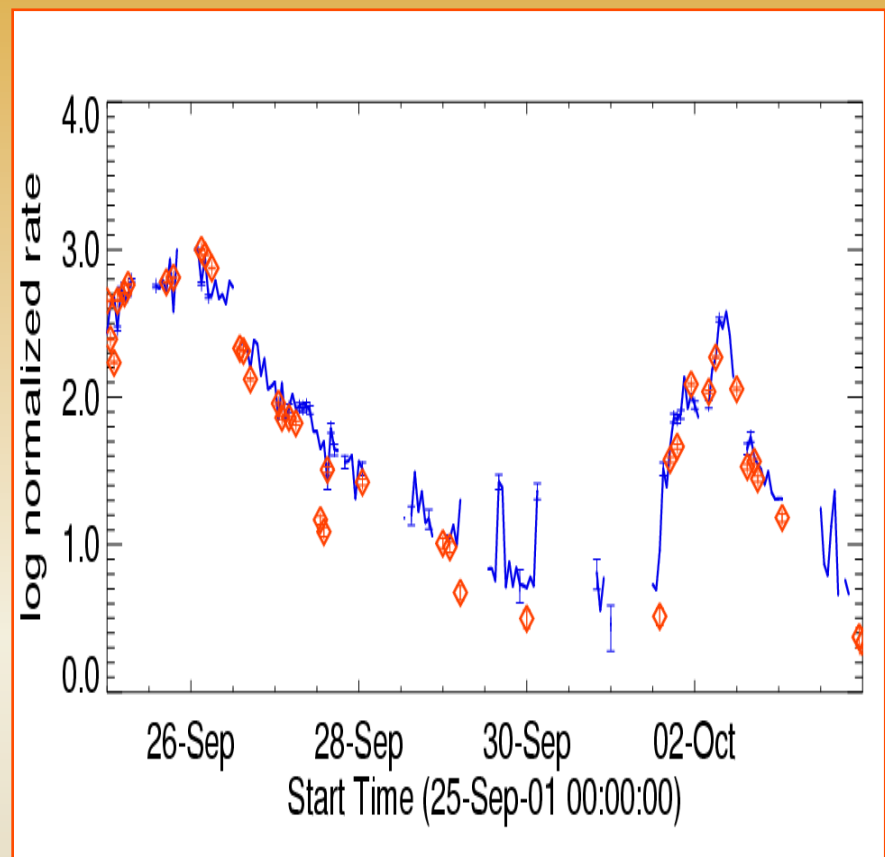


# MKL protons 26-50 keV & PIN



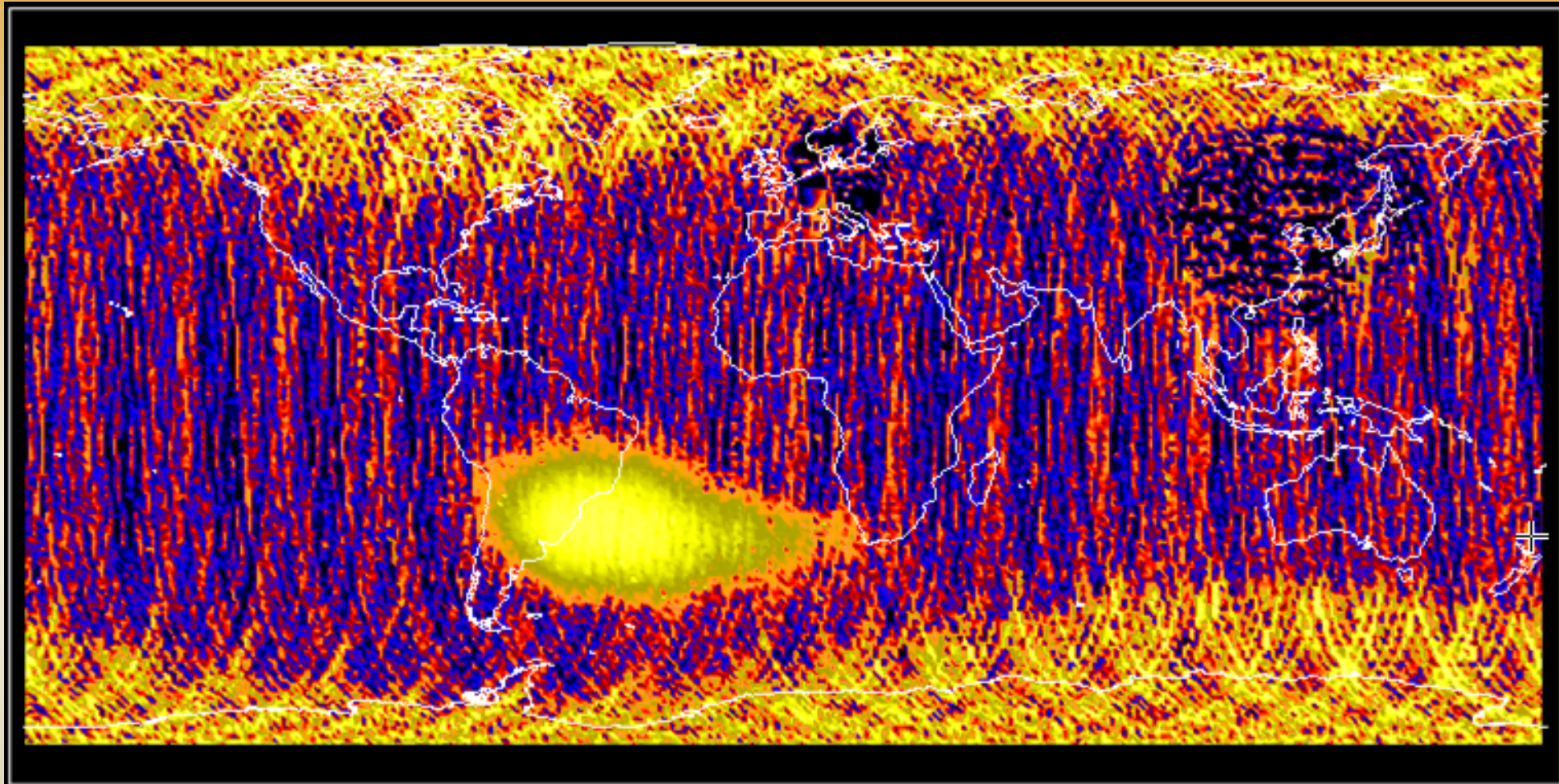


# corellation of PIN signal with MKL detector's



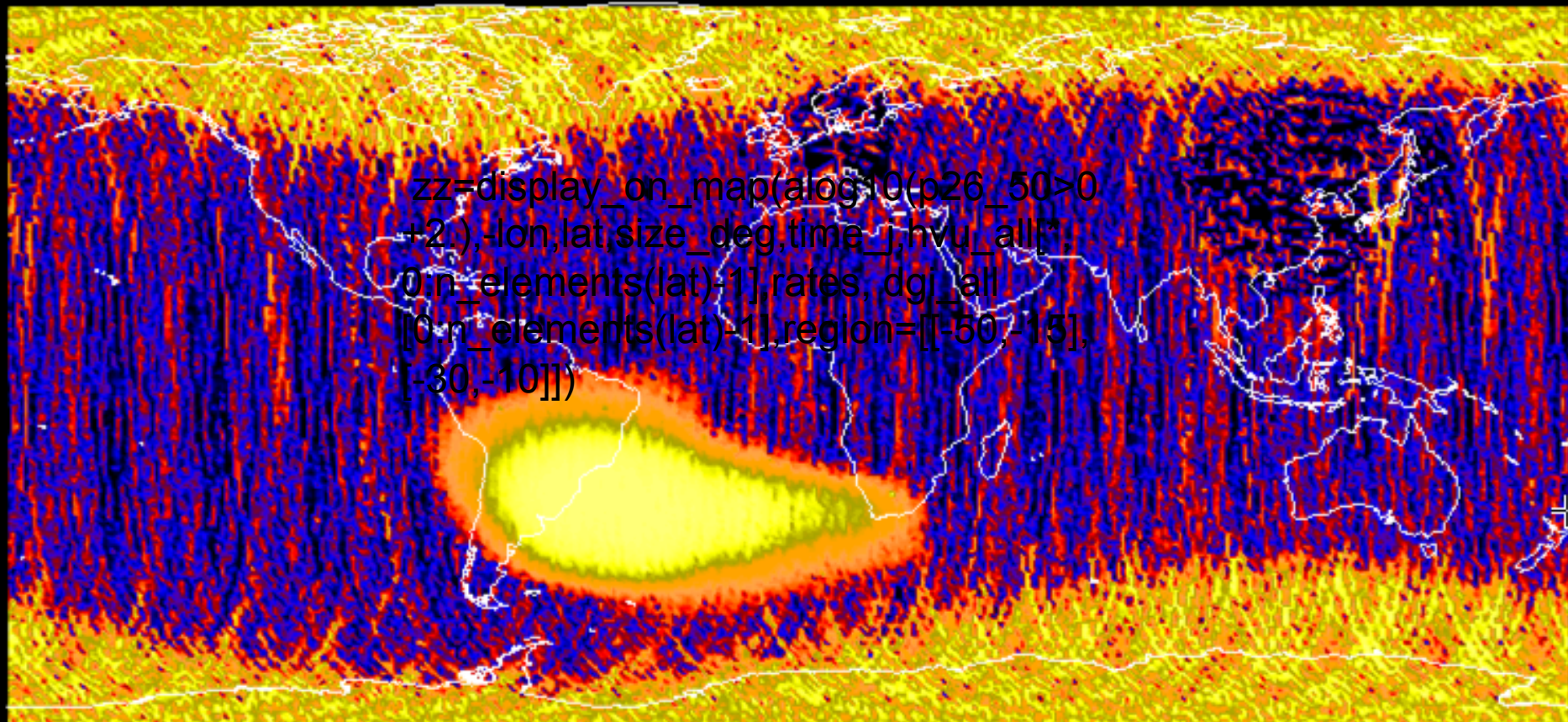
„early rather weak geomagnetic storms”

# MKL protons 23-26 keV



# MKL protons 26-50 keV

```
zz=display_on_map(alog10(p26_50>0  
