

# A common analysis of RESIK, SPIRIT and RHESSI observations for LDE flare of 15 April 2002

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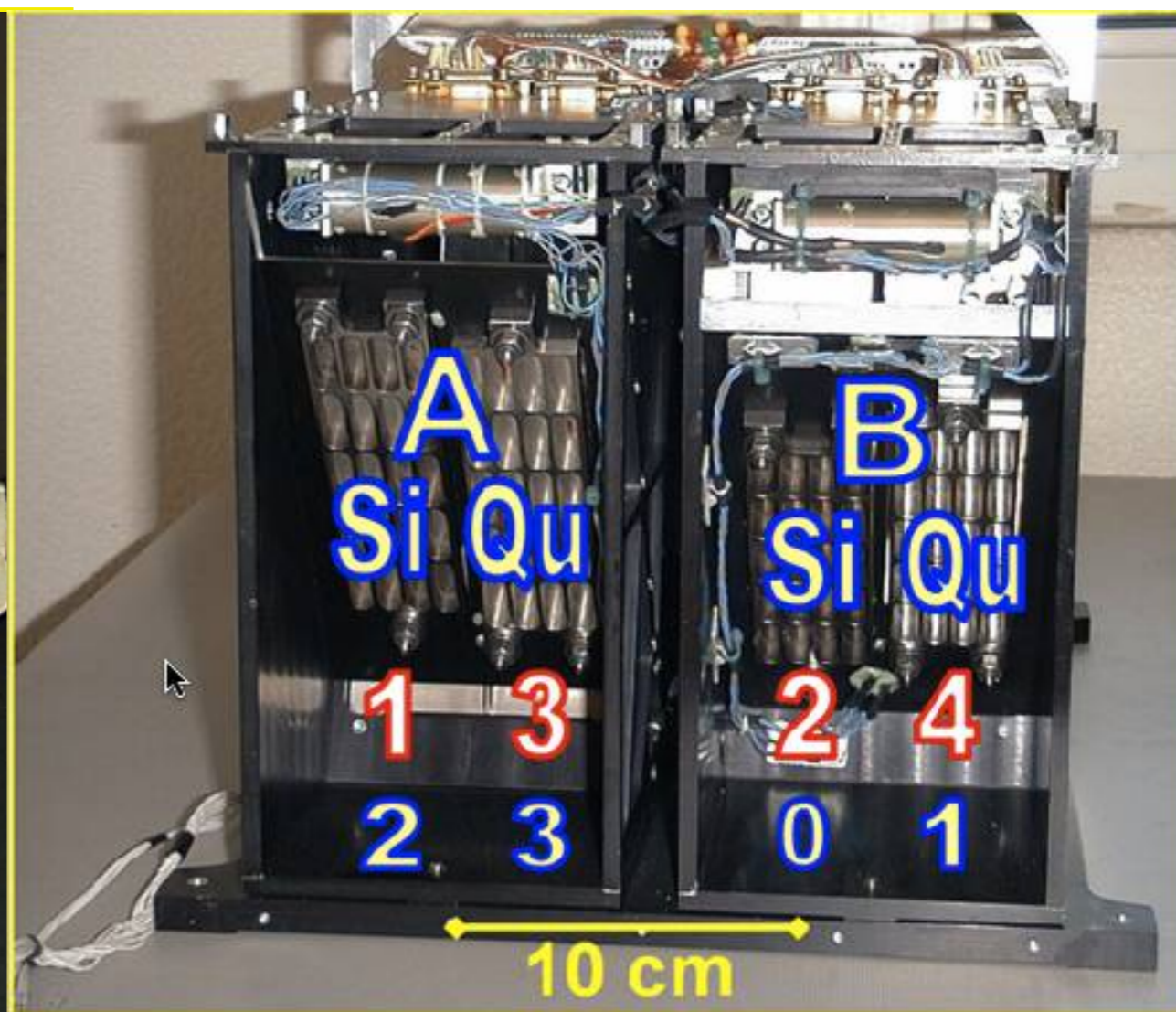
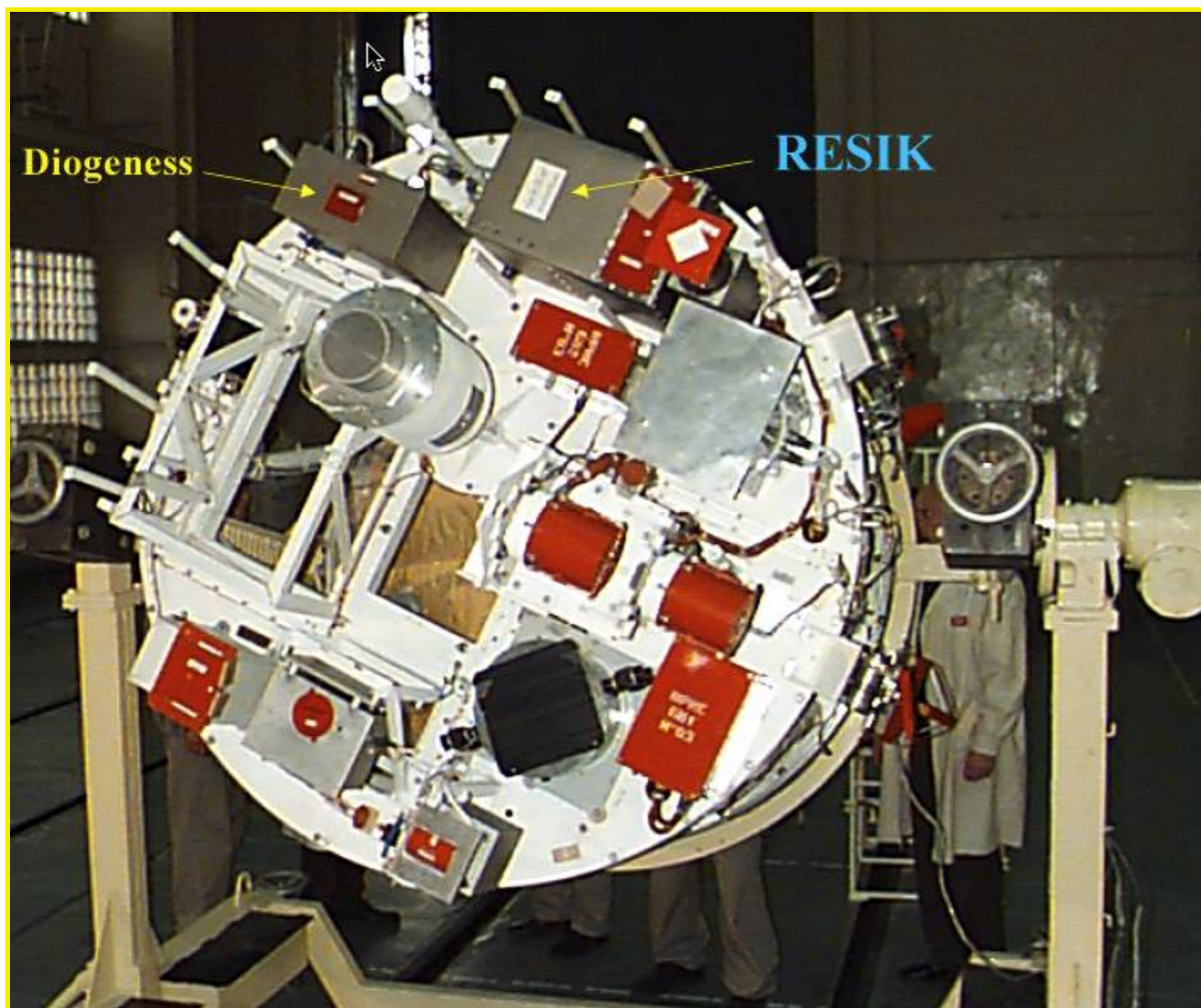
\*\* Lebedev Institute of the Russian Academy of Sciences



# Outline

- Coronas-F
- 15 April 2002 flare
- RESIK, SPIRIT, EIT/SoHO, RHESSI data
- Isothermal analysis & DEM

# Coronas-F



Characteristics of RESIK channels

Wavelength				
nominal range <sup>d</sup> [Å]	3.40–3.80	3.83–4.27	4.35–4.86	5.00–6.05
extreme range <sup>e</sup> [Å]	3.33–3.90	3.78–4.32	4.23–4.92	4.90–6.15
Wavelength				
resolution [mÅ]	8	9	12	17
Dispersion <sup>c</sup> [mÅ/bin]	2.49	2.28	2.85	4.99

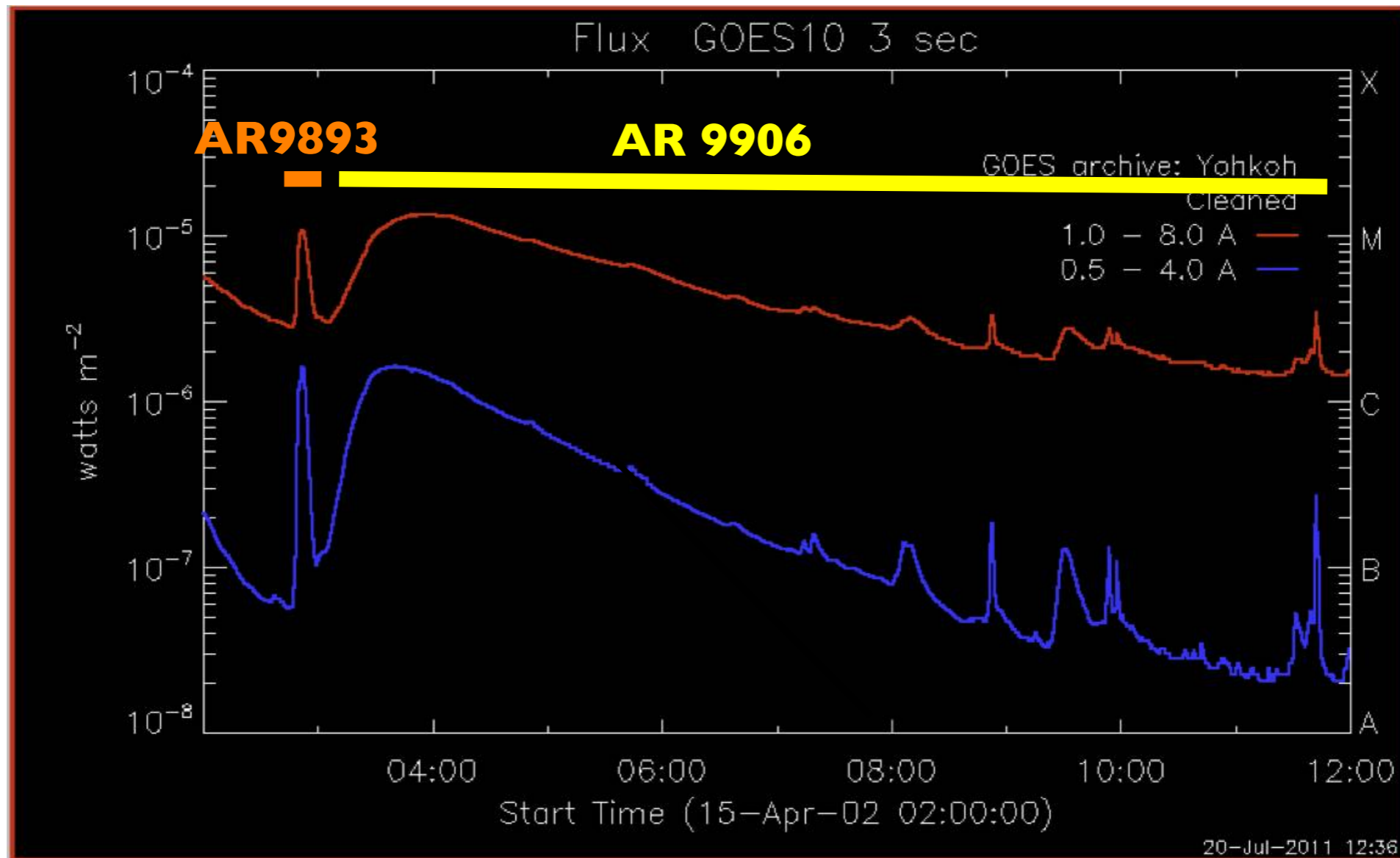
# RES SPIRIT

RES spectroheliograph of the SPIRIT instrumentation onboard the CORONAS-F satellite includes:

- X-ray channels to image the Sun in the monochromatic MgXII 8.42 Å line
- two channels to obtain monochromatic EUV solar images: 176–207 Å and 280–330 Å.

Wavelength range 280-330 Å includes lines of ions Fe VIII-XXII, Si VIII, XI, VII-VIII, Mg, Ni, Ca XVIII and XVIII, which are emitted in the temperature range from 0.8 MK to more than 20 MK.

# 15 April 2002 flare

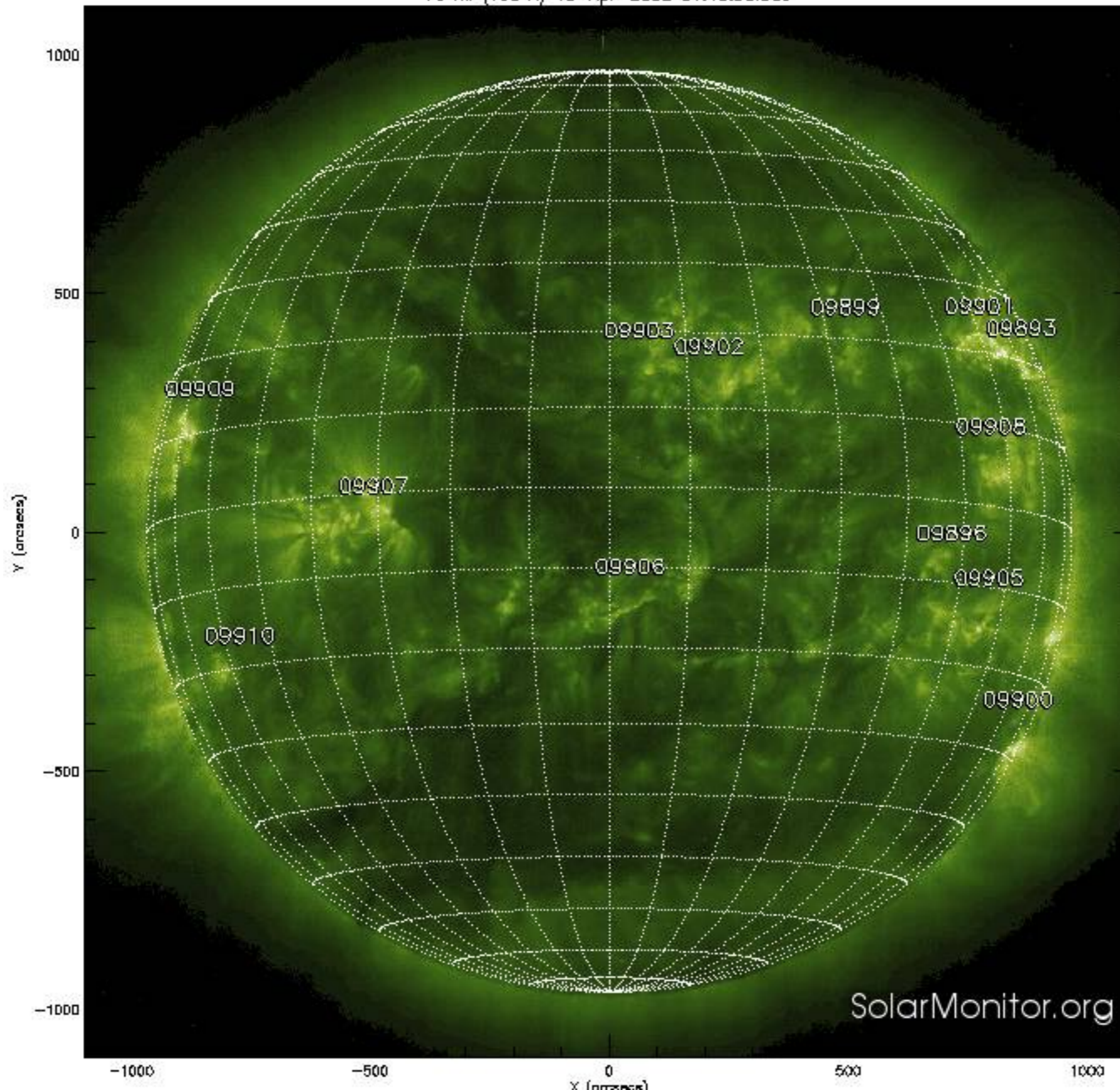


start: 2002-04-15 03:05:00  
max: 2002-04-15 03:55:00  
stop: 2002-04-15 05:06:00

