SphinX spectral synthesis for low activity corona

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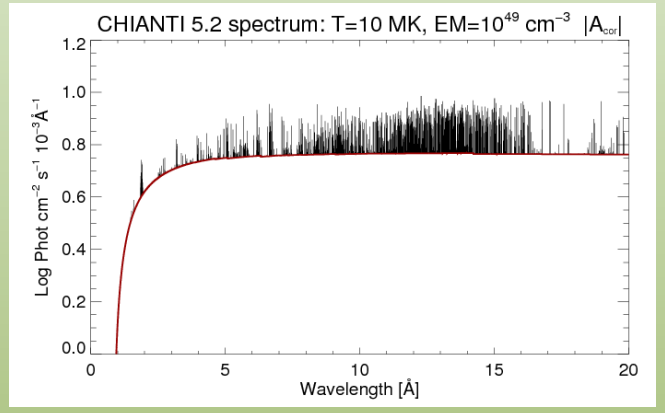
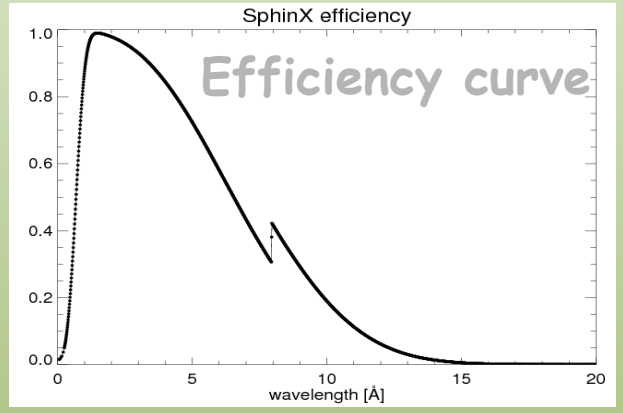
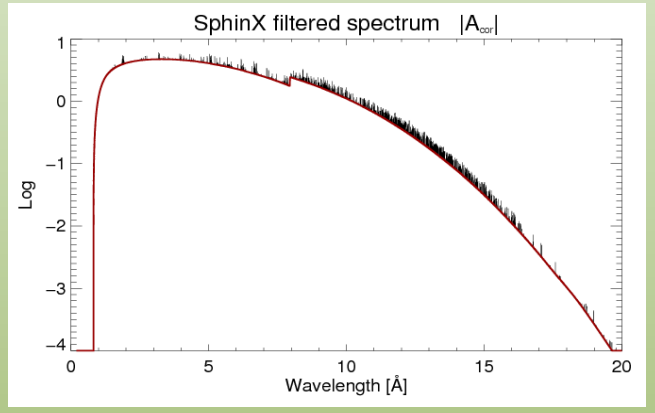
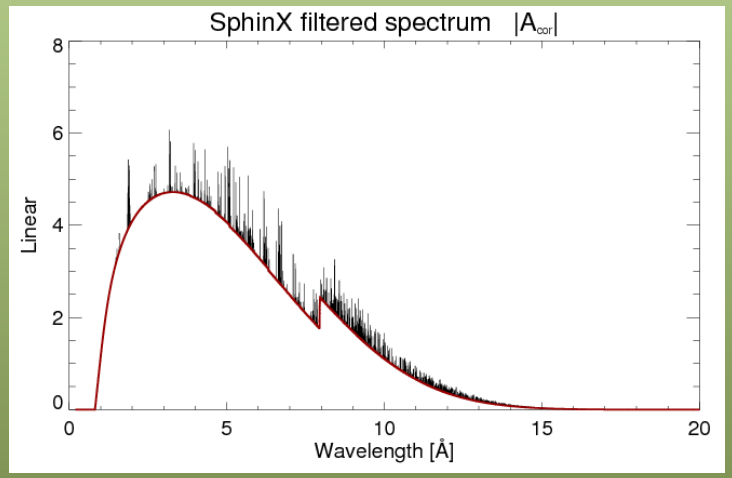
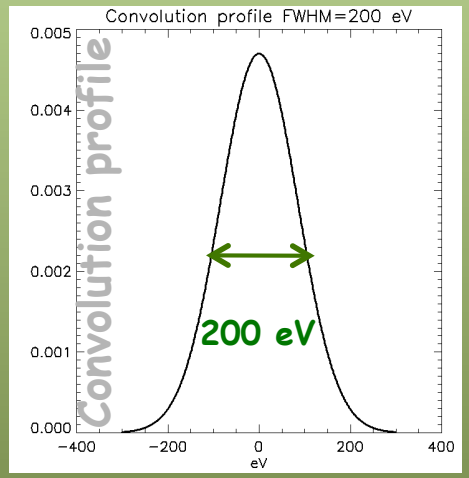
T=10 MK, EM=10^{49}cm^{-3}