+2.),-lon,lat,size_deg,time_j,hvu_all[*  
0:n_elements(lat)-1],rates,dgj_all  
[[0:n_elements(lat)-1],region=[[ -50, 15],  
[-30,-10]])
```



# Now back to RESIK gas detectors (Ar + Xe)

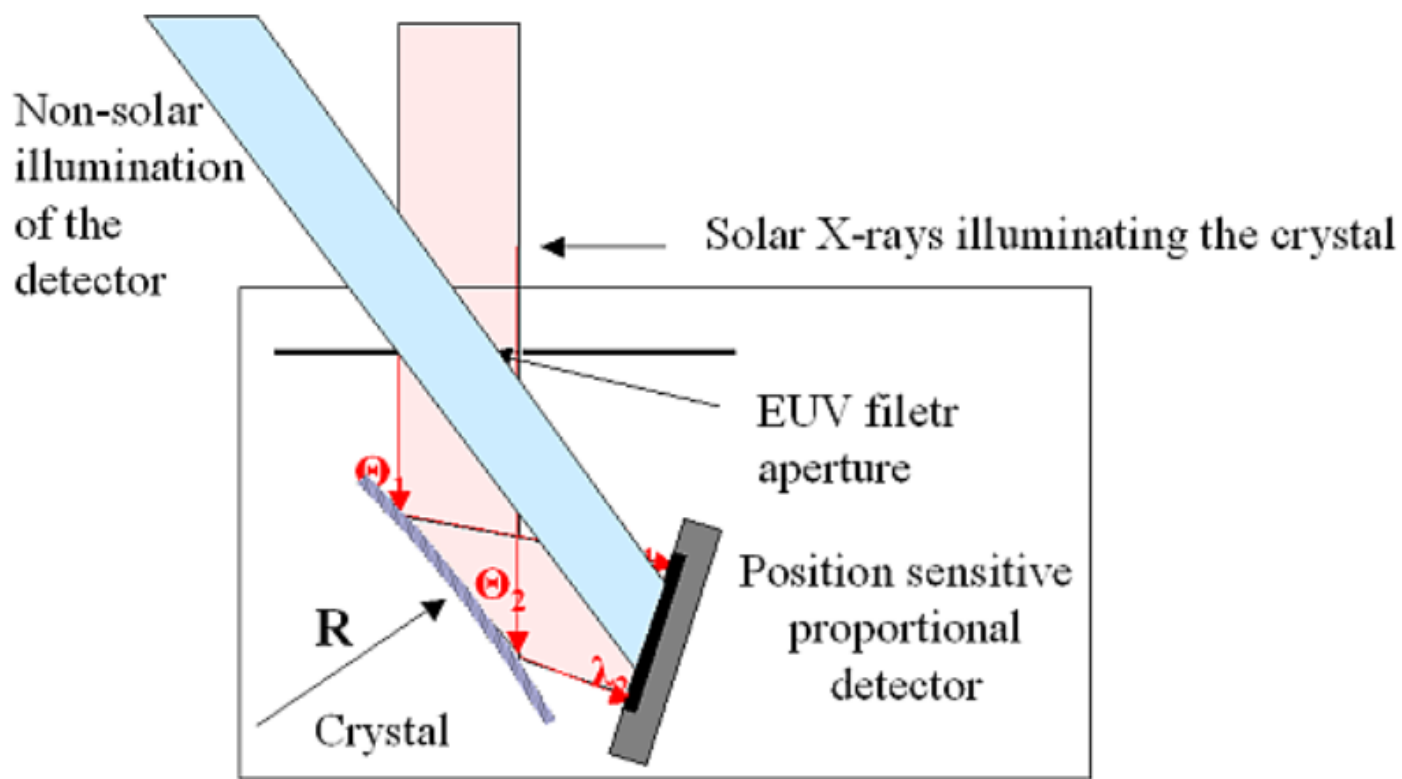
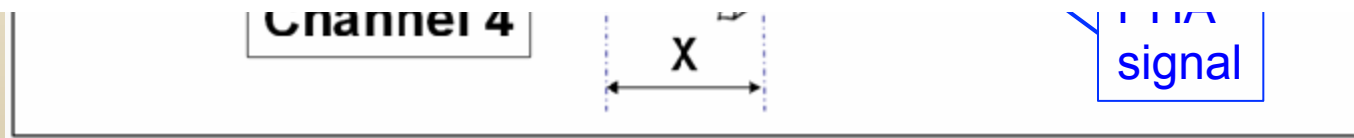
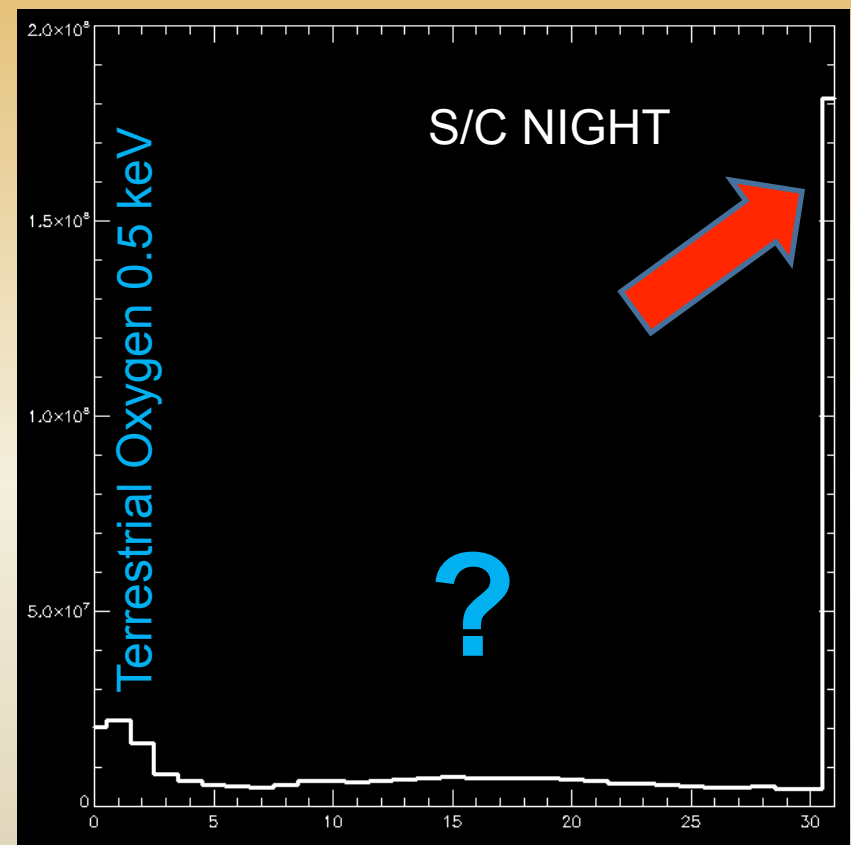
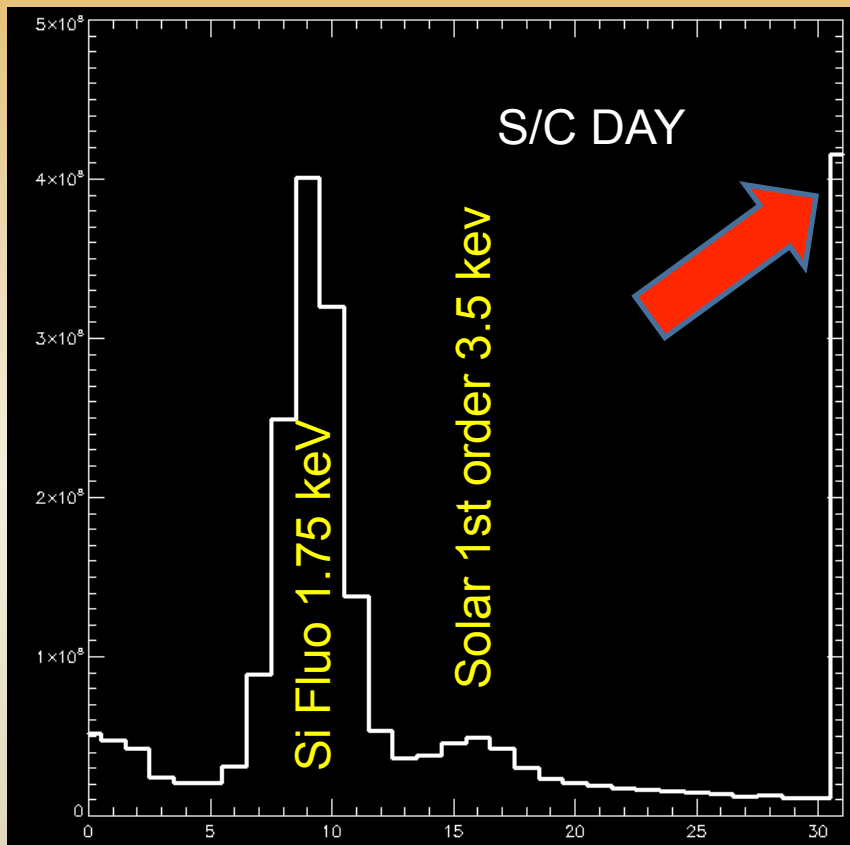