location: S15W01  
AR: 9906  
GOES class: M1.2

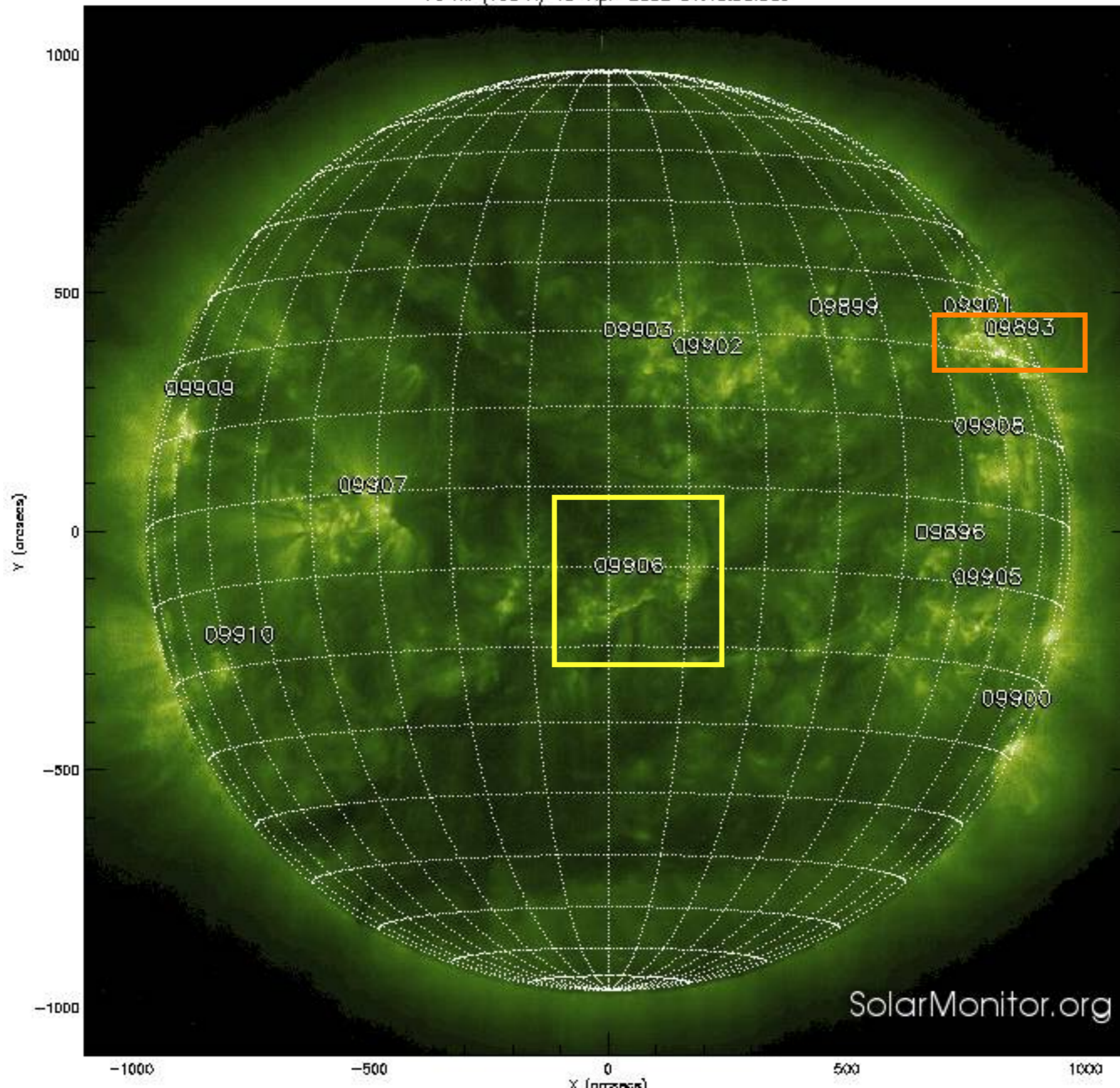
# 15 April 2002 flare

Fe XII (195 Å) 15-Apr-2002 01:13:36.589

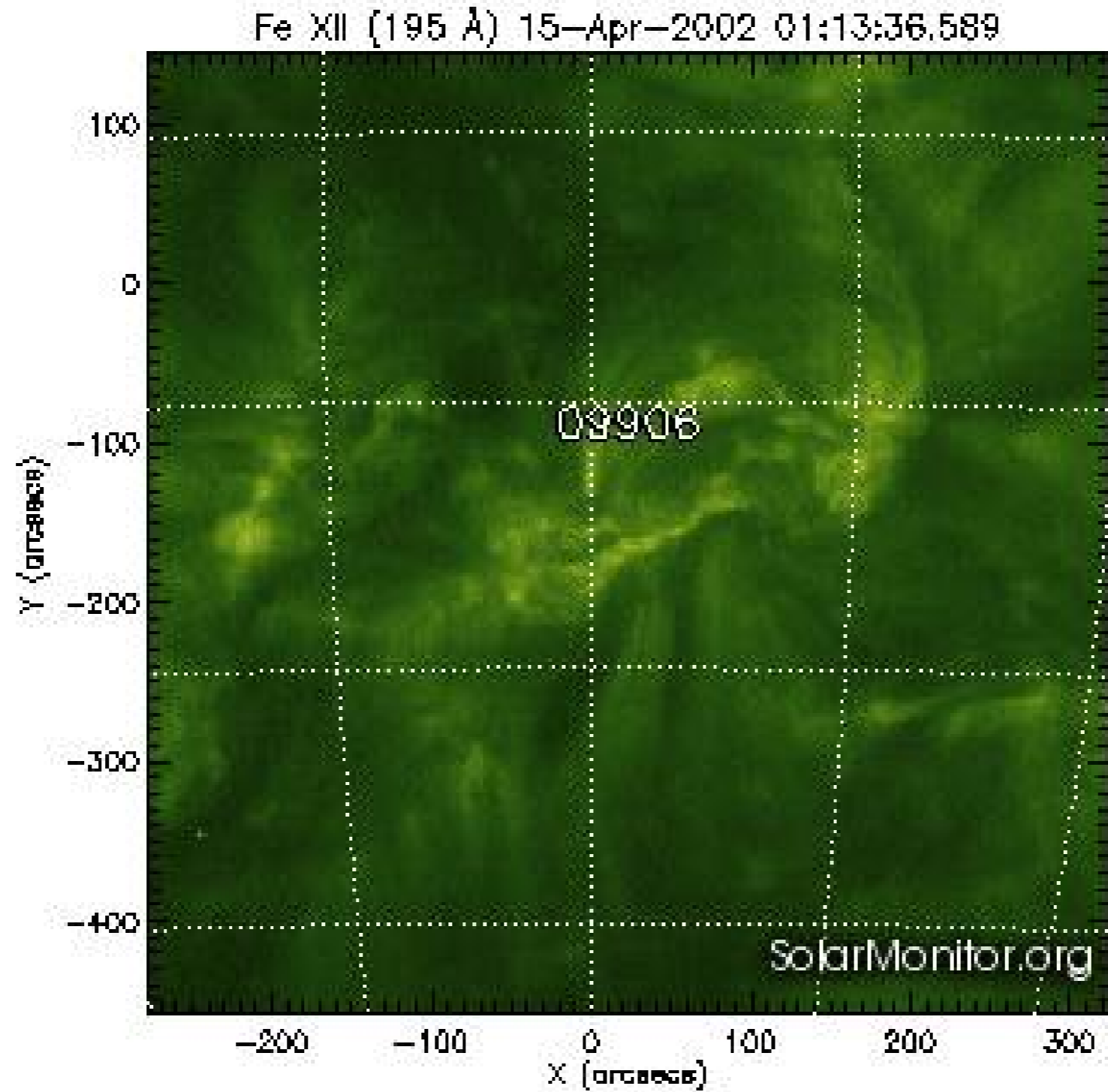


# 15 April 2002 flare

Fe XII (195 Å) 15-Apr-2002 01:13:36.589

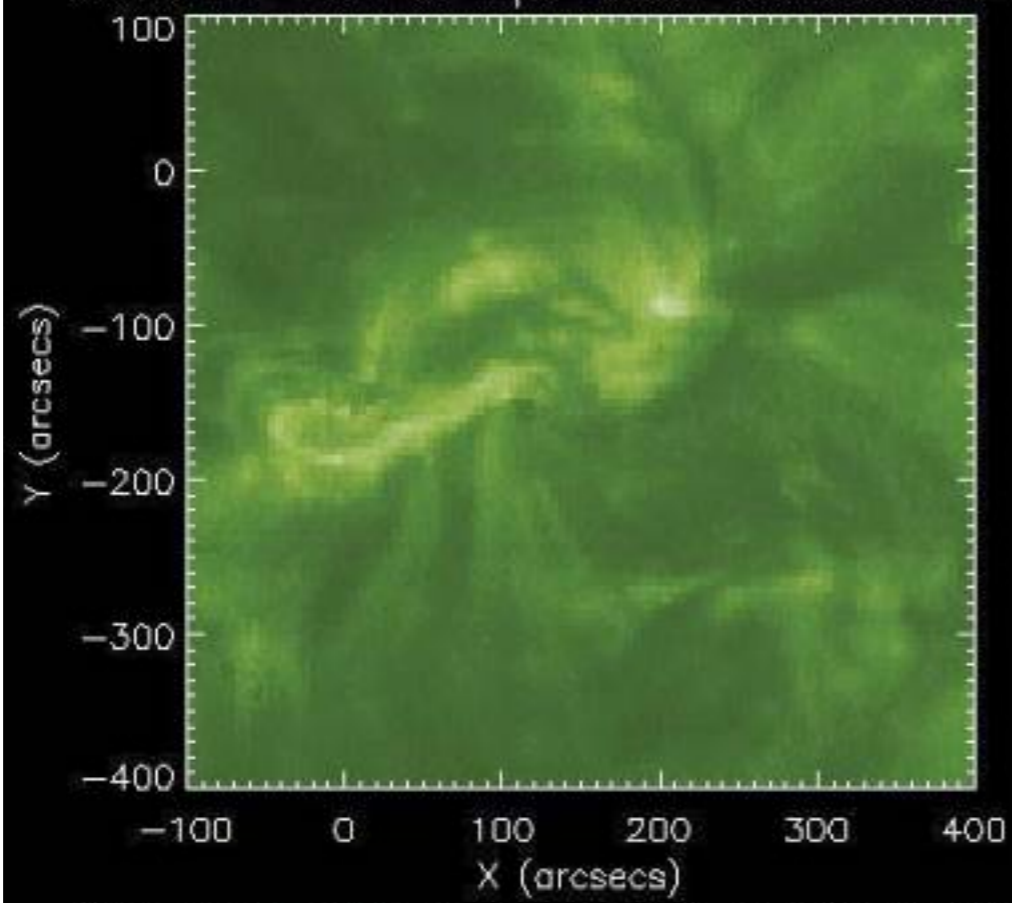


# 15 April 2002 flare

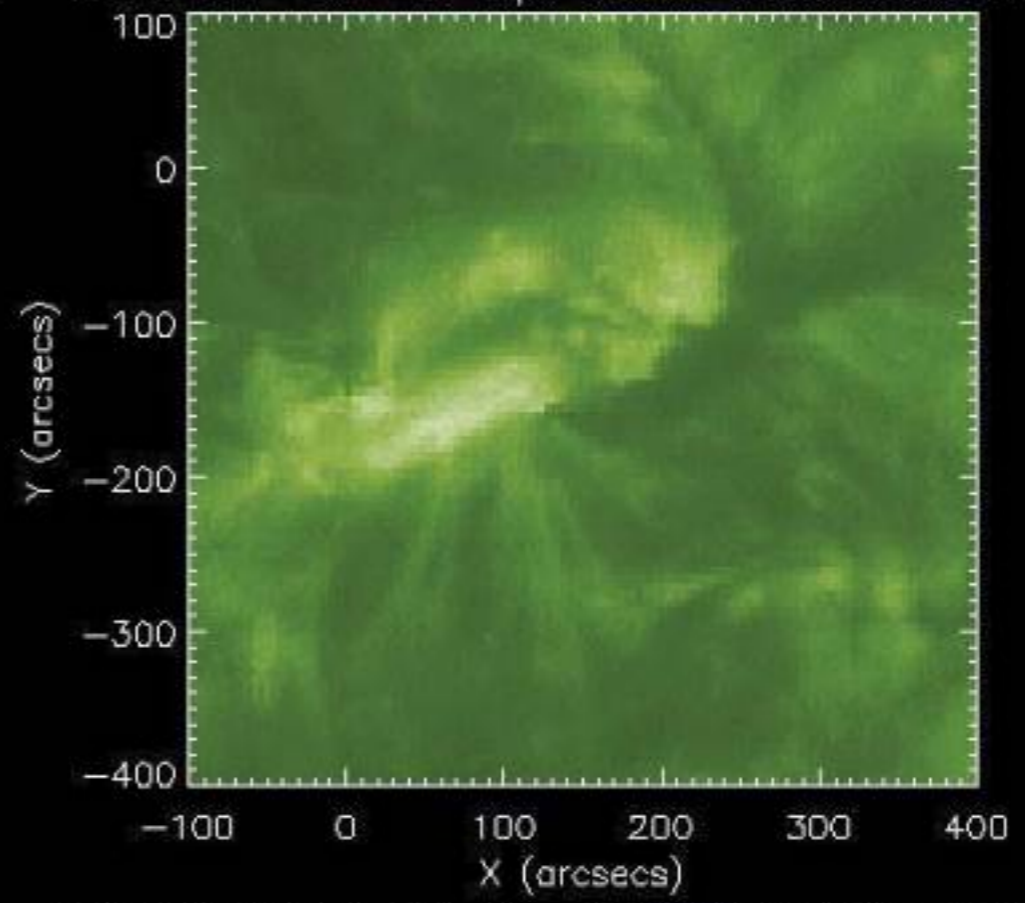




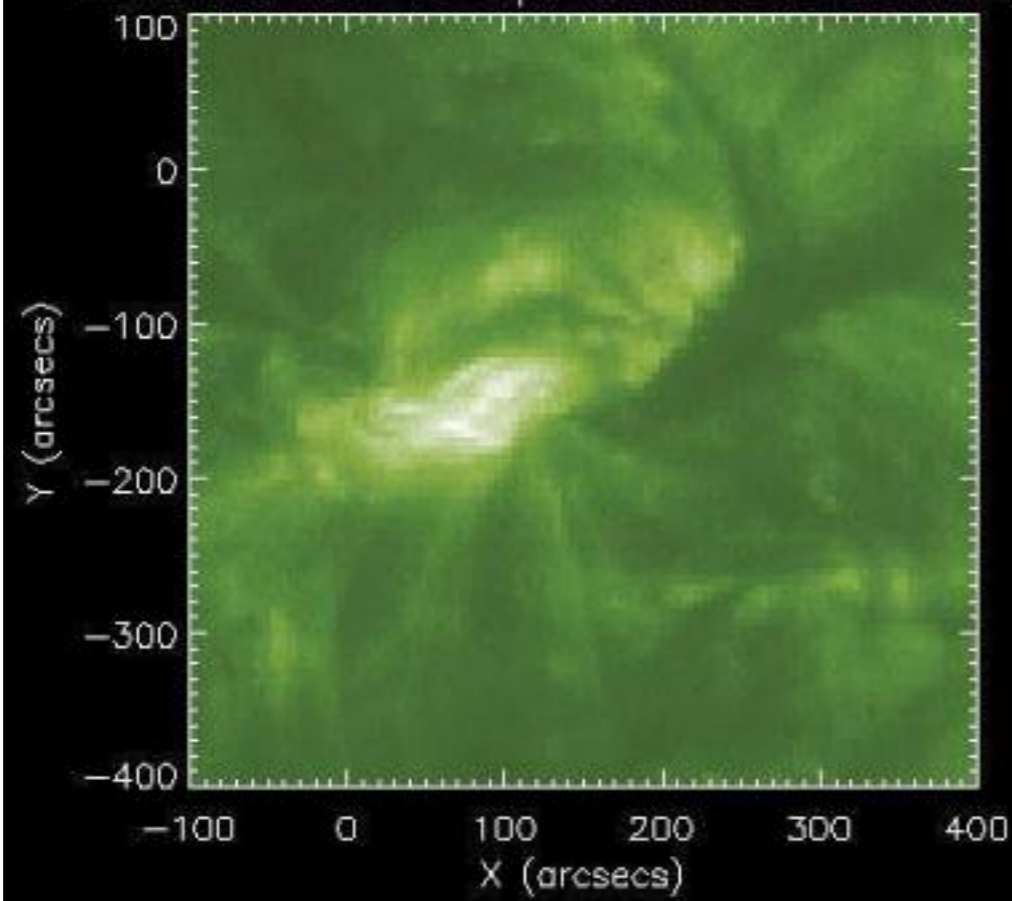
