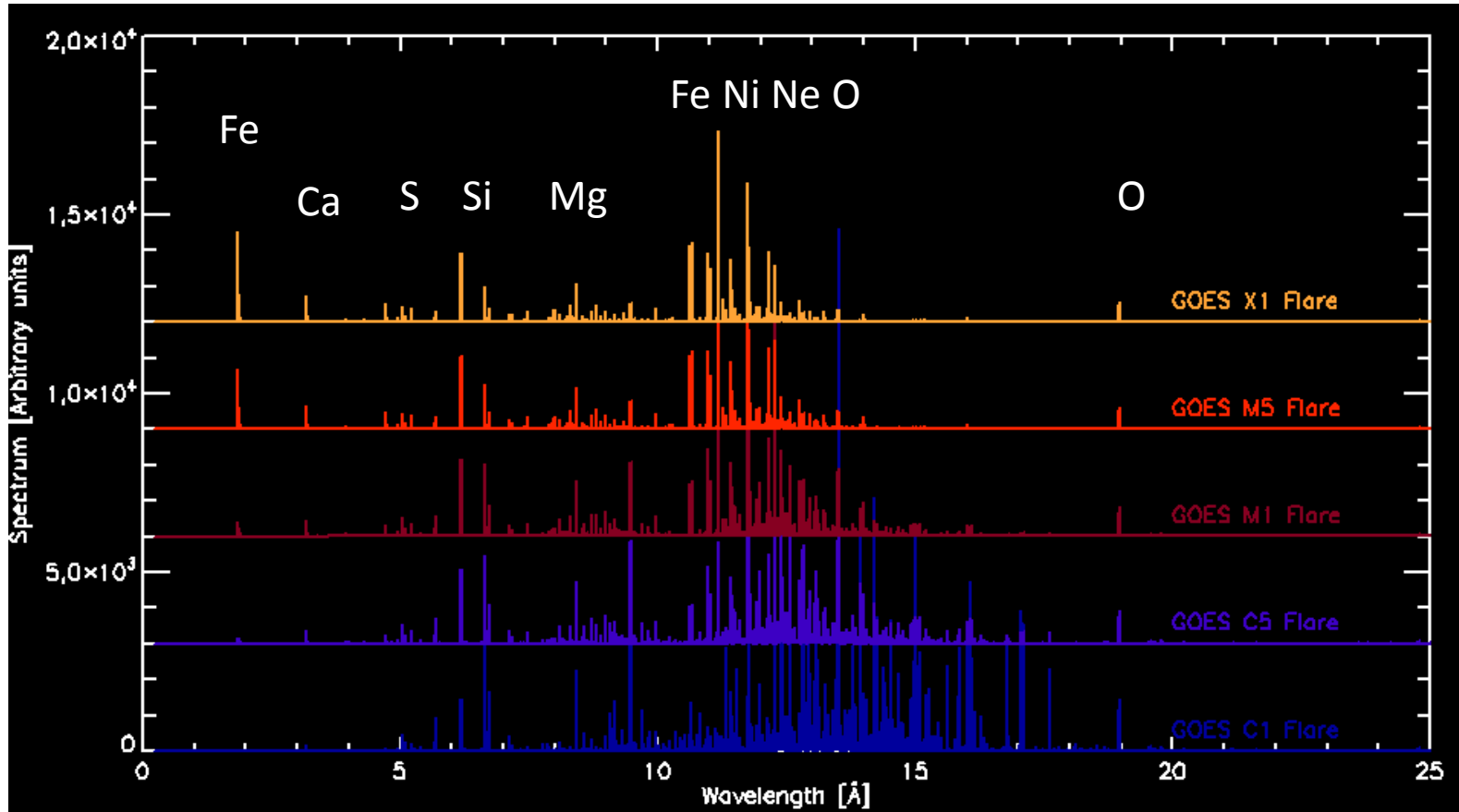


A blue-tinted image of the Sun, showing the solar corona and various solar activity features like flares and coronal loops. The Sun is the central focus of the upper half of the slide.