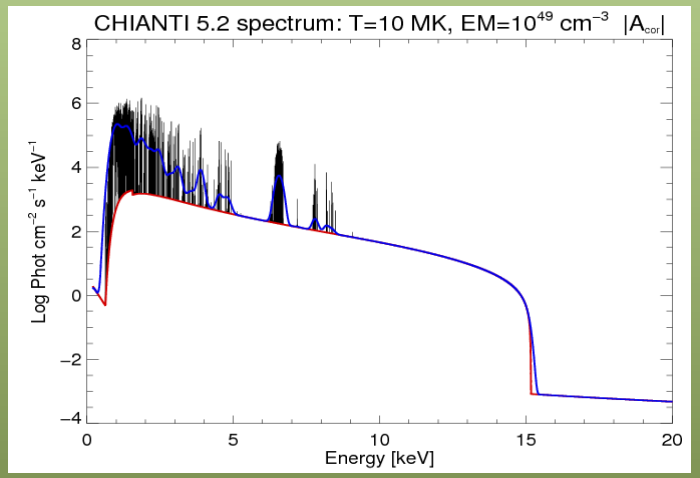




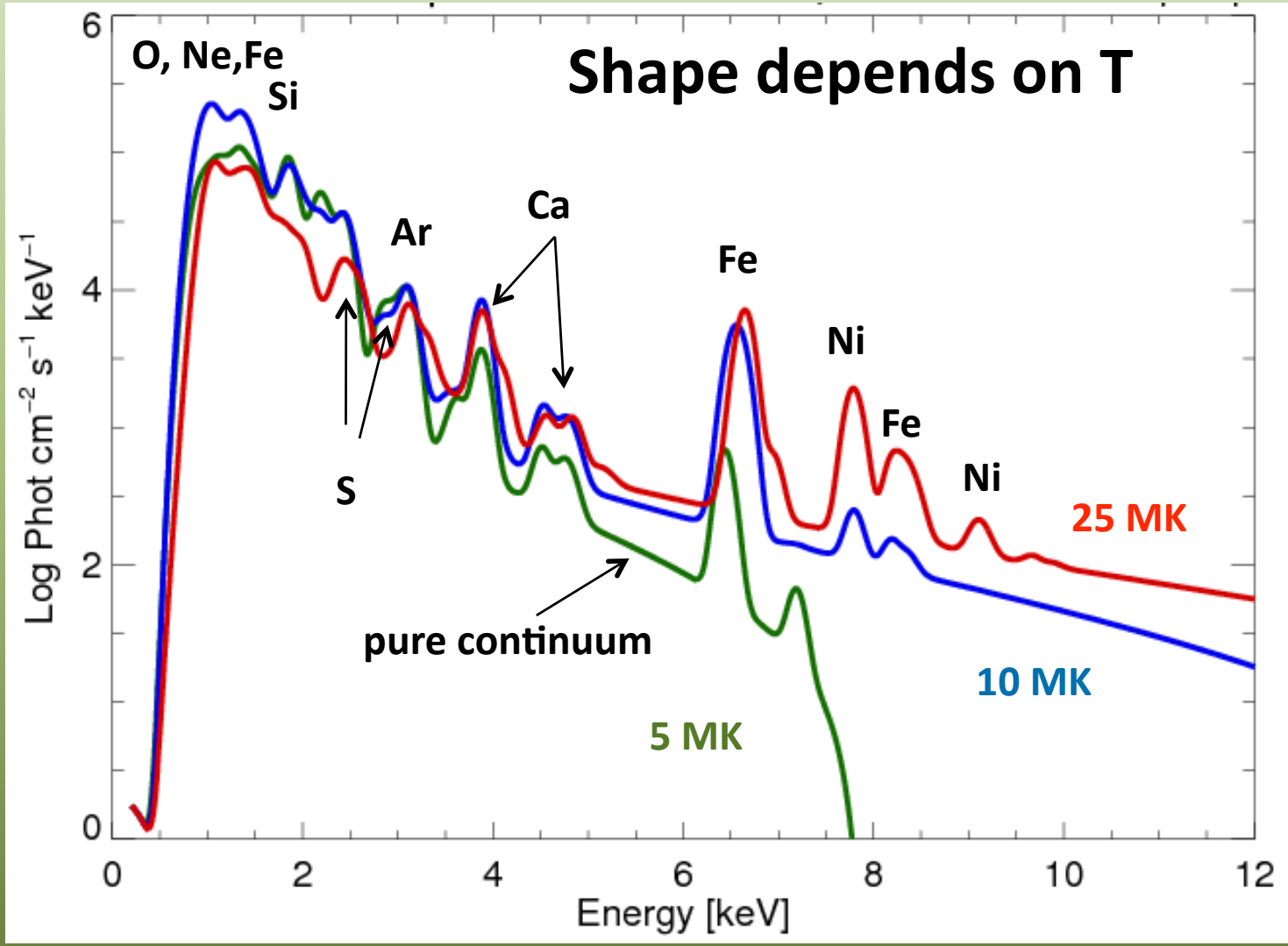








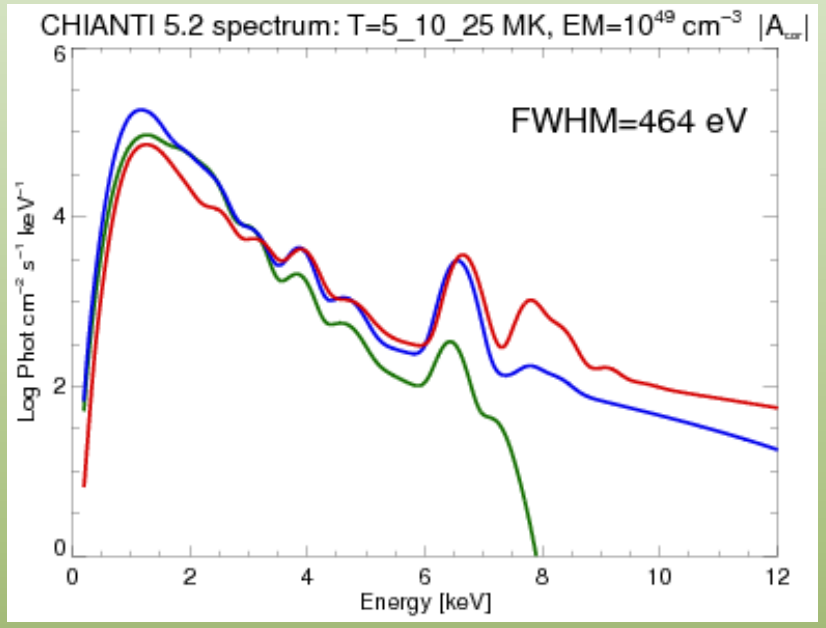