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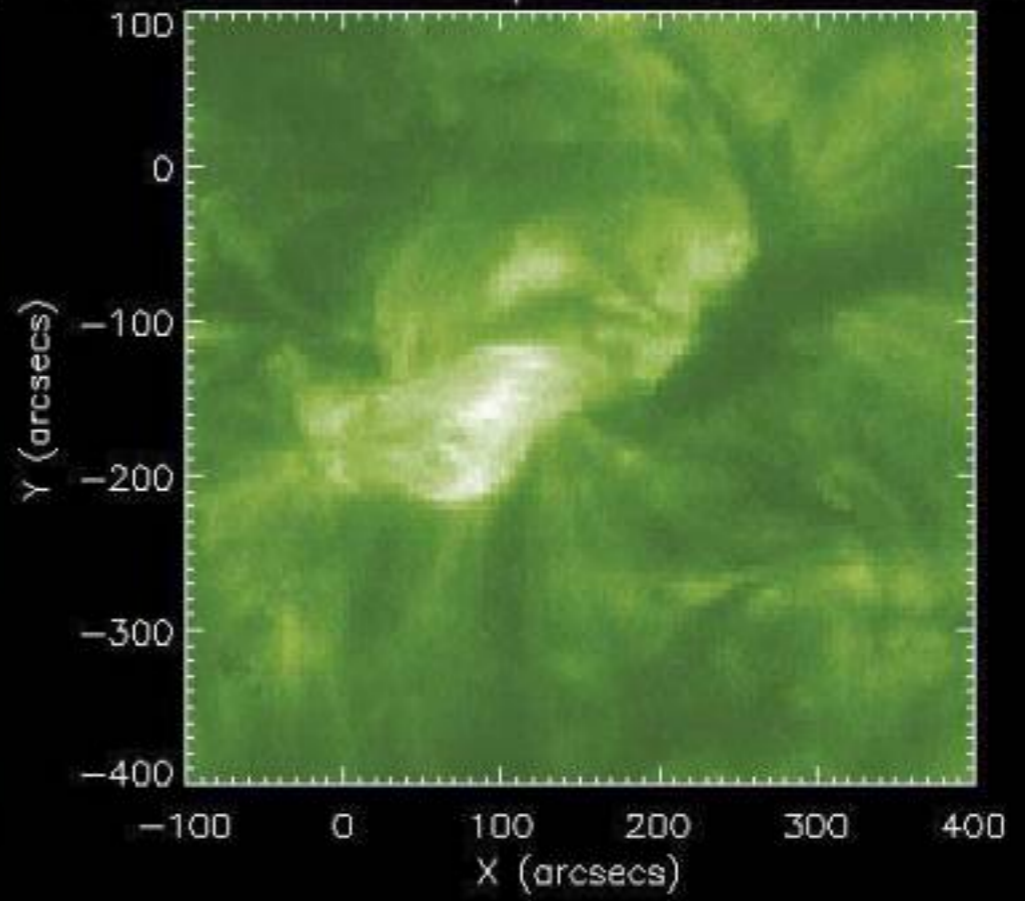
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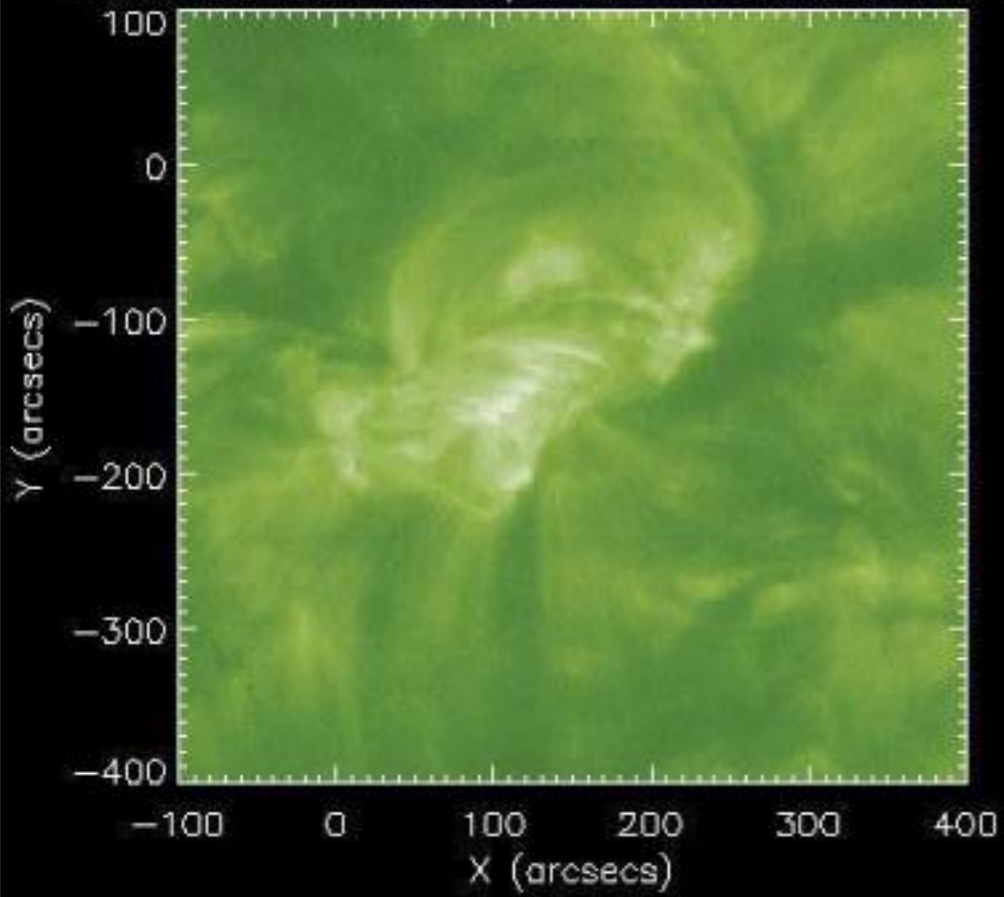
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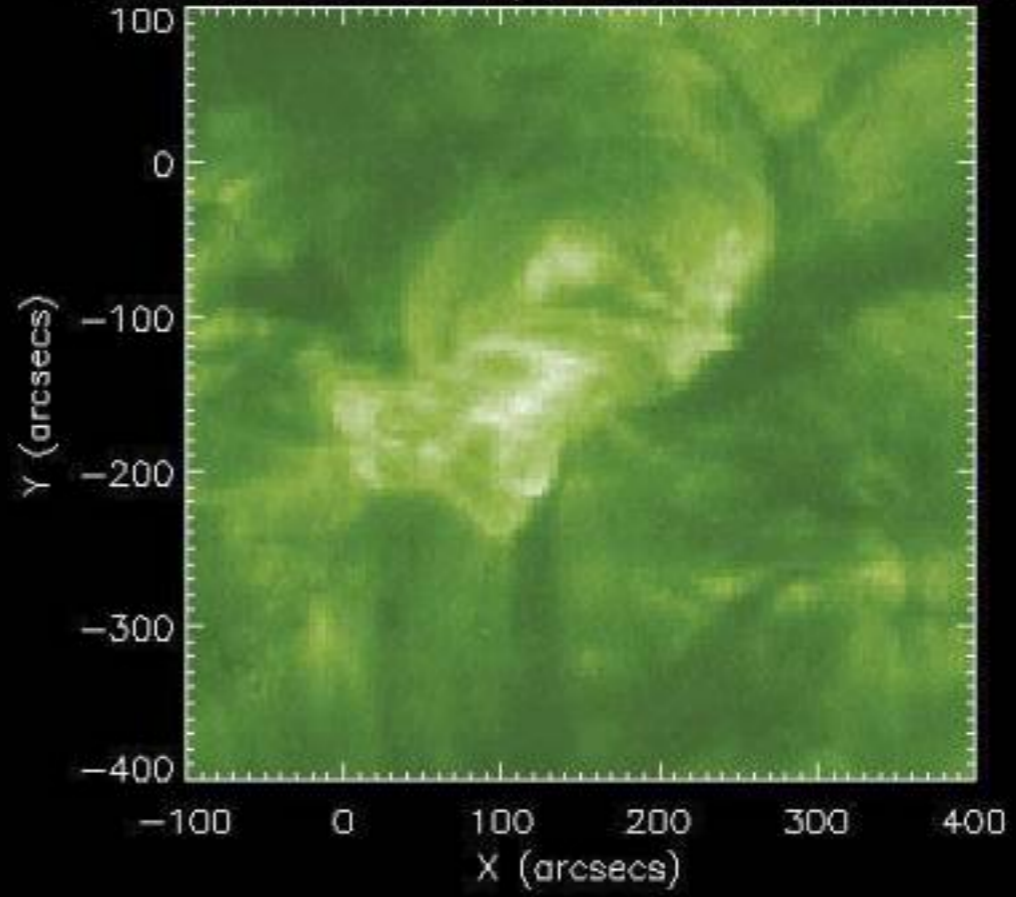
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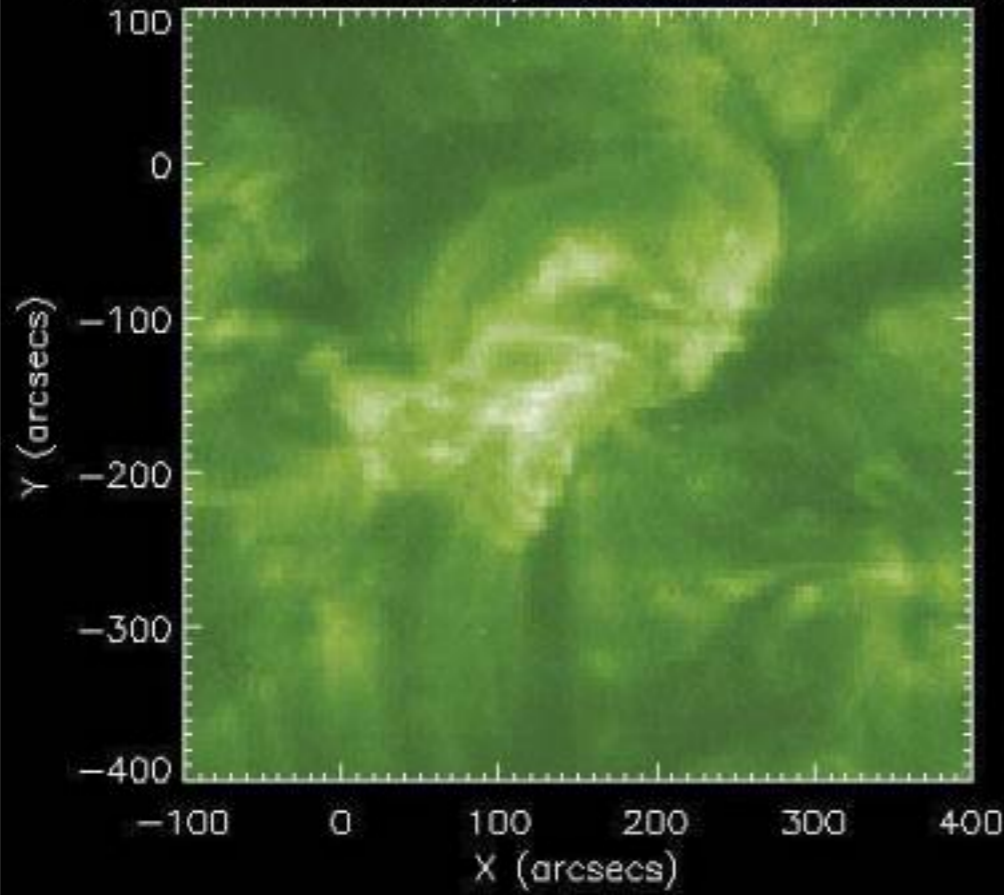