# Bragg X-ray spectroscopy of the solar corona

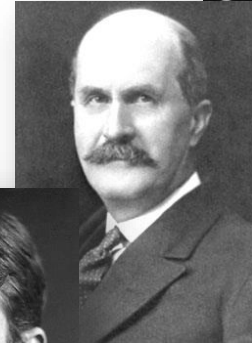
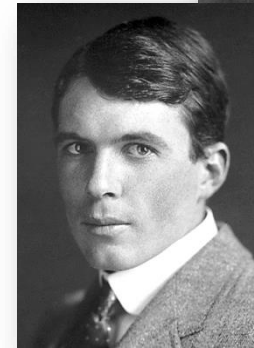
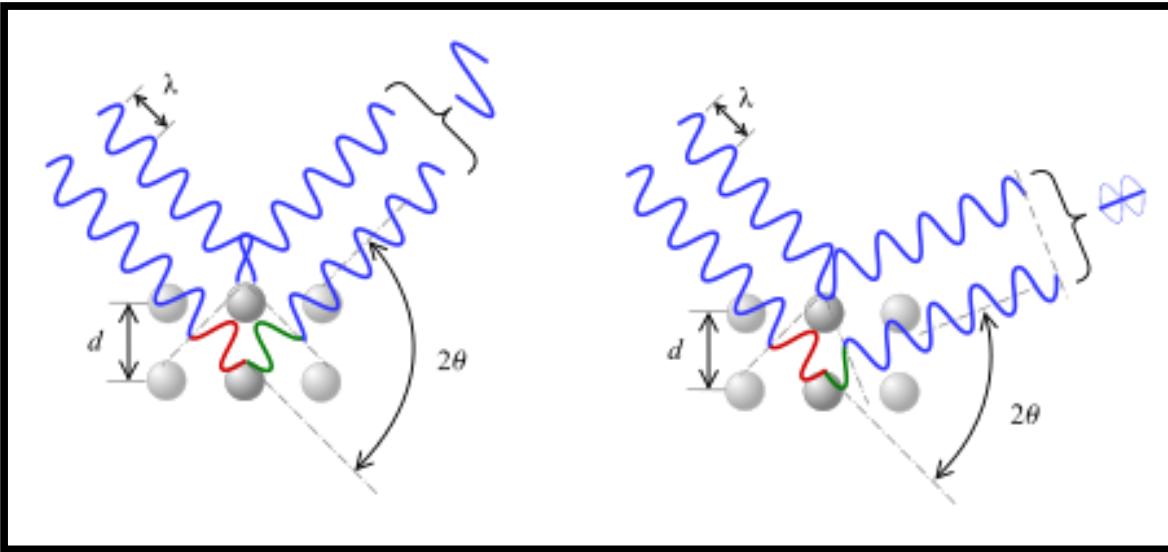
**Żaneta Szaforz**, Stefan Płoceniak, Janusz Sylwetrer, Jarosław Bąkała, Mirosław Kowaliński, Daniel Ścisłowski, Piotr Podgórski, Marek Siarkowski, Barbara Sylwester, Marek Stęślicki, Zbigniew Kordylewski