Figure 1. A scheme of X-ray beams within one section of RESIK Bragg spectrometer



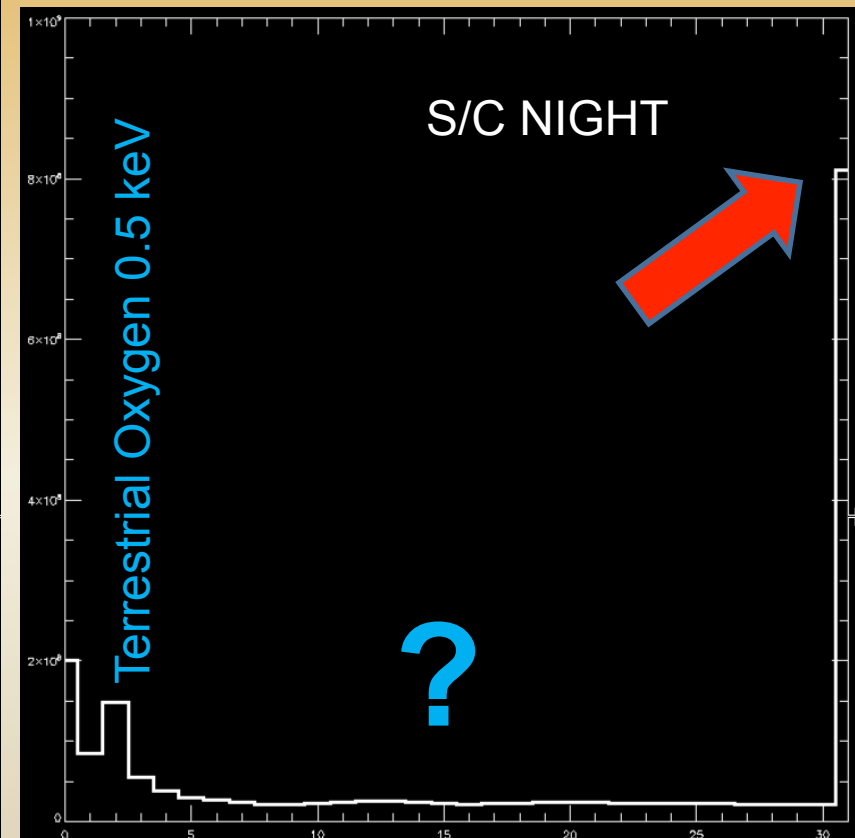
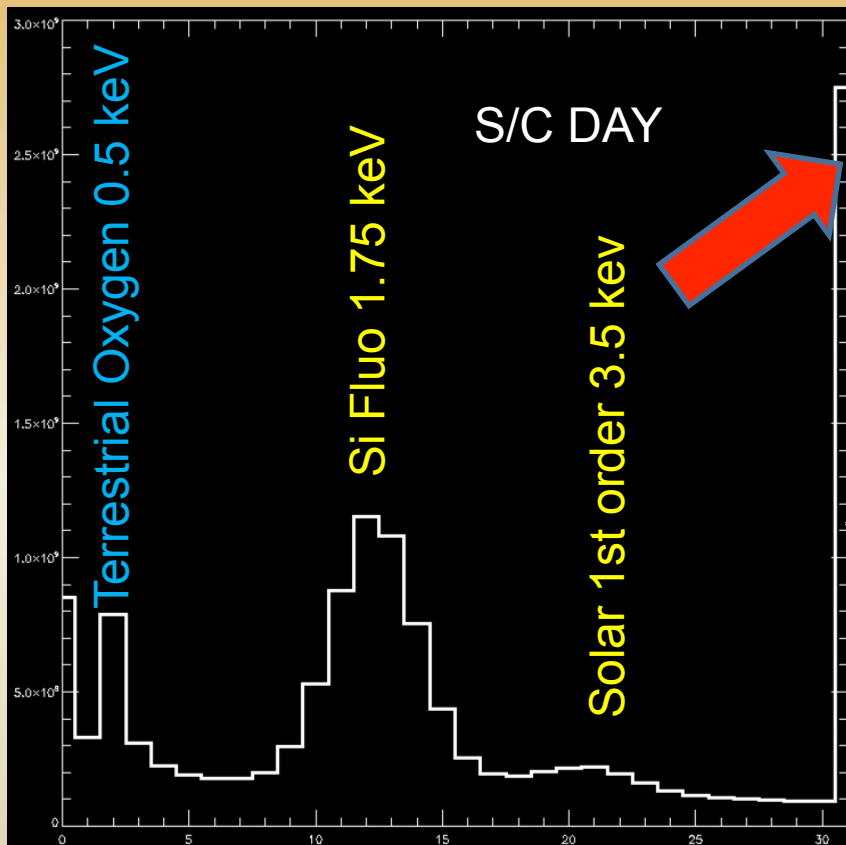
# RESIK PHA Spectra