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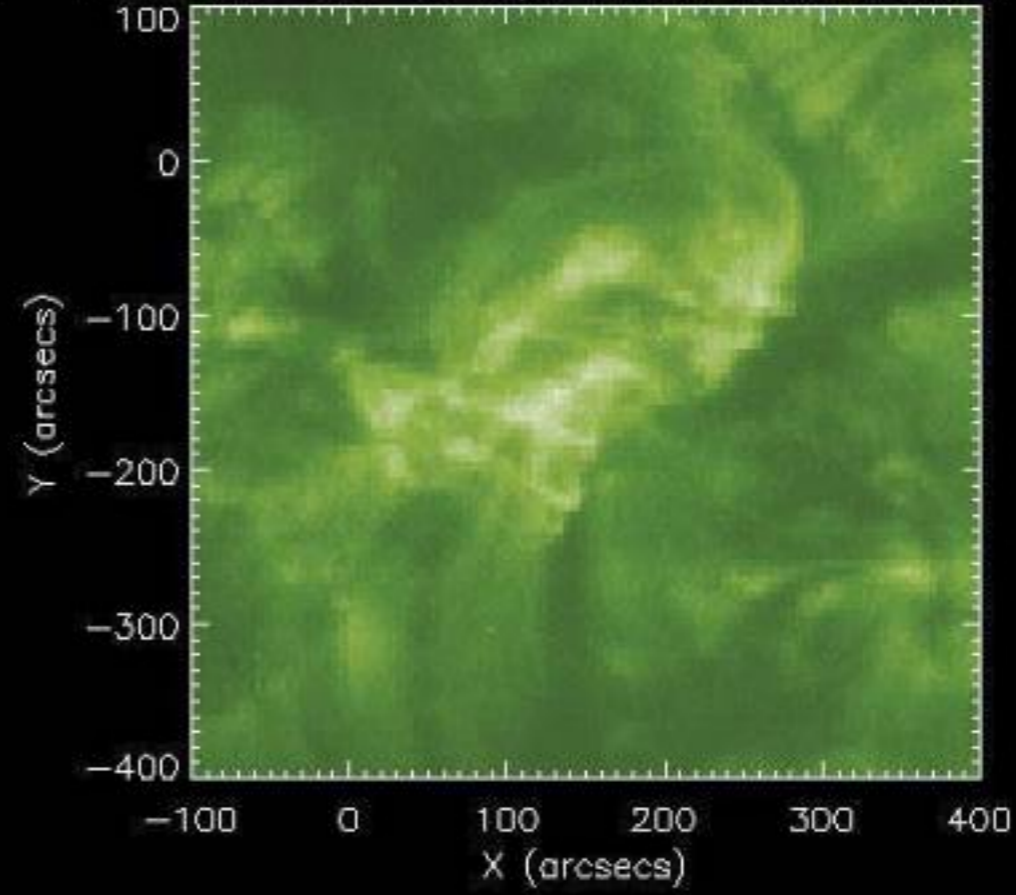
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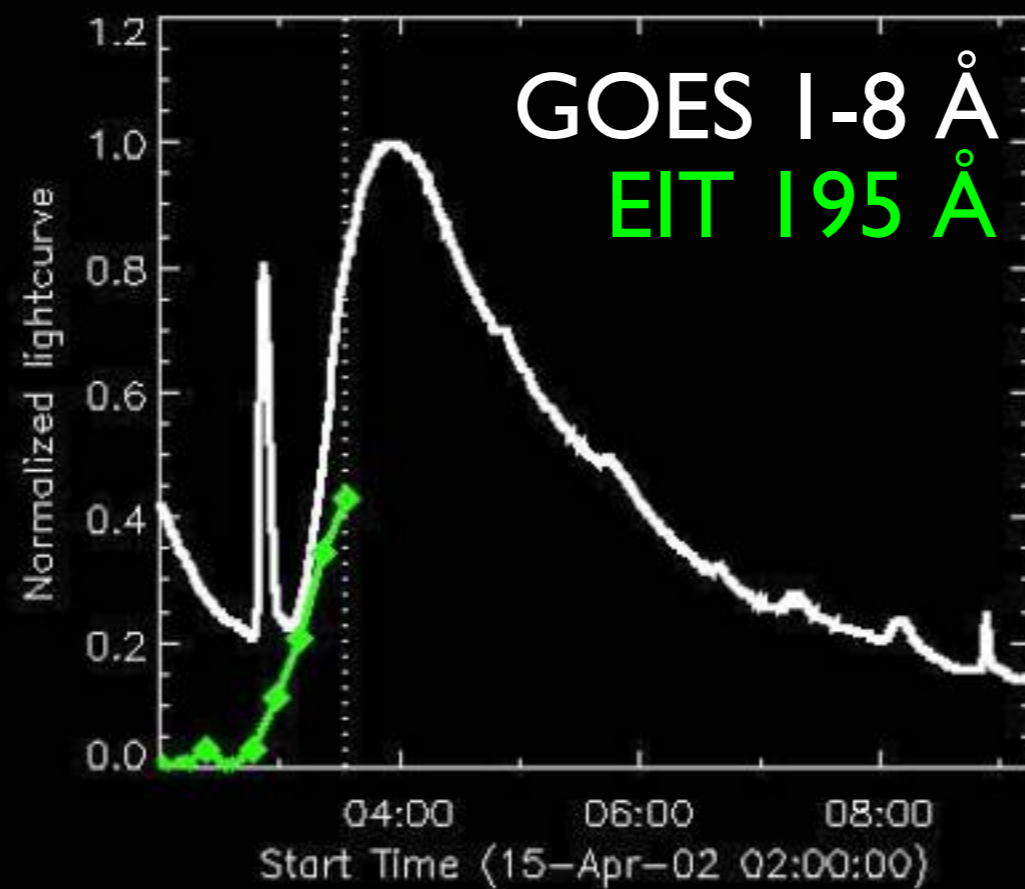
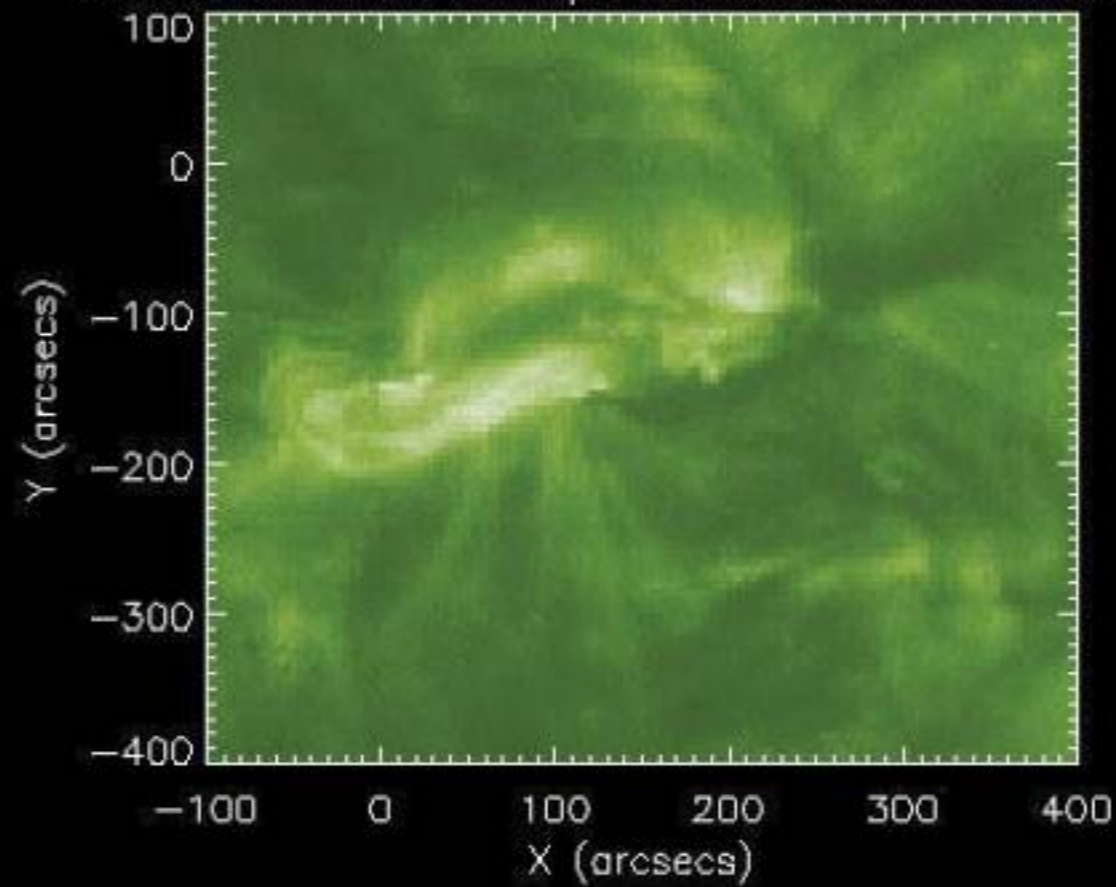
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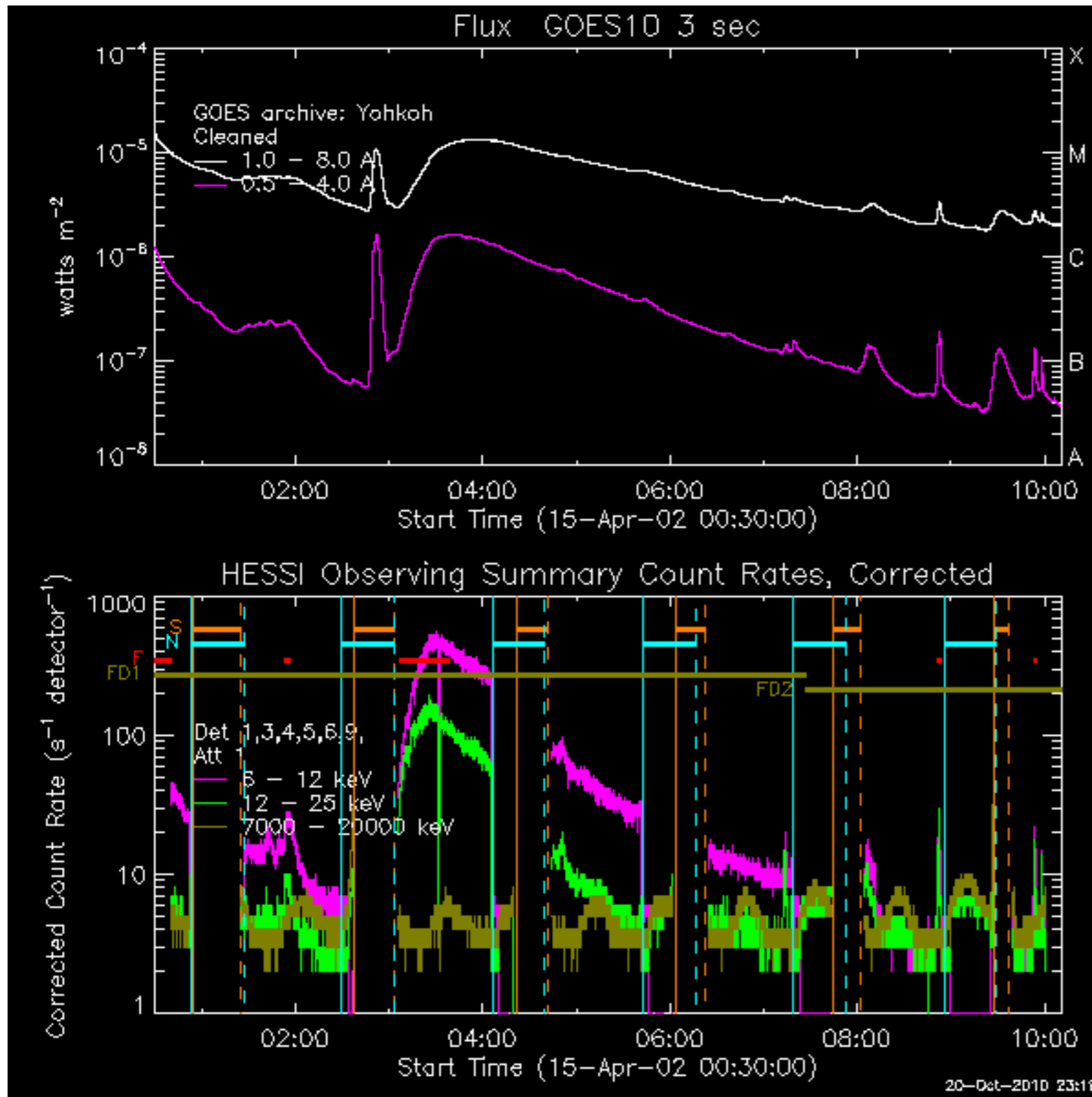
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SOHO EIT 195 15-Apr-2002 03:36:05.433