# Solar X-ray spectrum



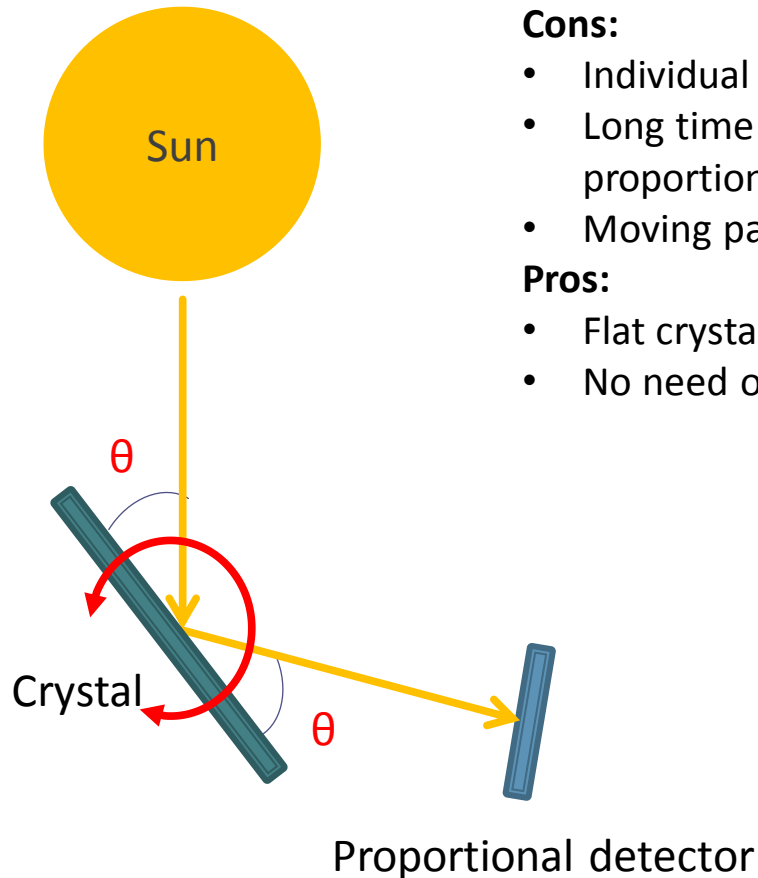
# Bragg's law

$$2 d \sin \theta_B = n \lambda$$



Max von Laue  
William Henry Bragg  
William Lawrence Bragg

# Flat crystal spectrometers



## Cons:

- Individual wavelengths are not recorded at the same time!
- Long time needed to observe the whole spectrum (time proportional to the length of observed spectral range)
- Moving parts

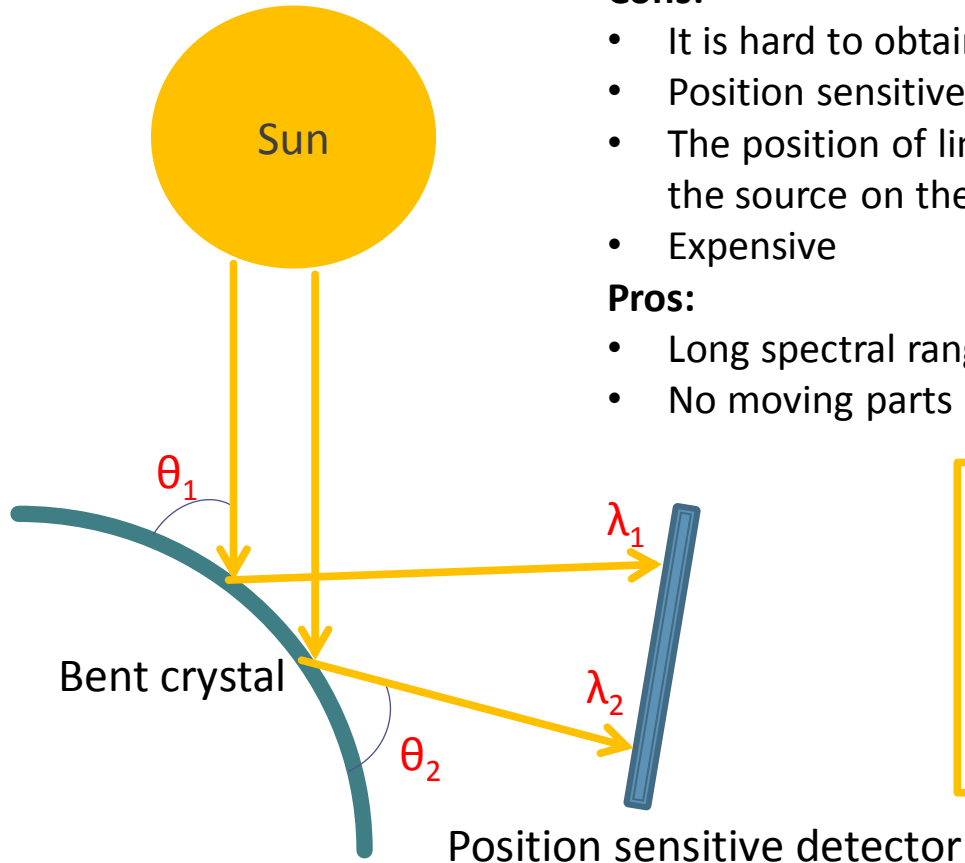
## Pros:

- Flat crystals are easier to obtain (than the alternative bent crystals)
- No need of position sensitive detector

## Previous instruments:

- SOLFLEX on board the P78-1 (1979)
- Flat Crystal Spectrometer on-board the Solar Maximum Mission (1980)
- Solar soft X-ray bright line spectrum analyzer on-board the HINOTORI (1981)
- DIOGENESS on-board the CORONAS-F (2001)

# Bent crystal spectrometers



## Cons:

- It is hard to obtain the perfect cylindrical shape
- Position sensitive detector is needed
- The position of lines on the detector can depend on the position of the source on the solar disk. Collimator or tracing system is needed
- Expensive

## Pros:

- Long spectral ranges can be observed simultaneously
- No moving parts

## Previous instruments:

- Bent Crystal Spectrometer on-board the Solar Maximum Mission (1980)
- Bragg Crystal Spectrometer on-board the Yohkoh (1991)
- Resik on-board the CORONAS-F (2001)

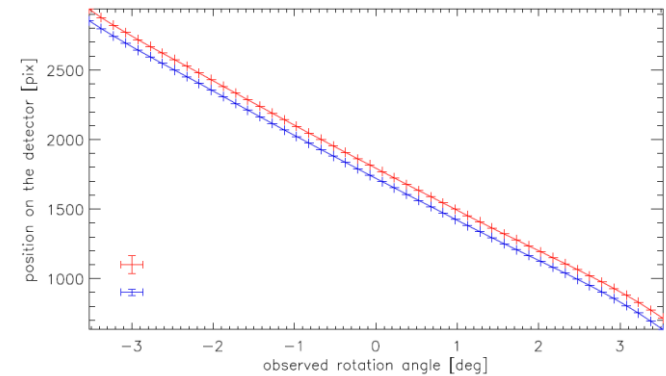
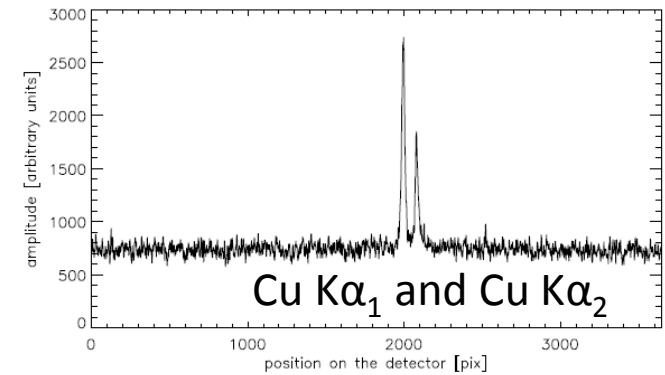
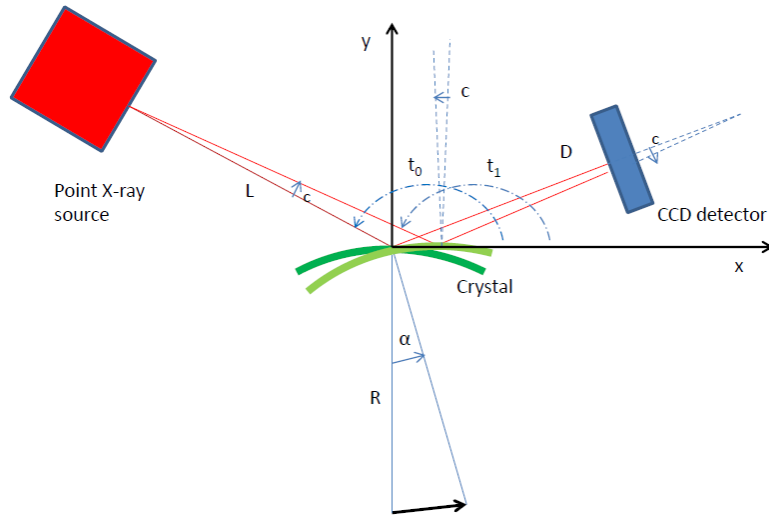
# Designing a spectrometer