## Example PHA channel #2 HV:1450 V

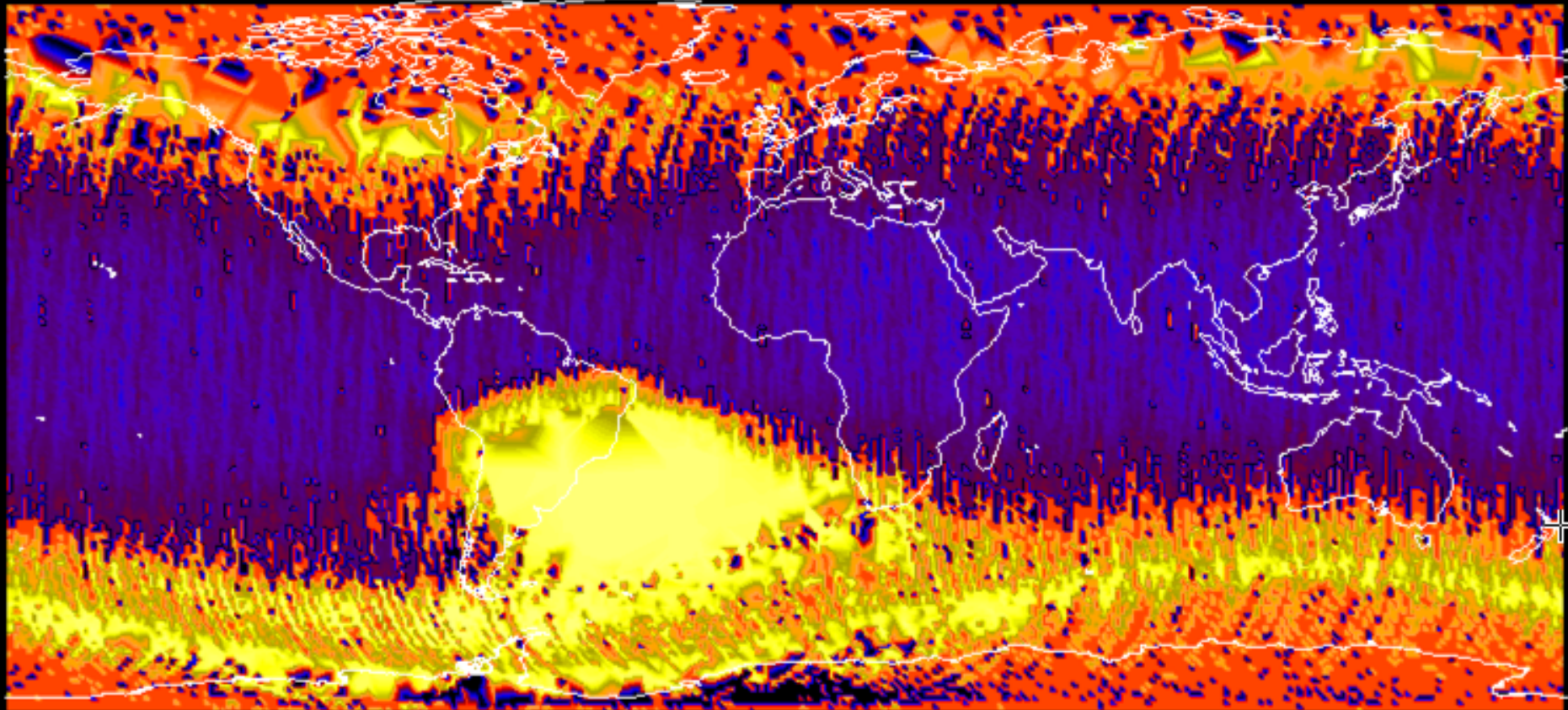


# RESIK PHA Spectra

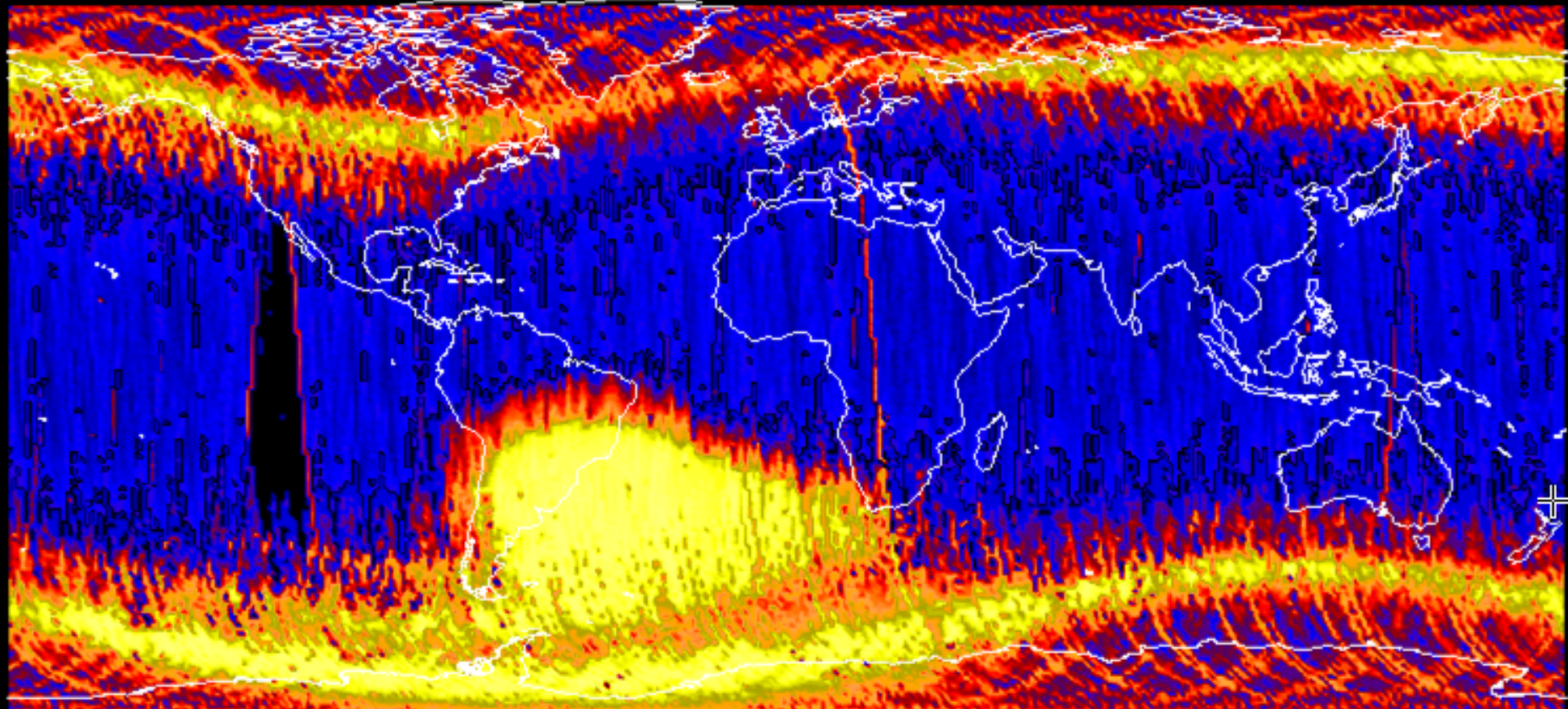
## Example PHA channel #2 HV:1480 V