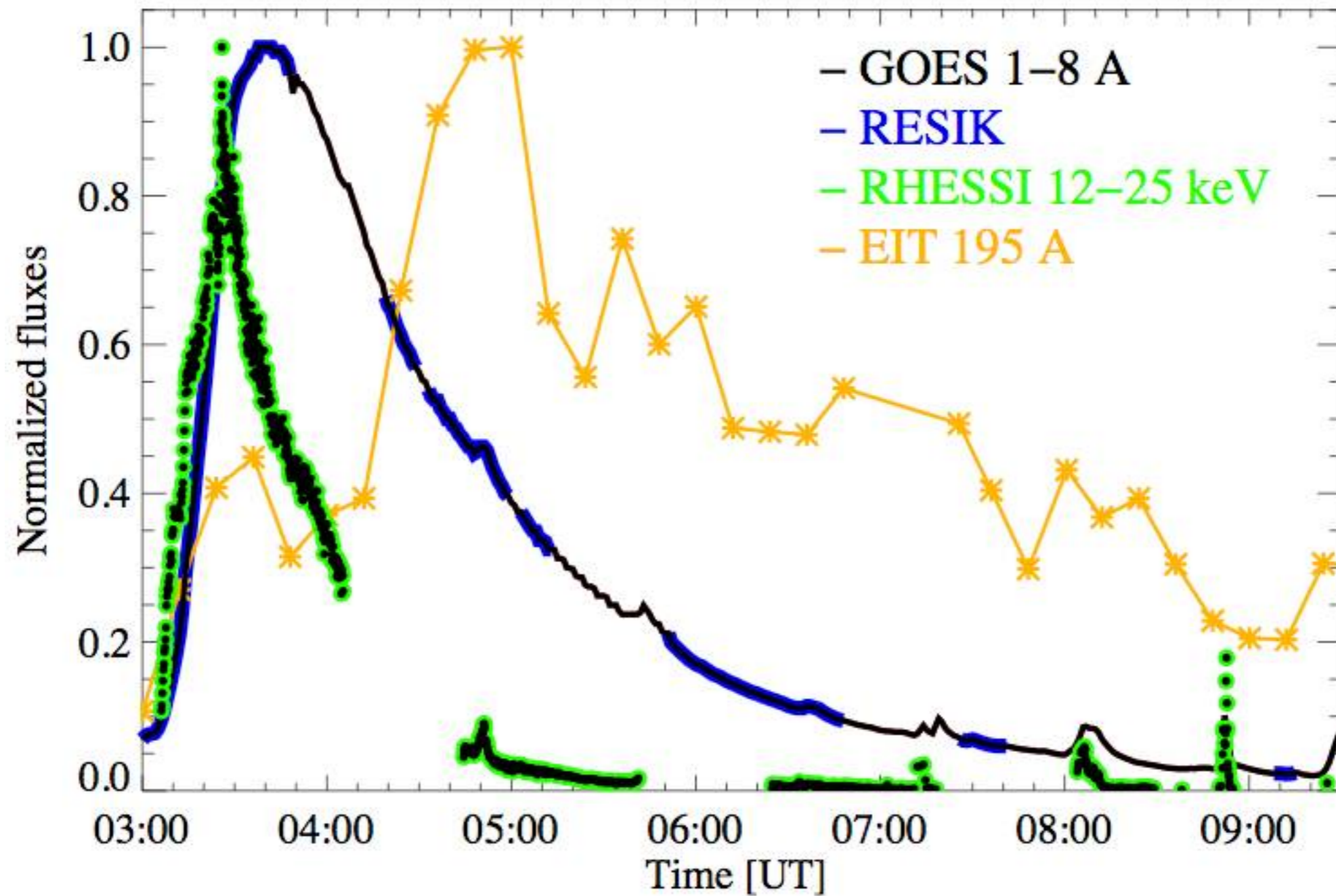
# RHESSI - 15 April 2002 flare



Almost entire rising phase

About 5 hours of decay in an energy range of 6-12 keV

# GOES, RHESSI and EIT/SoHO light curves

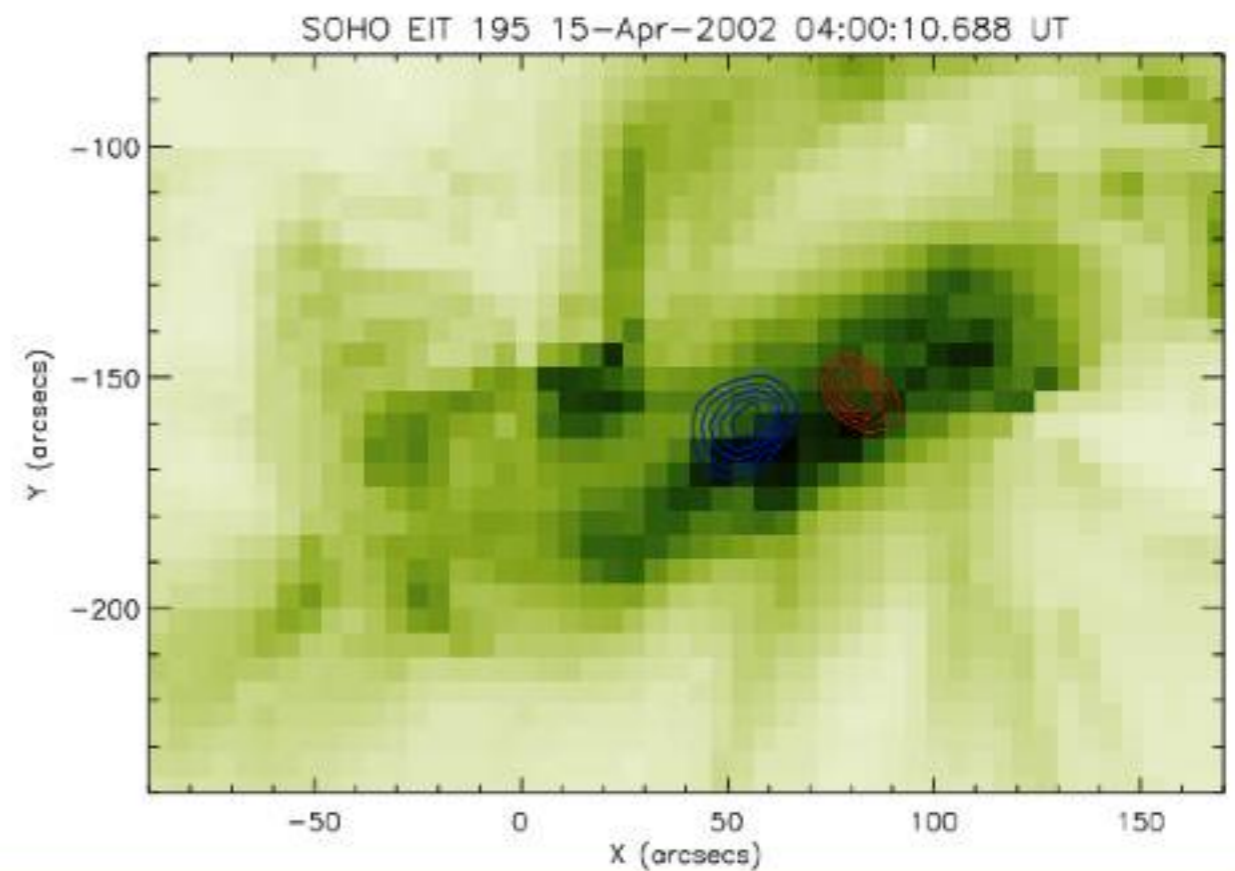
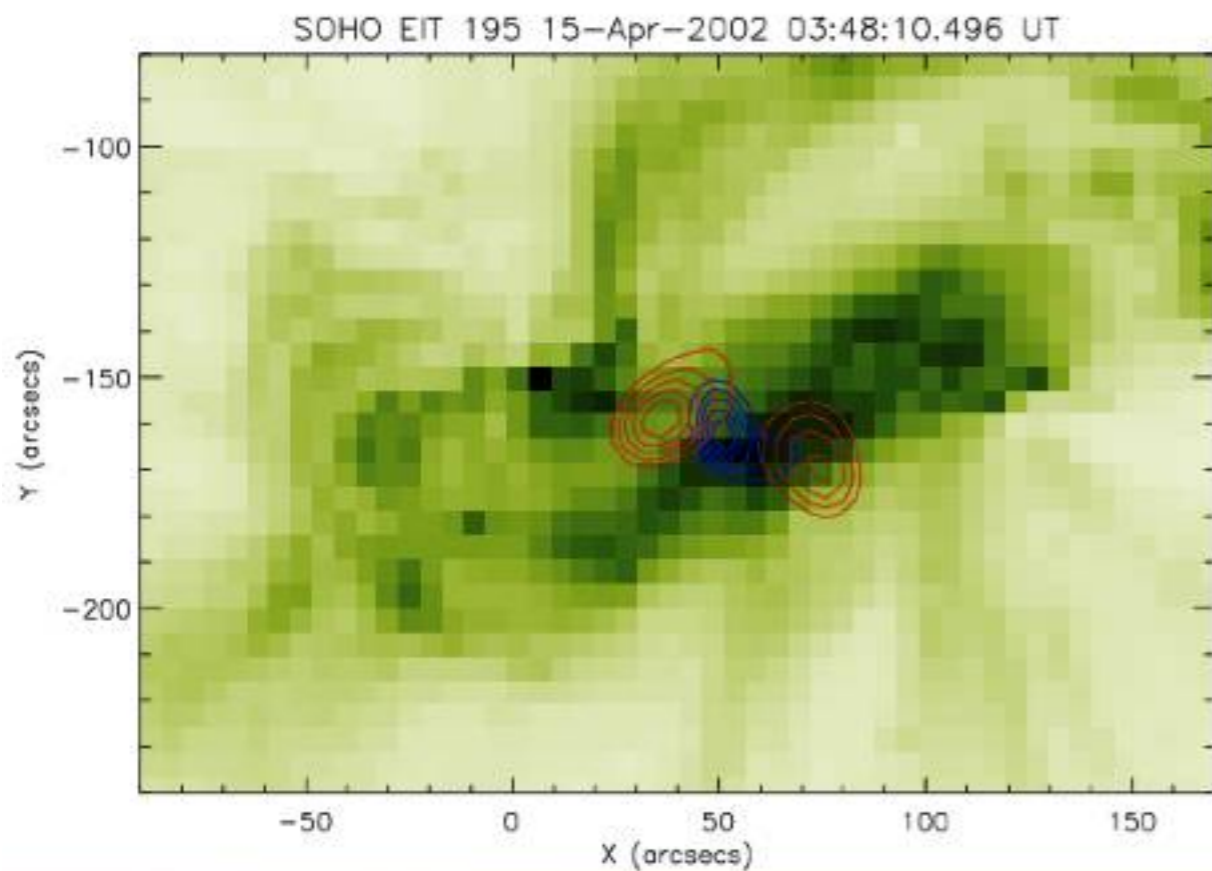
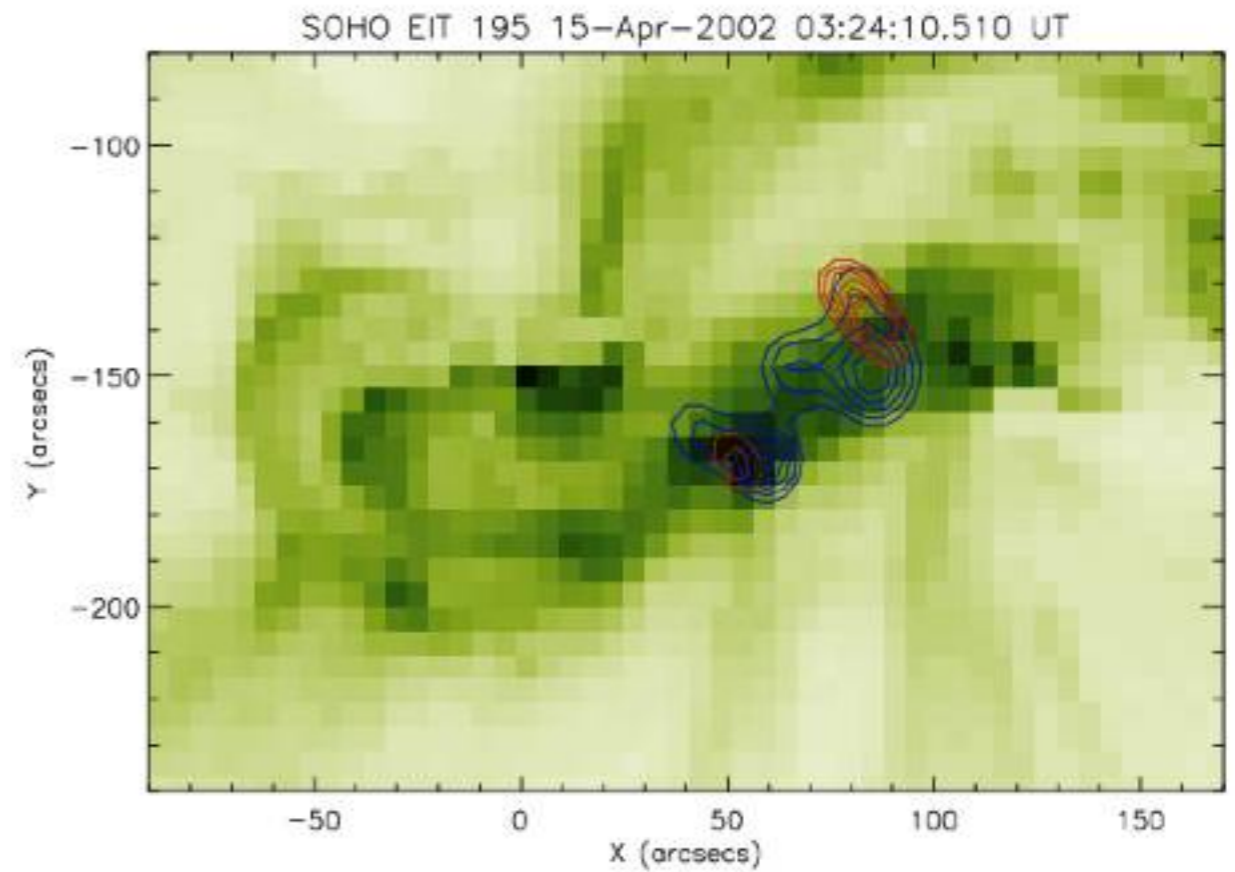
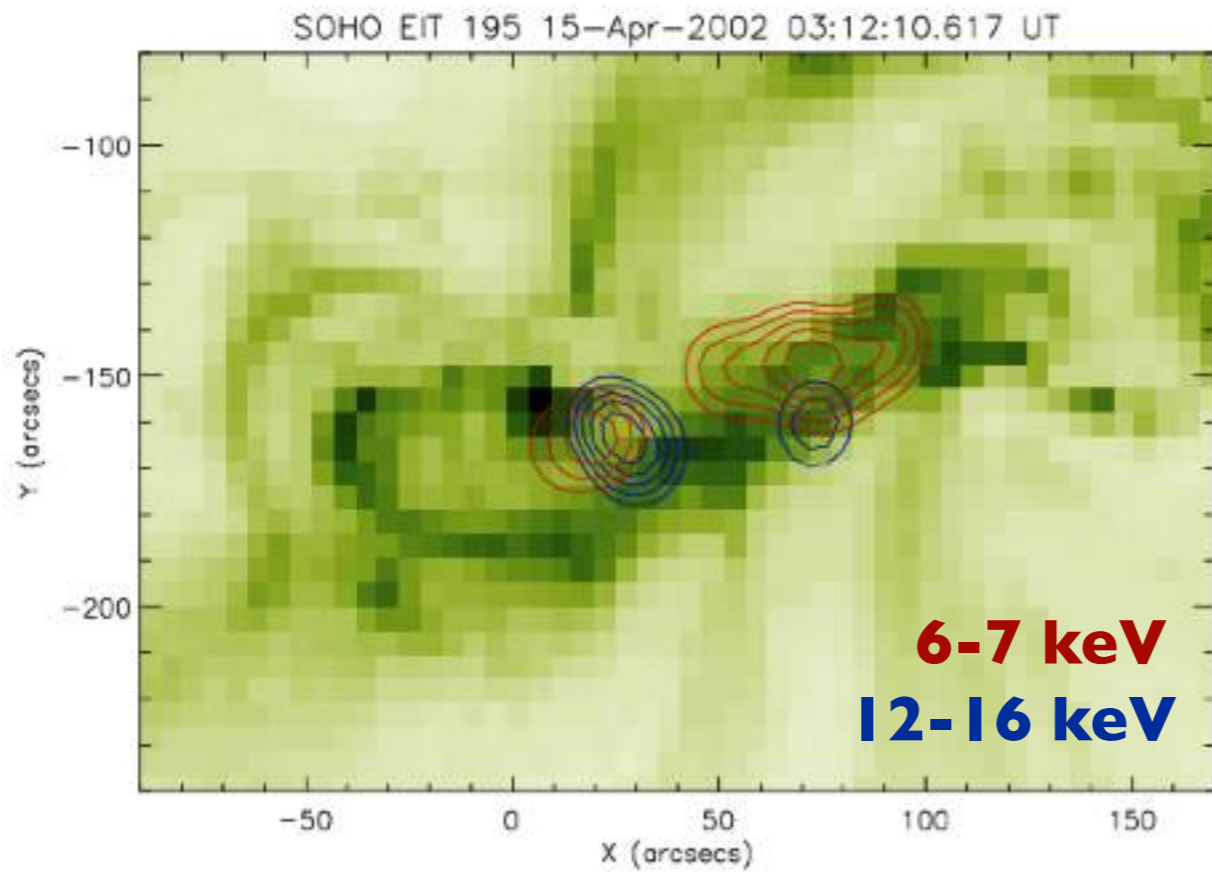