- The Bragg diffraction is very inefficient!
- Because the X-ray spectrometer had to be placed outside the atmosphere, in most cases strong restrictions to the weight or size of instrument are placed
- Fluorescence
- The signal measured by the X-ray spectrometer is the convolution of the solar radiation spectrum and the response function of the instrument. Precise knowledge of the properties of all elements in the path of the incident beam is needed to determine the instrument response function
- Calibration is needed

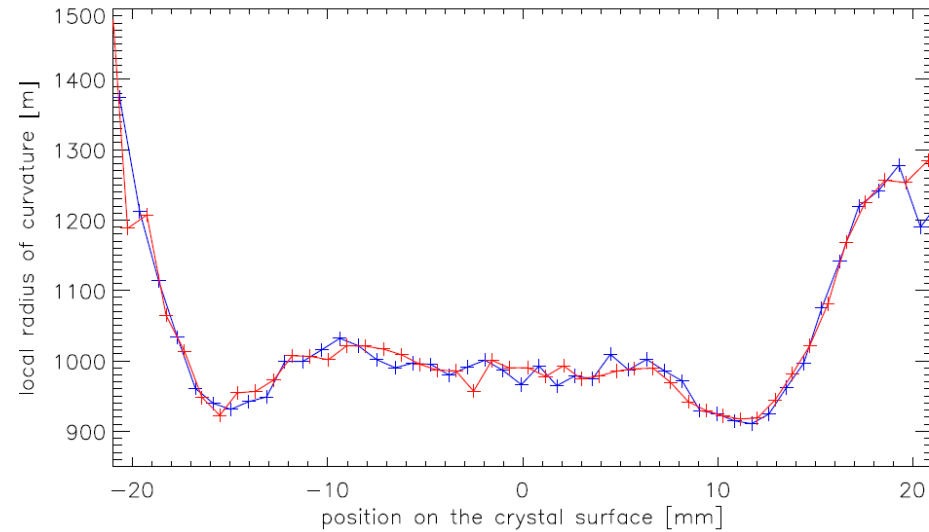
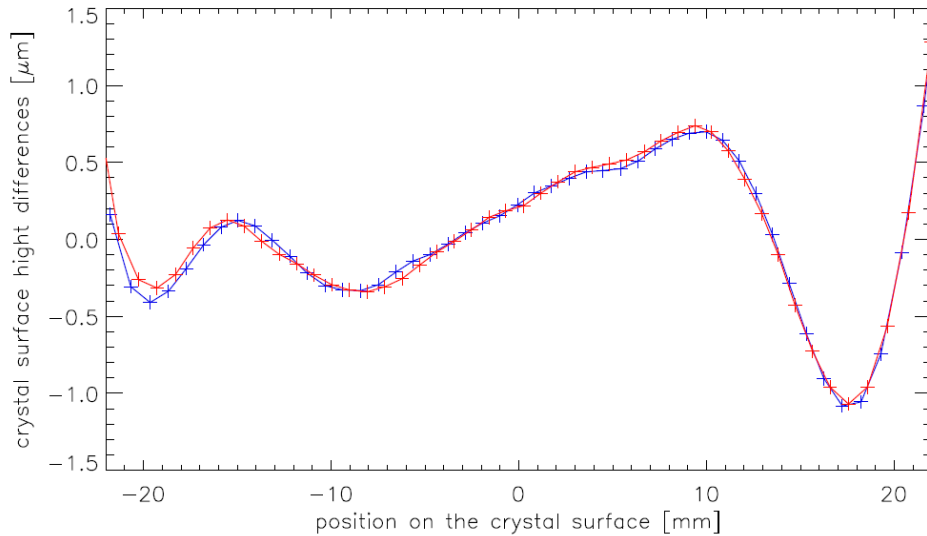
# Crystal's properties measurements

Laboratory characterization of bent monocrystal wafers for Bragg X-ray spectroscopy