# Terrestrial 0.5 keV emission S/C nights

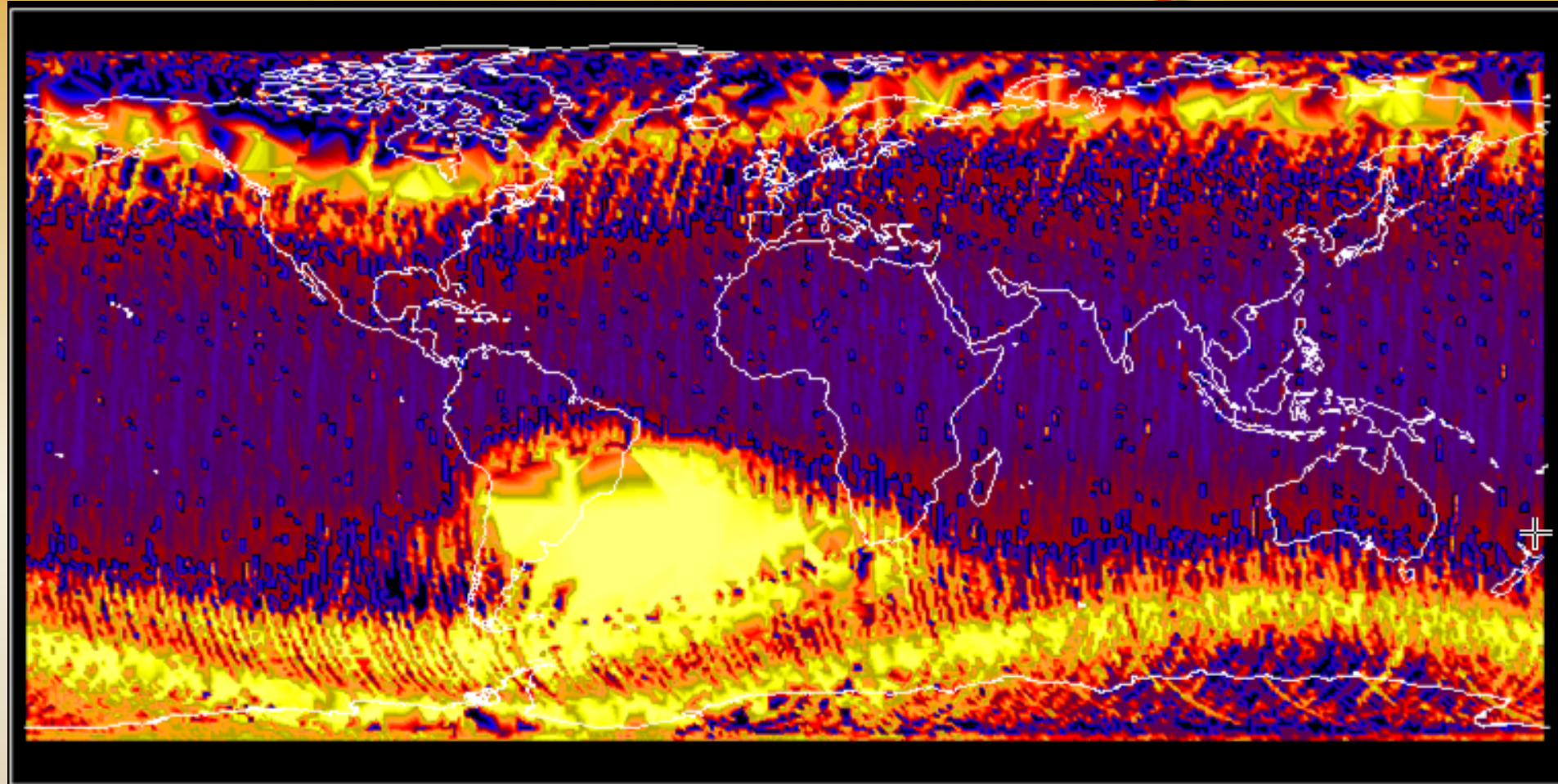


# Terrestrial 0.5 keV emission S/C days





# Terrestrial 1.5 -7 keV emission S/C nights