$t_{\text{RHESSI\_GOES}} \sim 20 \text{ min}$

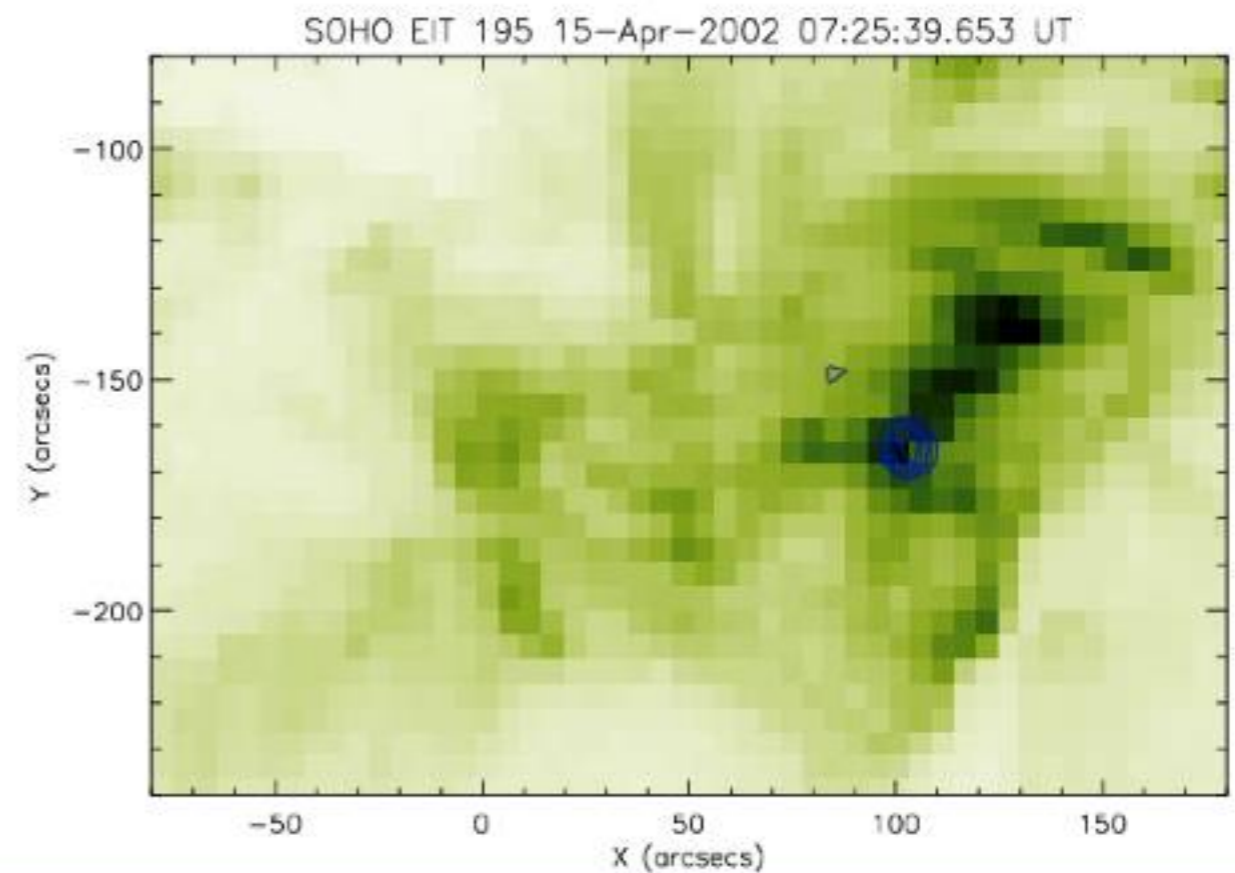
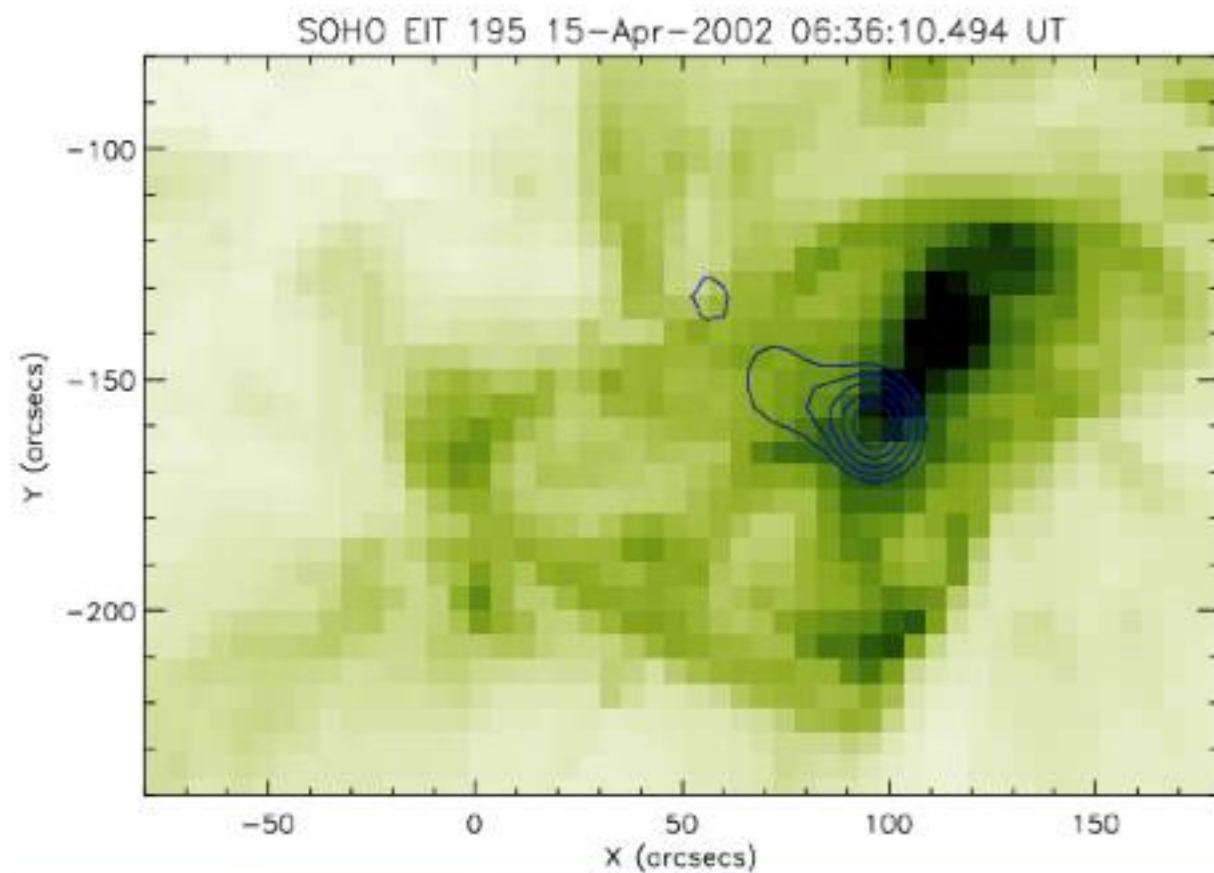
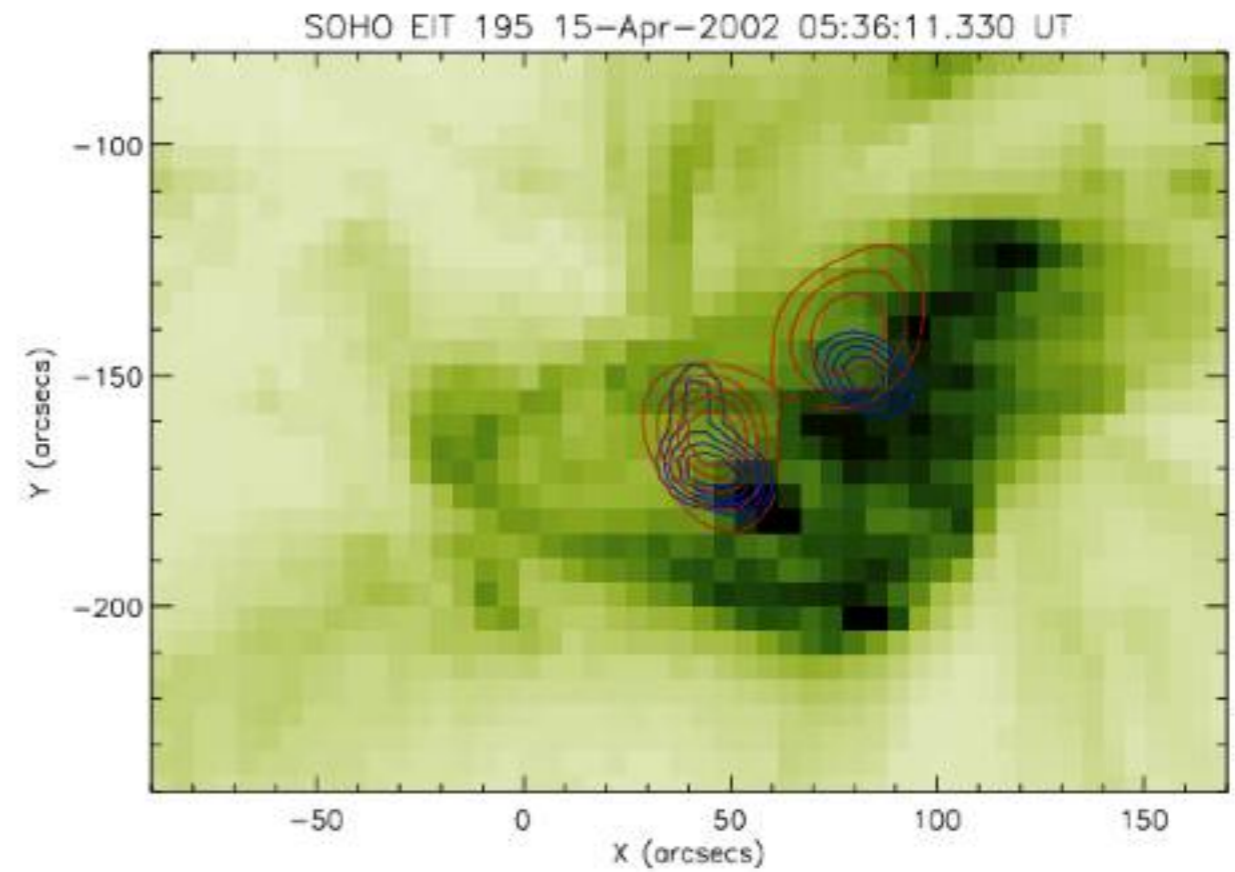
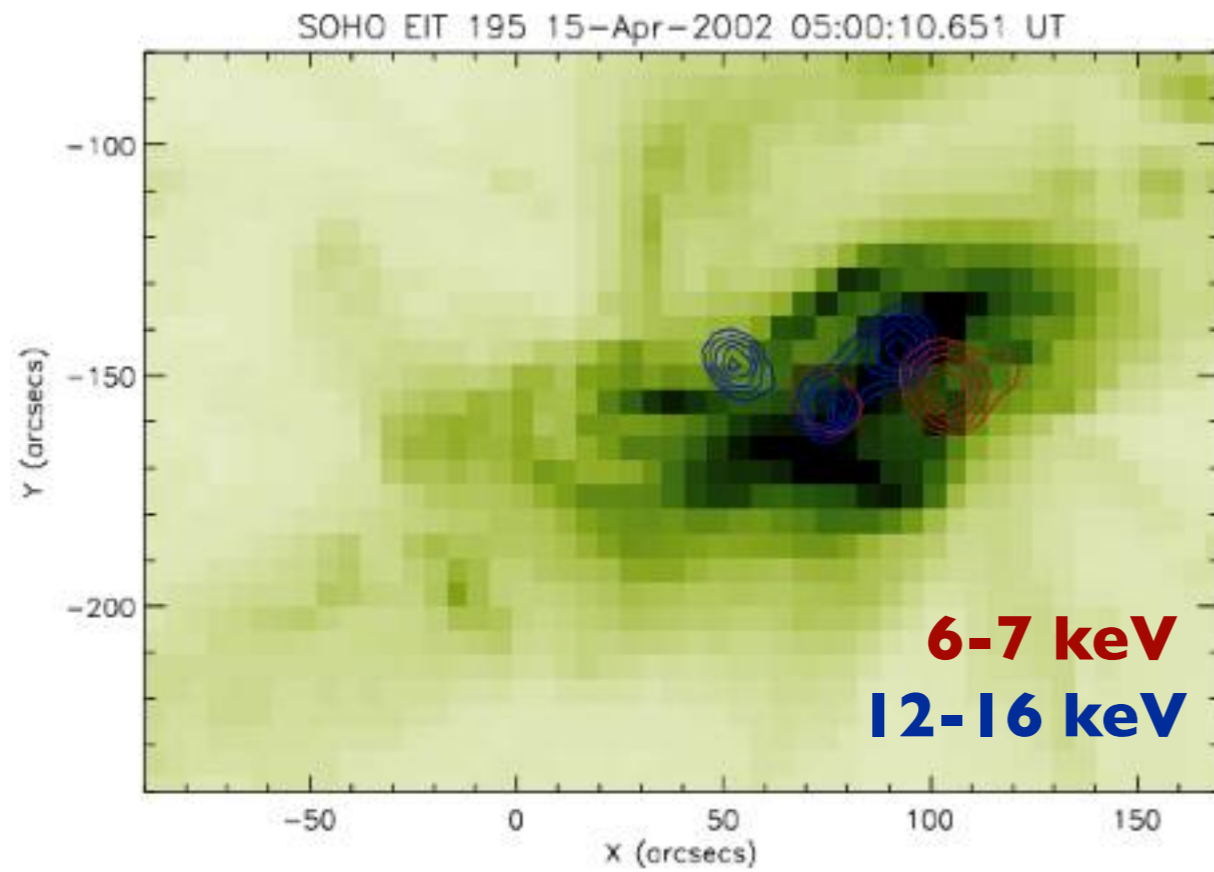
$t_{\text{RHESSI\_EIT}} \sim 1 \text{ h } 30 \text{ min}$

$t_{\text{GOES\_EIT}} \sim 1 \text{ h } 10 \text{ min}$

# RHESSI - 15 April 2002 flare

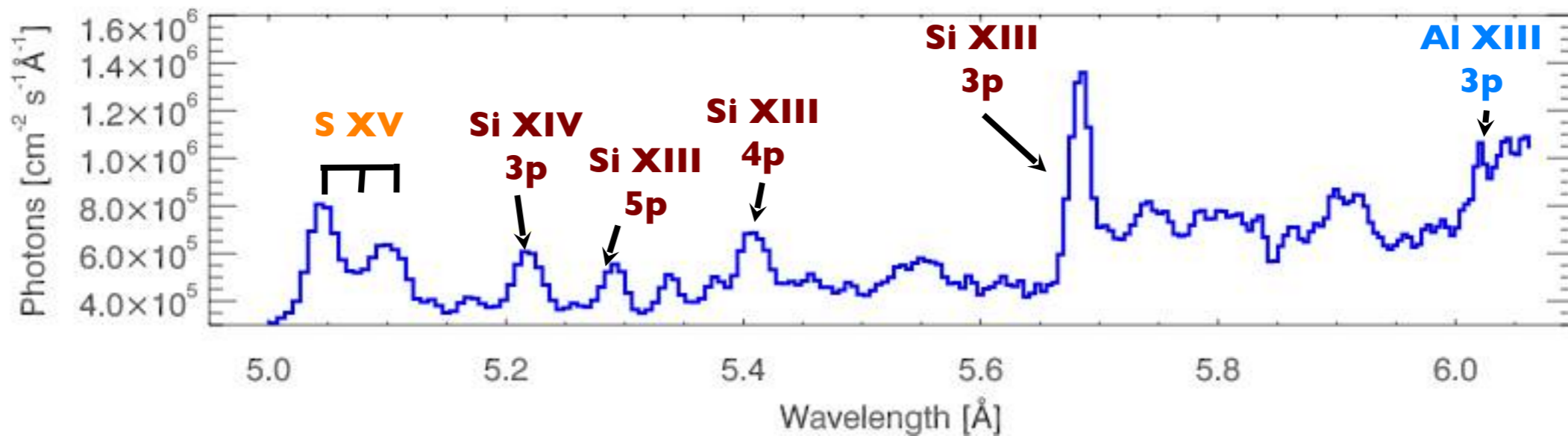
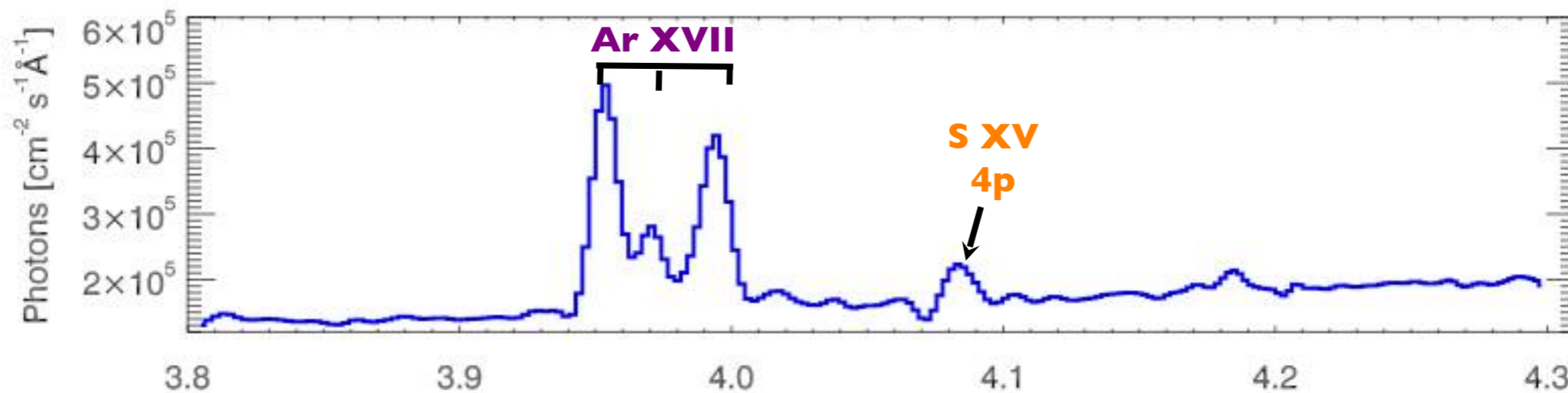
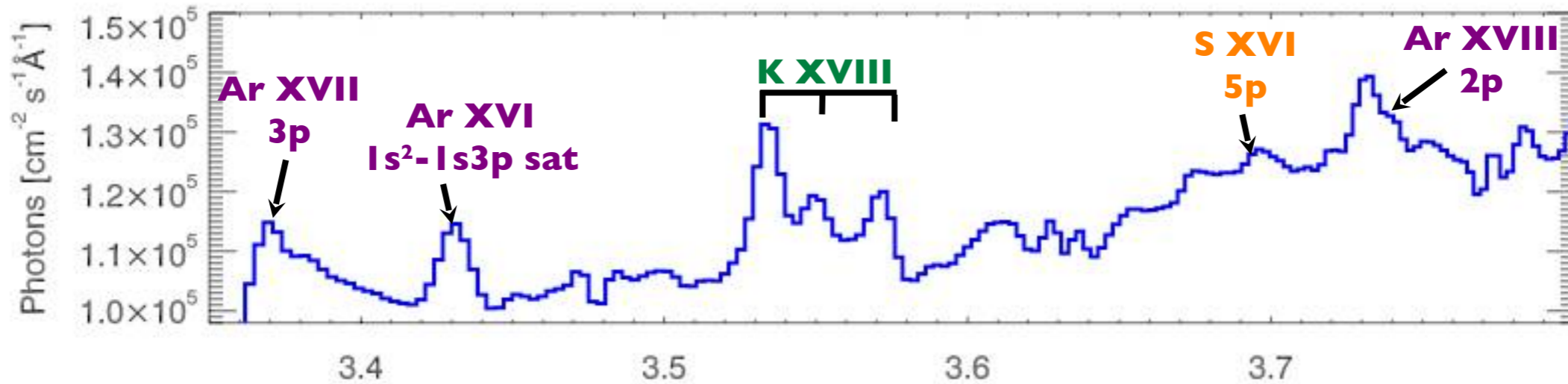