Stefan Płoceniak · Żaneta Szaforz



# Crystal's properties measurements



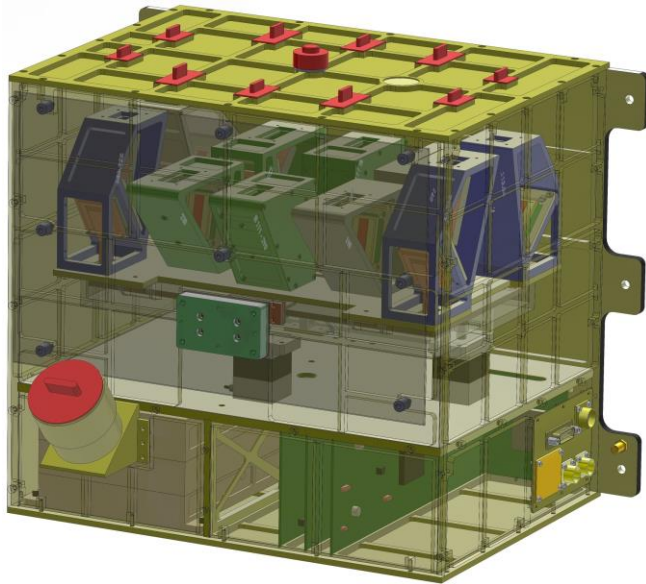
Silicon (400) crystal wafer bent convex to radius of  $R = 984$  mm  
The crystal length was 463 mm



# Future instruments

## ChemiX

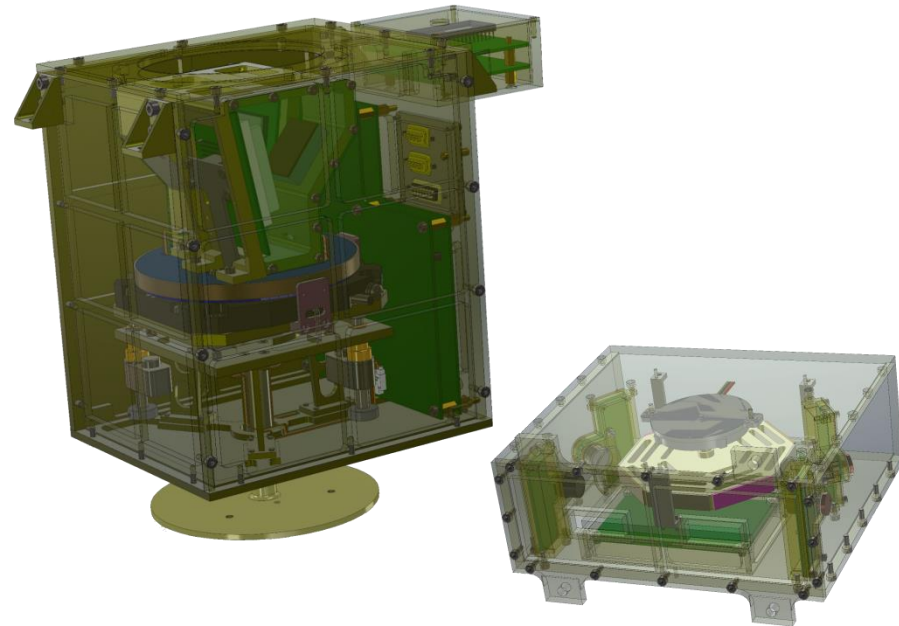
for Interhelioprobe Mission



M. Siarkowski, J. Sylwester, J. Bakała, Ż. Szaforz, M. Kowaliński, Z. Kordylewski, S. Płoceniak, P. Podgórski, B. Sylwester, W. Trzebiński, M. Stęślicki, K. J. H. Phillips, O. V. Dudnik, E. Kurbatov, V. D. Kuznetsov, S. Kuzin, I. V. Zimovets; **Experimental Astronomy (2016)**, Volume 41, Issue 3, pp 327–350

## SolpeX

for KORTES modul on ISS



J. Sylwester, S. Płoceniak, J. Bakała, Ż. Szaforz, M. Stęślicki, D. Ścisłowski, M. Kowaliński, P. Podgórski, J. Hernandez and S. Shestov; **Proceedings of the International Astronomical Union (2015)**, Volume 305, pp. 114-120

**Thank you for your attention**