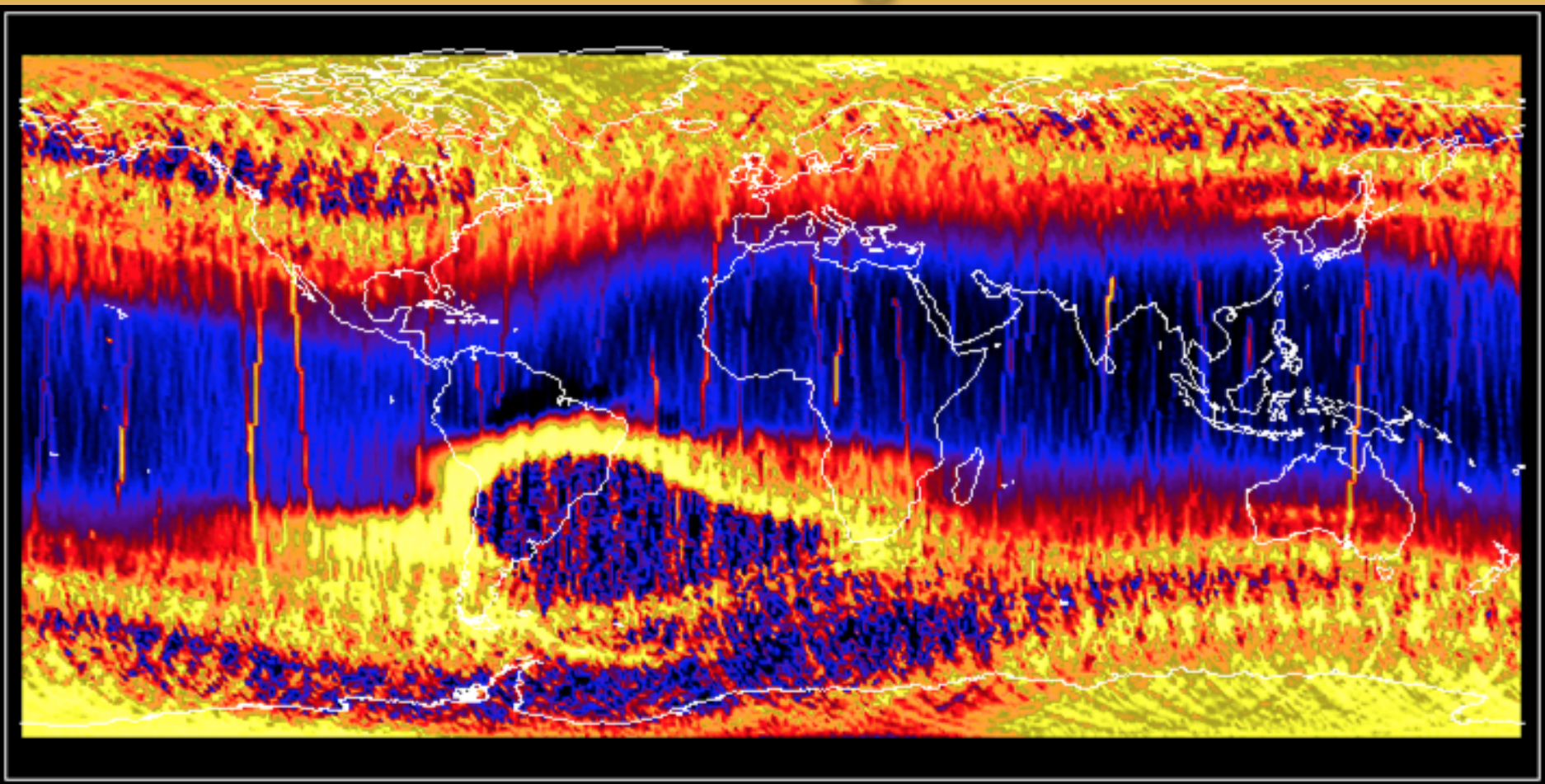


CBK PAN

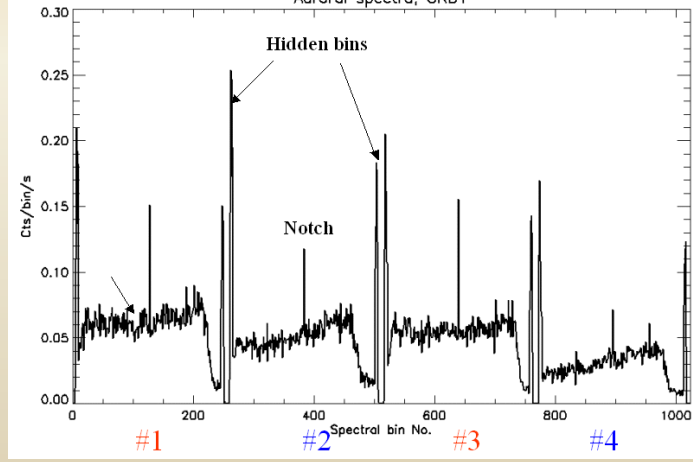
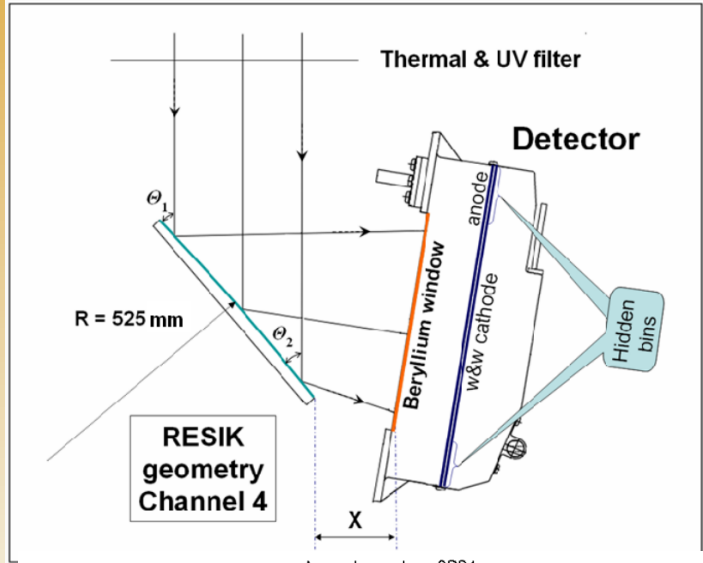
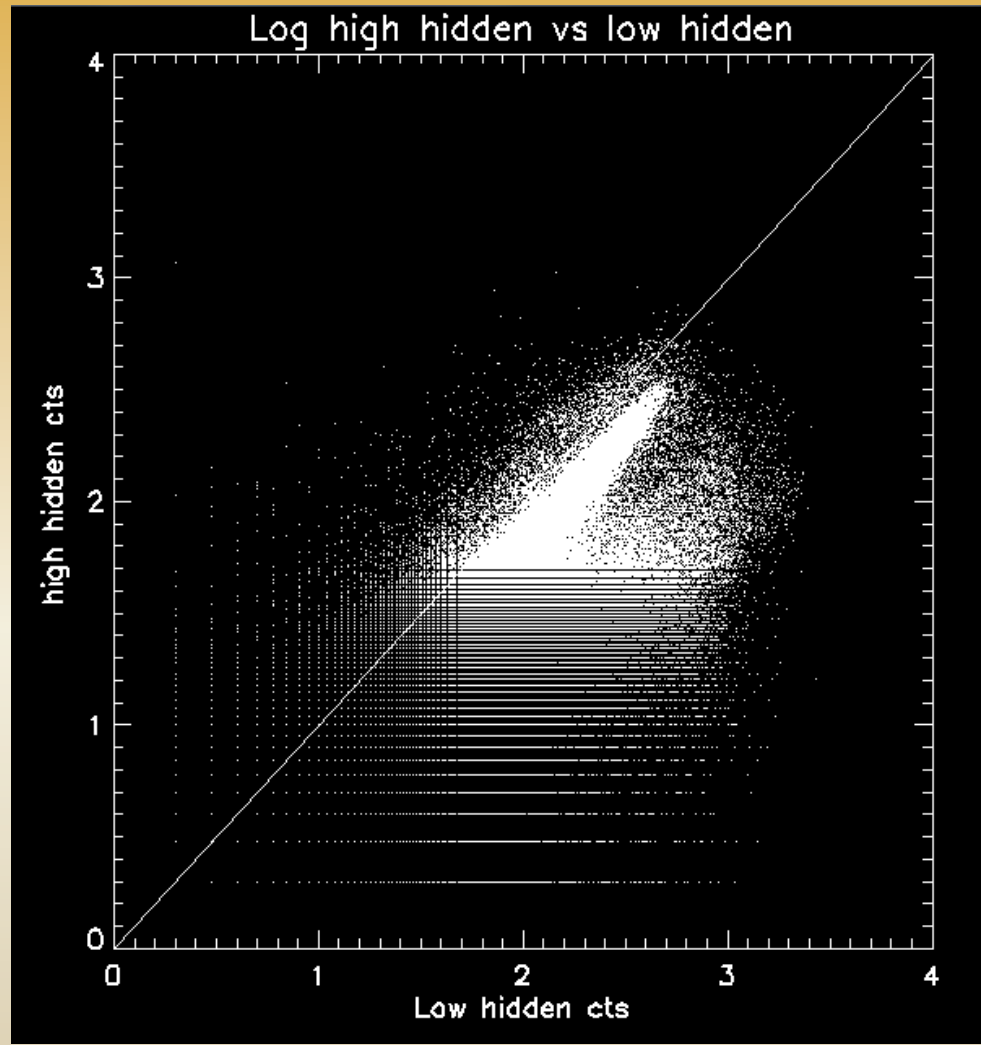
SRC PAS