# RHESSI - 15 April 2002 flare



# RESIK - 15 April 2002 flare

No record on channel 3 is associated with instrumental problems.





# DEM - Withbroe-Sylwester method

DEM distribution  $\varphi_{j+1}(T)$  calculated from preceding one  $\varphi_j(T)$  using the following expression:

$$\varphi_{j+1}(T) = \varphi_j(T) \frac{\sum_{i=1}^k c_i w_i(T)}{\sum_{i=1}^k w_i(T)}$$

$w_i(T)$  is the weight function defined by the following relationship:

$$w_i(T) = f_i(T) \varphi_j(T) \frac{\int_0^\infty f_i(T) \varphi_j(T) dT}{\int_0^\infty [f_i(T) \varphi_j(T)]^2 dT} \left[ \frac{|F_{oi} - F_{ci}|}{\sigma_i} + 1 \right]^a$$

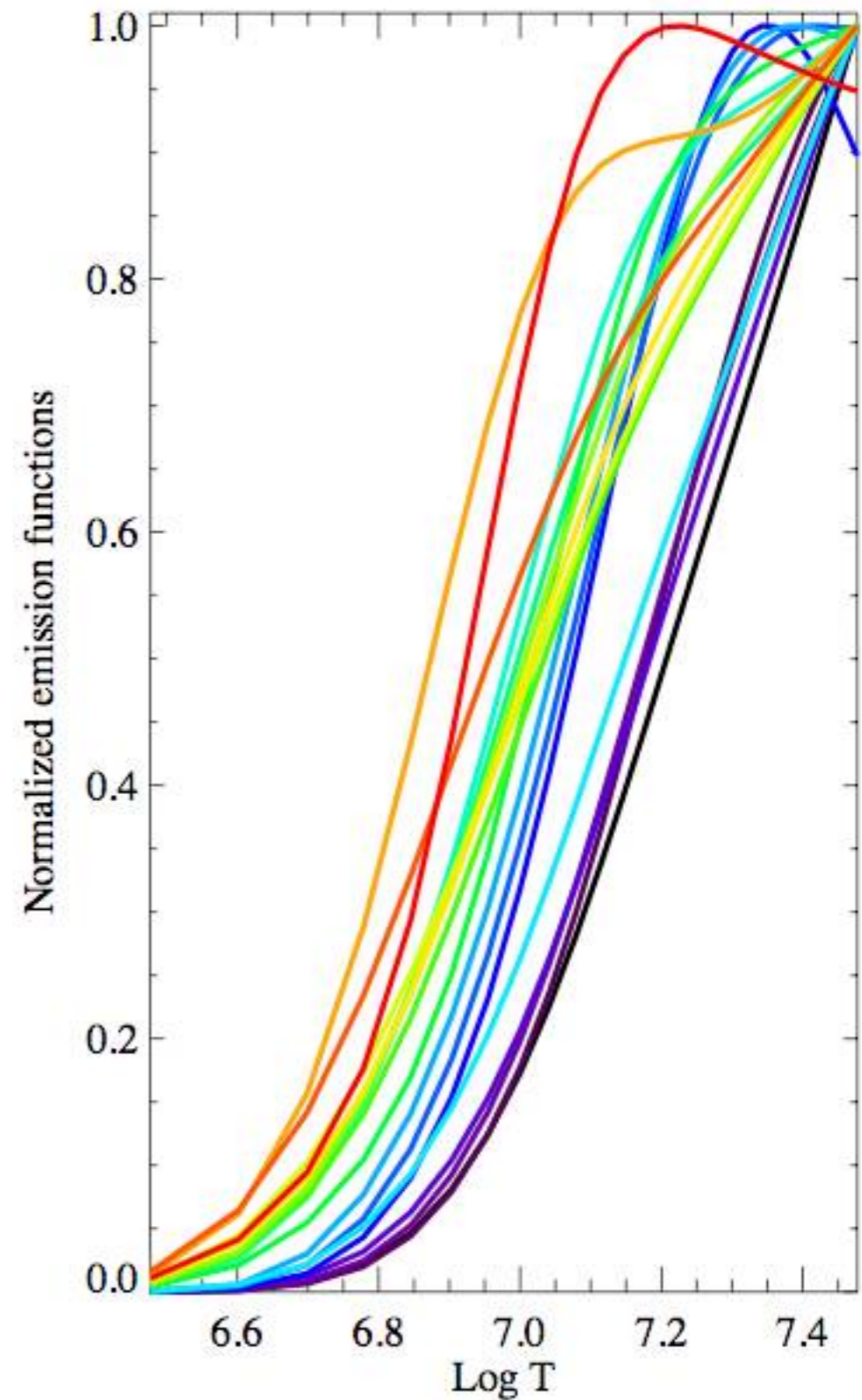
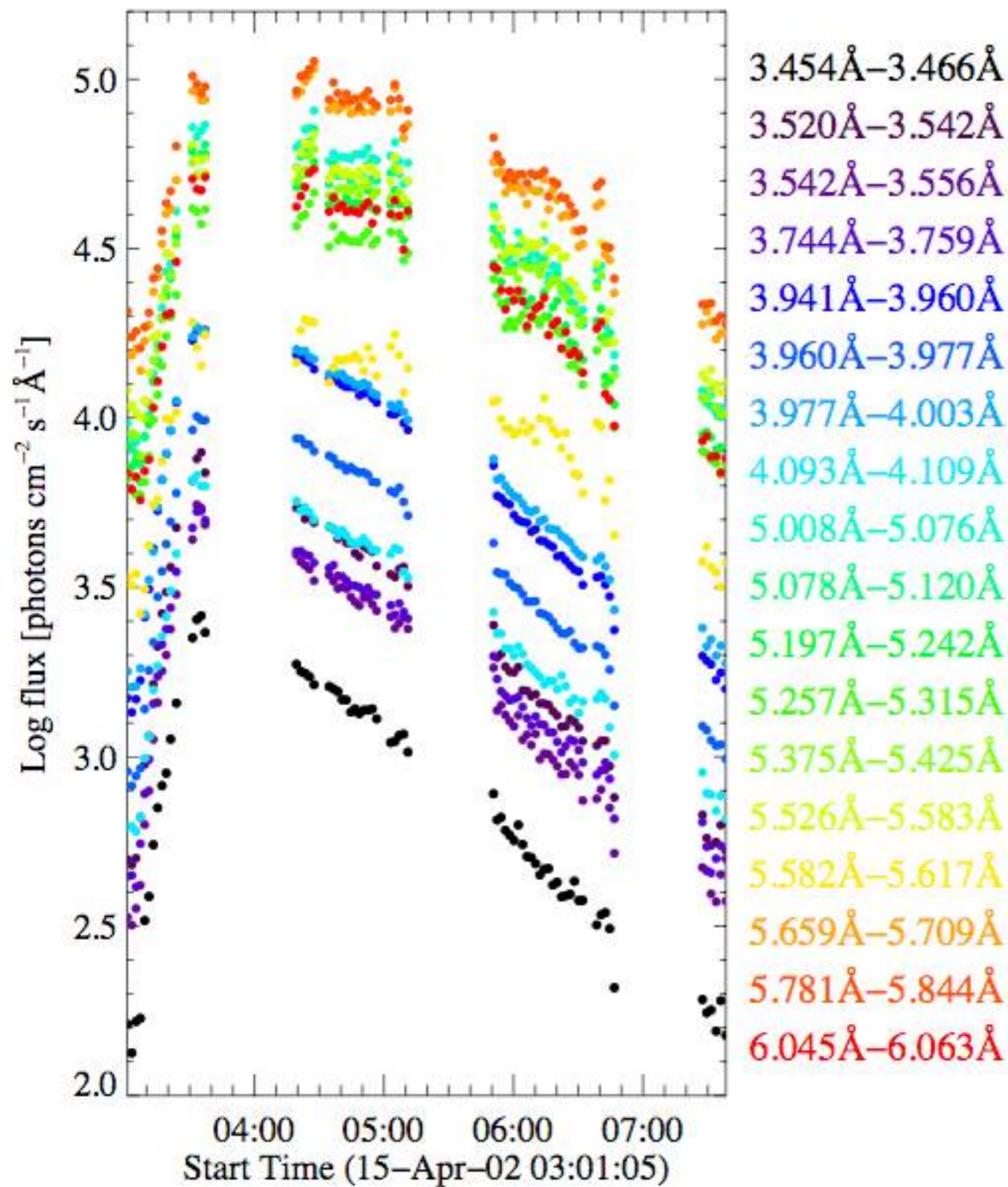
$c_i$  is the correction factor taken as  $c_i = F_{oi}/F_{ci}$

$f_i(T)$  is the emission function in line/band  $i$

$F_{ci}$  is the flux calculated (in  $j^{\text{th}}$  iteration) as:

$$F_{ci} = \int_0^\infty f_i(T) \varphi_j(T) dT$$

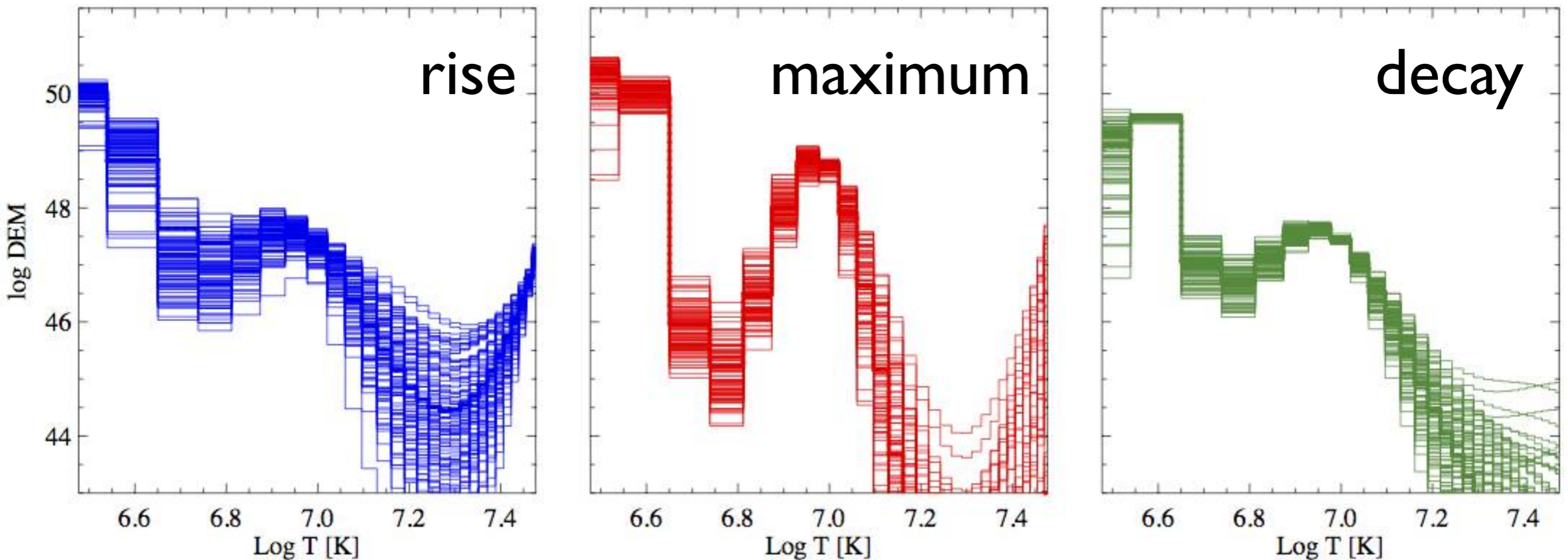
# DEM - input data



# DEM - input data

- Temperature range = 3MK - 30 MK
- 3000 iterations
- CHIANTI package ver. 6.0.1
- ionization equilibrium by Mazzotta et al. 1998
- coronal abundances by Feldman
- new abundances for S, Si, K

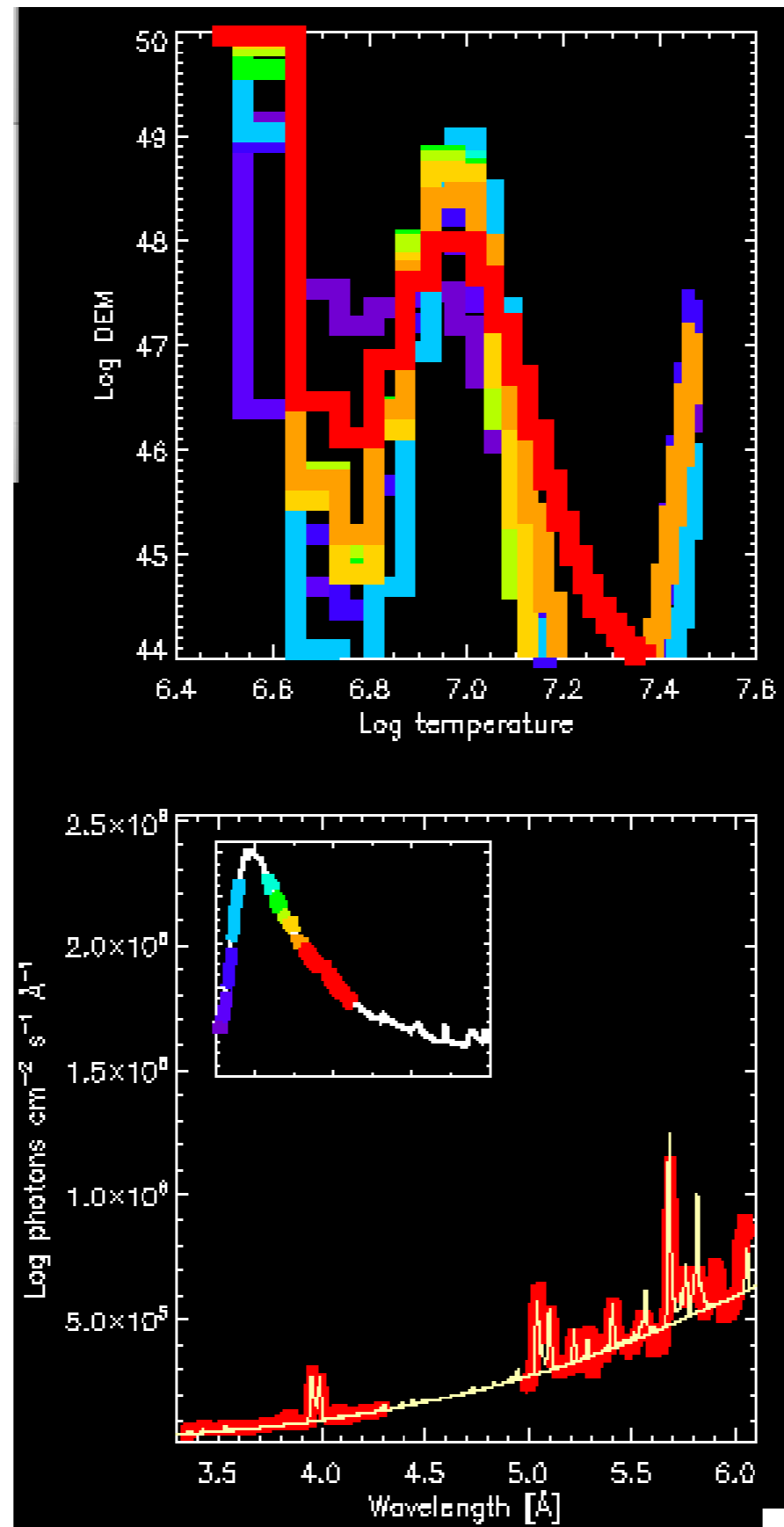
# DEM - results



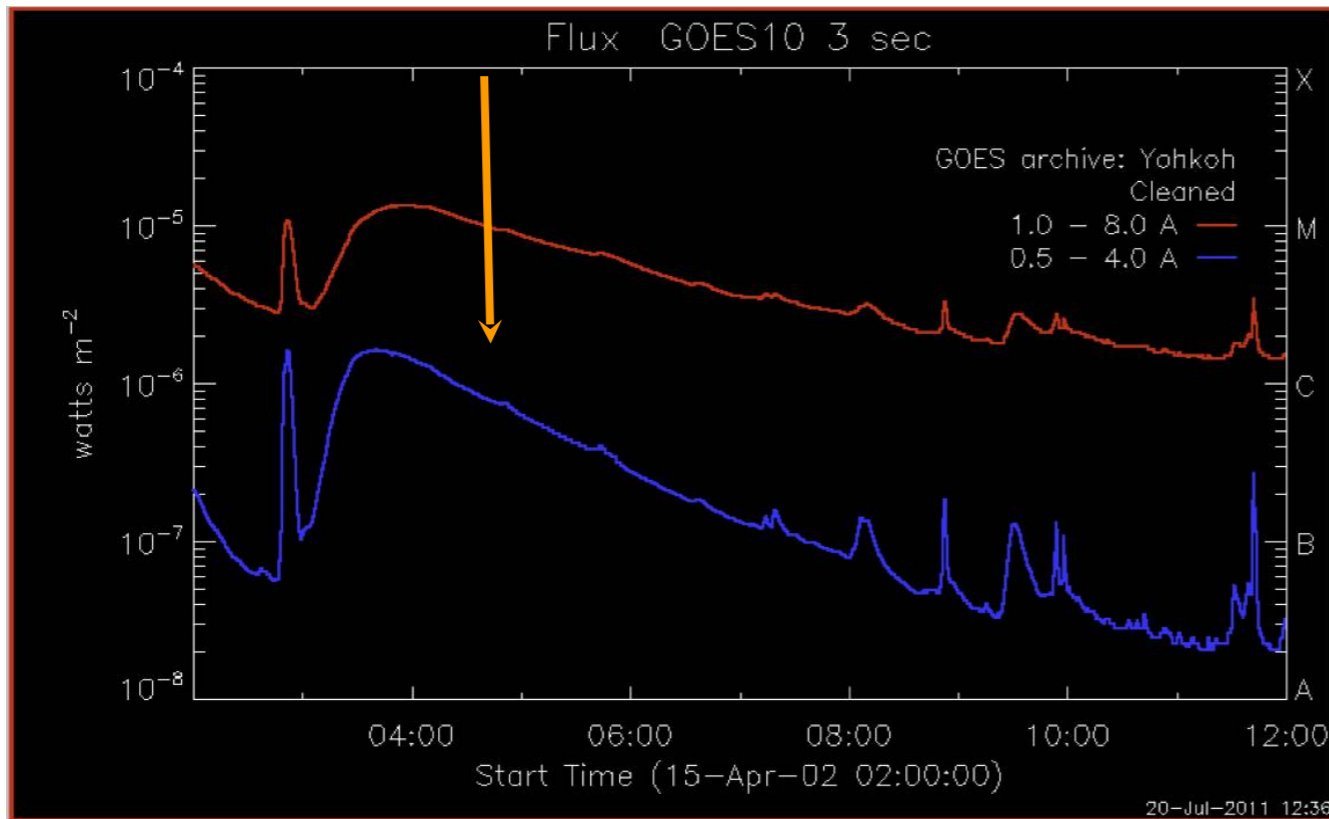
The DEM distributions as calculated for rise, maximum and decay phase of 15 April 2002 flare. They have been obtained from 100 Monte-Carlo realizations of DEM calculations.

The calculated DEM distributions are two- component independent on the evolutionary phase.

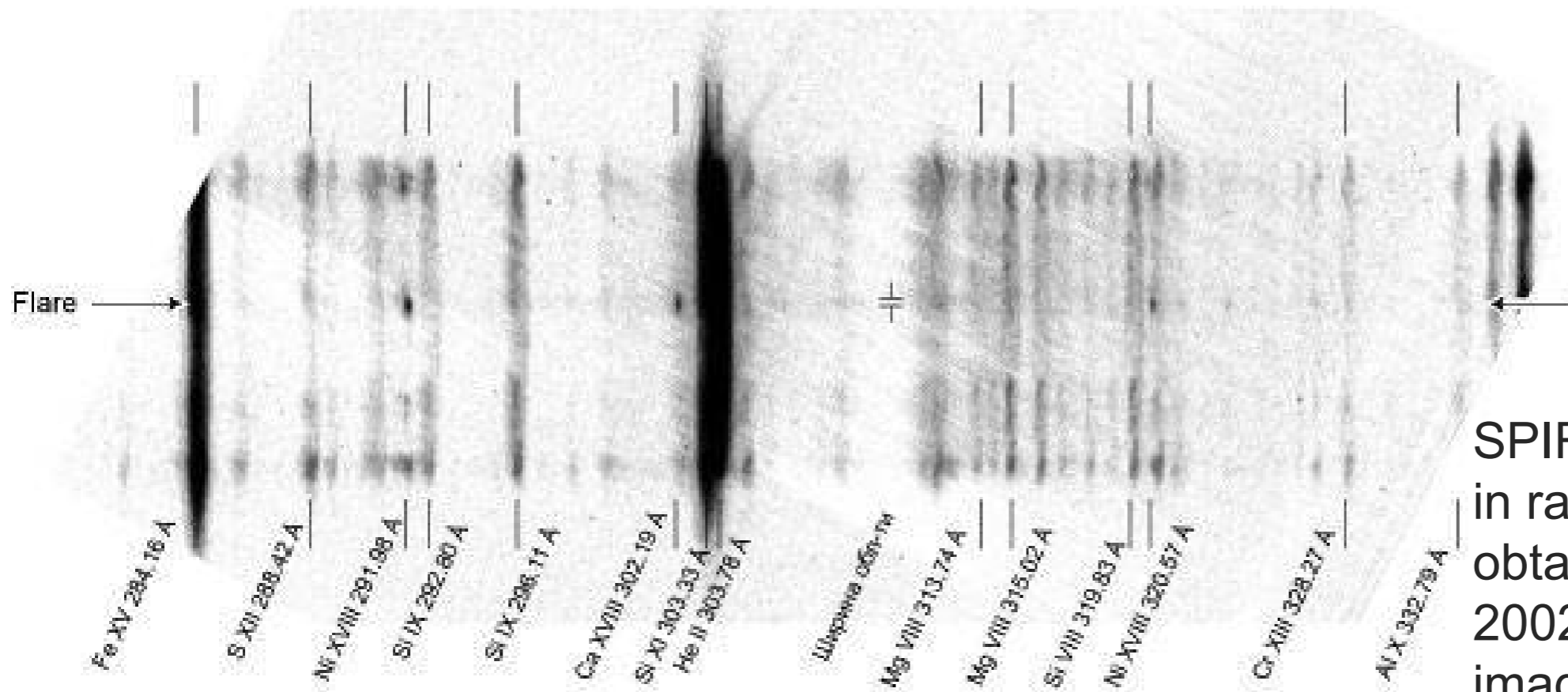
# DEM - results



# SPIRIT

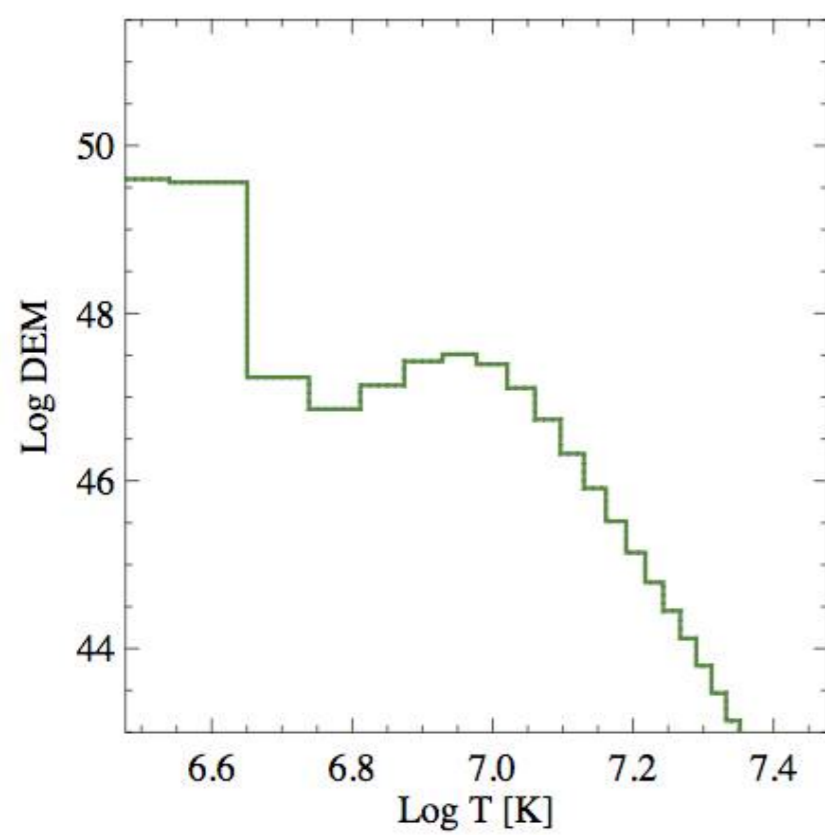


Simultaneous observations RESIK and SPIRIT were made on the flare decay starting at 4:24 until 4:26. This period was marked by an arrow on GOES light curve.

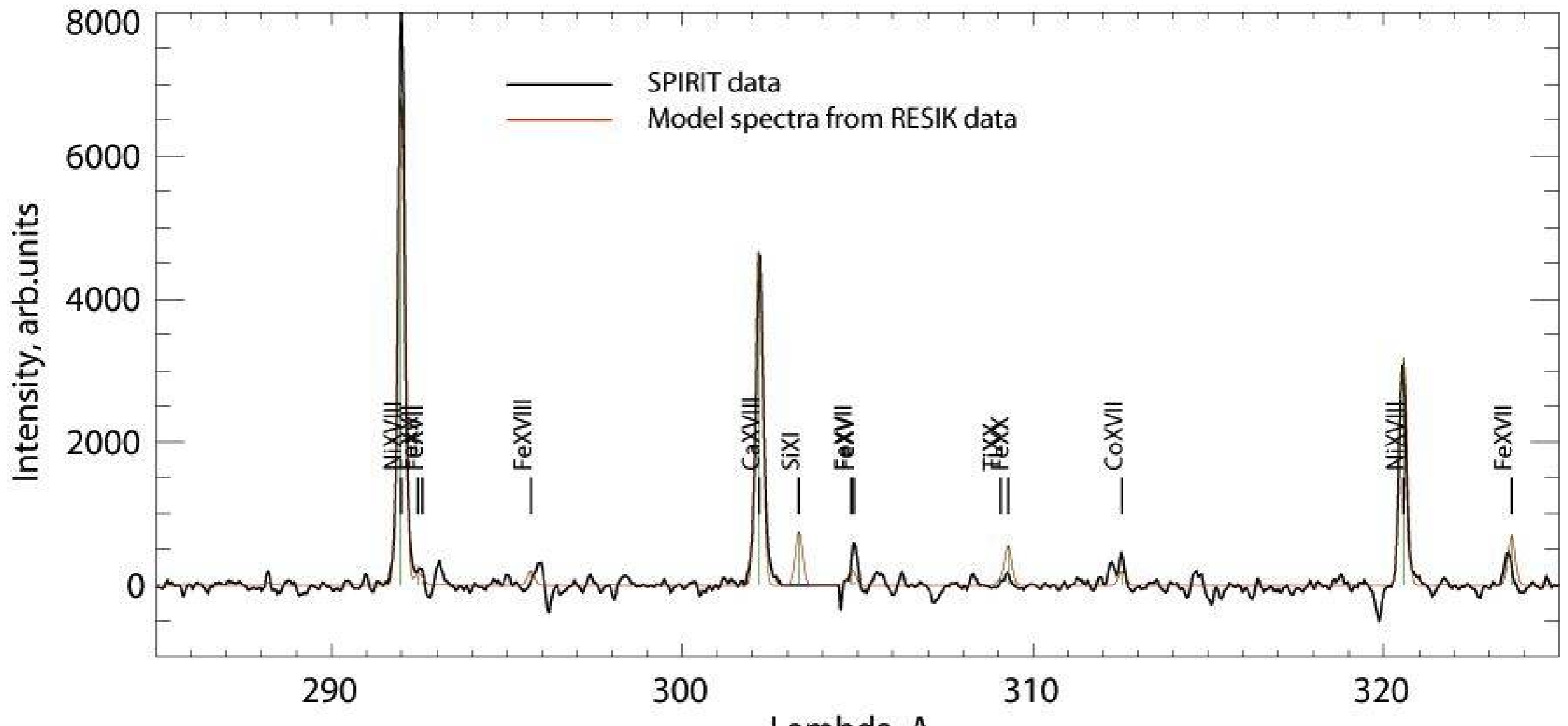


SPIRIT spektroheliogram in range 280-330 Å, obtained for 15 April 2002. Individual spectral images are marked.

DEM distribution obtained based on RESIK spectra for time period observed by SPIRIT



Comparison synthetic and observed SPIRIT spectrum



# Plans on the future

The simultaneous use of data from SPIRIT and RESIK instruments will allow to determine differential emission measure distributions in wide temperature range from 0.8 MK to 30 MK. The first obtained results are promising.

Work is in progress.

Thank you for your attention