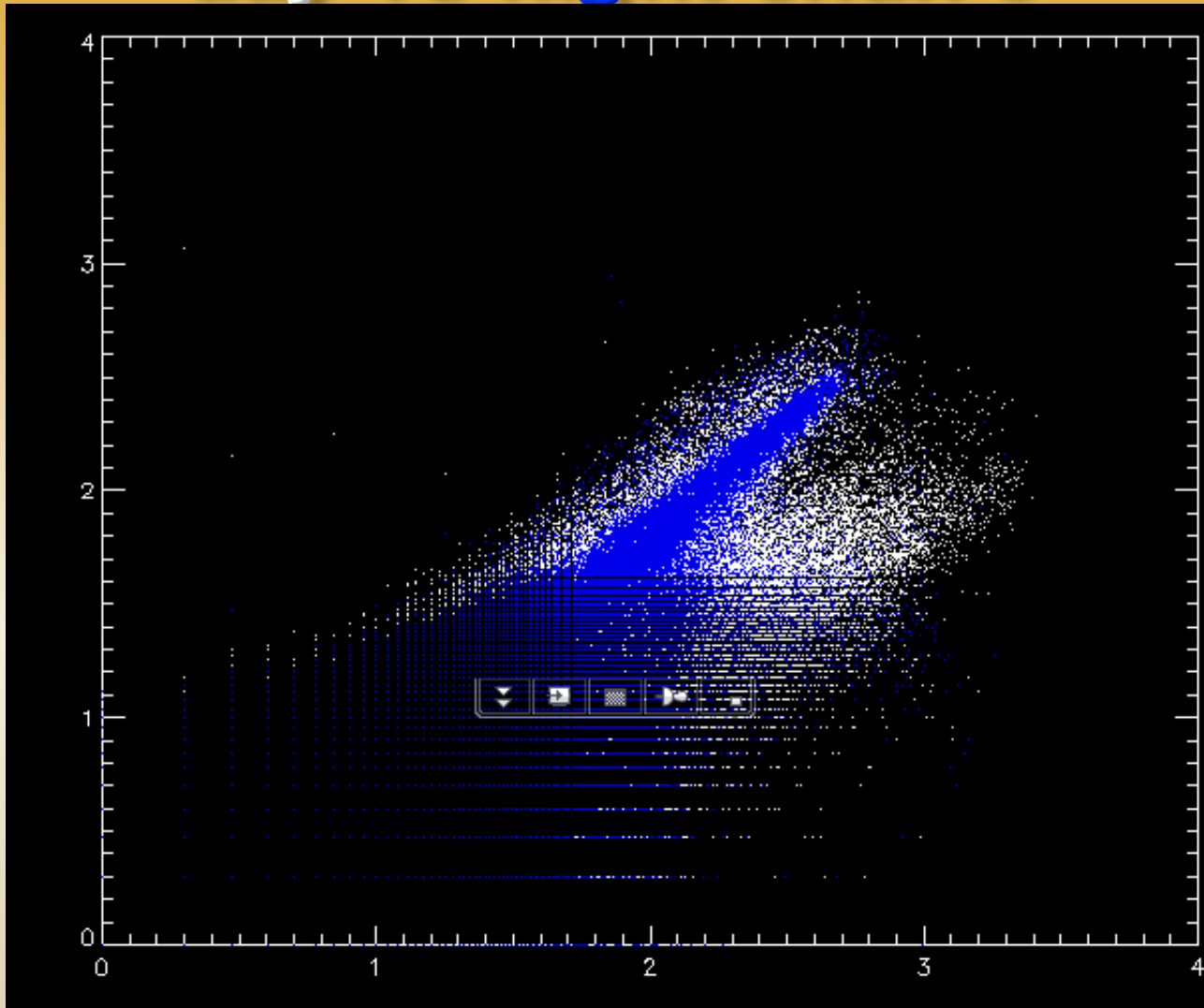
31<sup>st</sup> bin 1 x 1 deg resolution



# Hidden bins correlations

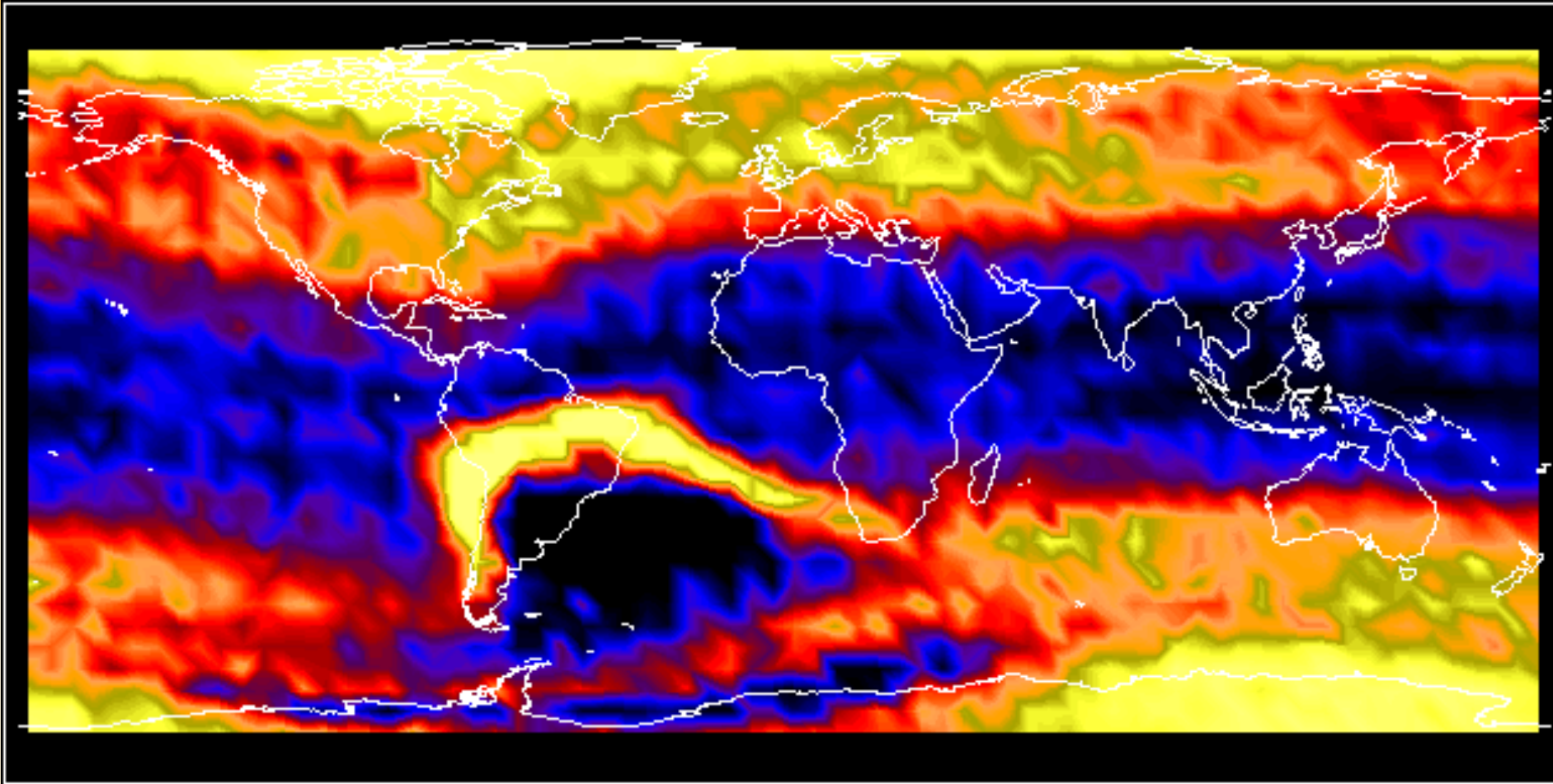


# Hidenn bins day vs night chan 0

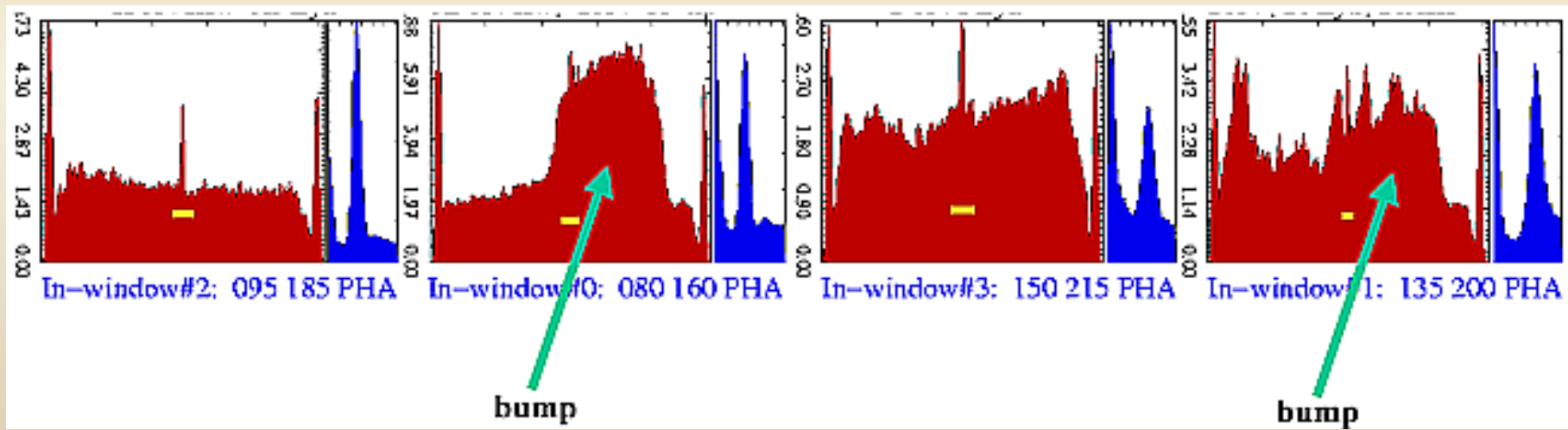
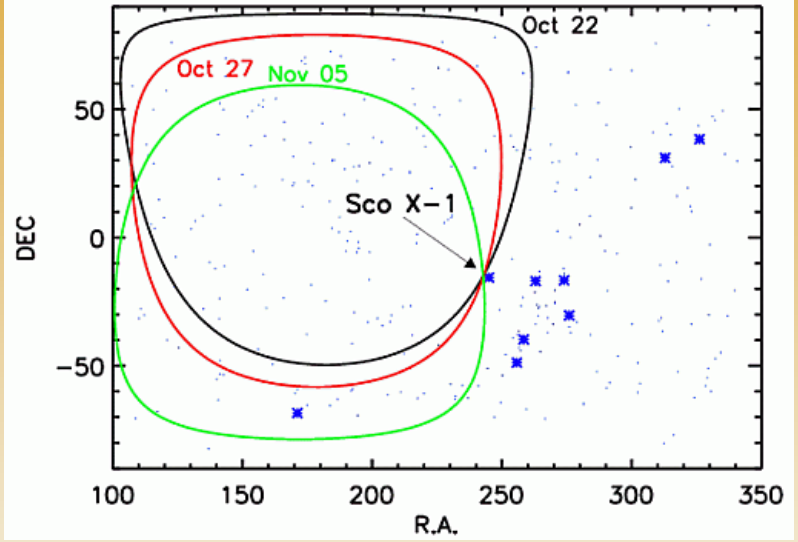
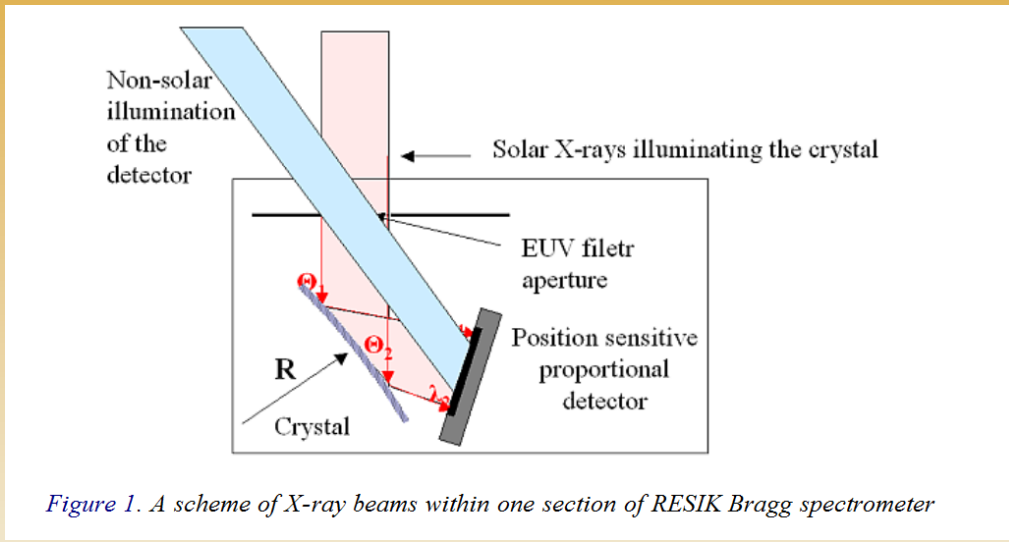




# High hidden chan 0



# Non-solar contributions





# Ongoing & Possible collaborations



- ▣ Skobeltsyn Institute MGU, Moscow
  - Understanding & Calibrations of various RESIK Signals- for quiet and stormy conditions
- ▣ Plasma physics group, SRC Warsaw
  - Understanding signals due to severe storms
- ▣ Lebedev Physical Institute, Moscow
  - Relation between the PIN signals and CCD tracks signal
- ▣ ROB & SPENVIS groups, Belgium
  - Quiet & stormy condition response of PROBA-2



# Thank you eHeroes Team



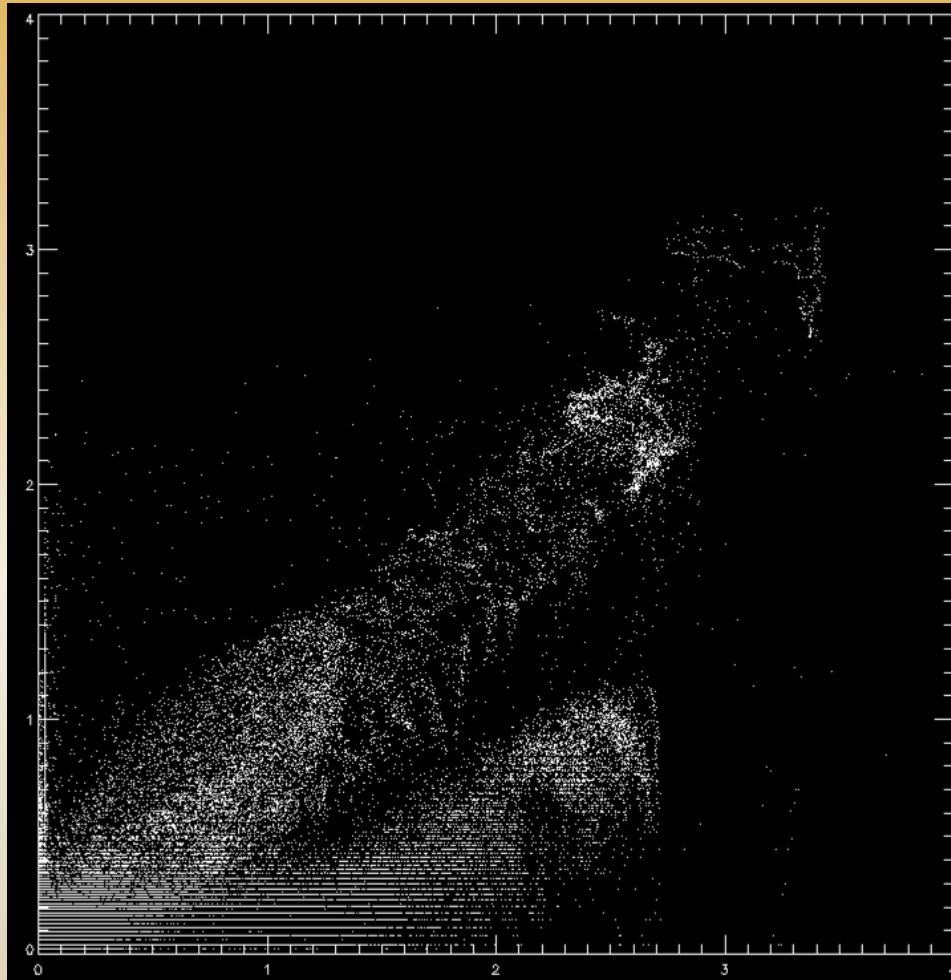




# RESIK and MKL on CORONAS-F MKL data on:



<http://smdc.sinp.msu.ru/index.py?nav=coronasf>



# RESIK within CORONAS-F

# RESIK within CORONAS-F

# RESIK within CORONAS-F