Temperature dependence: FWHM=200 eV

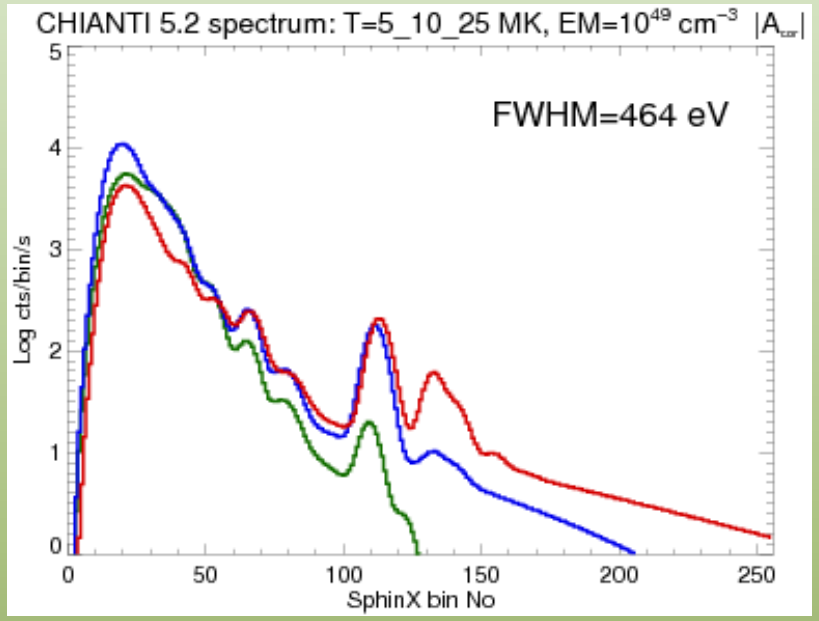


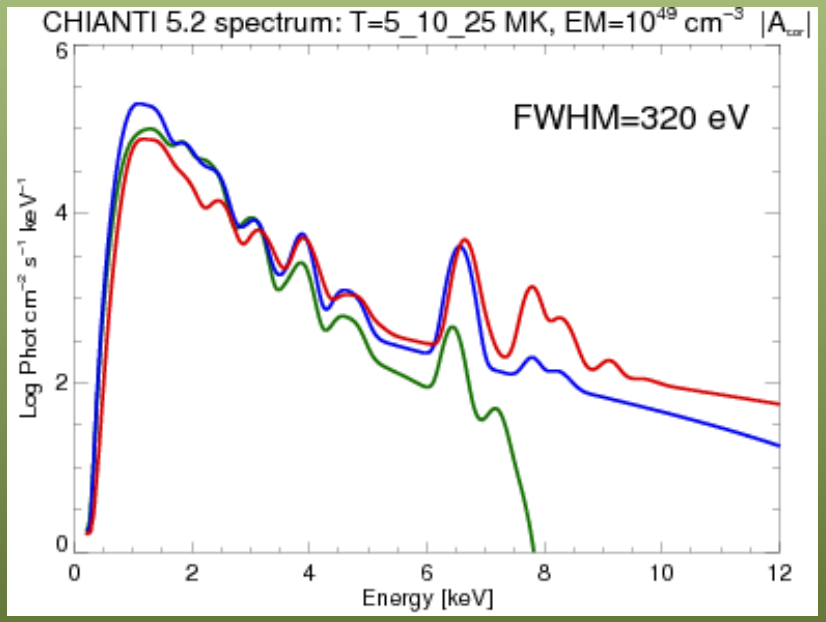
FWHM dependence; 464 keV & 320 keV (1cm^2)

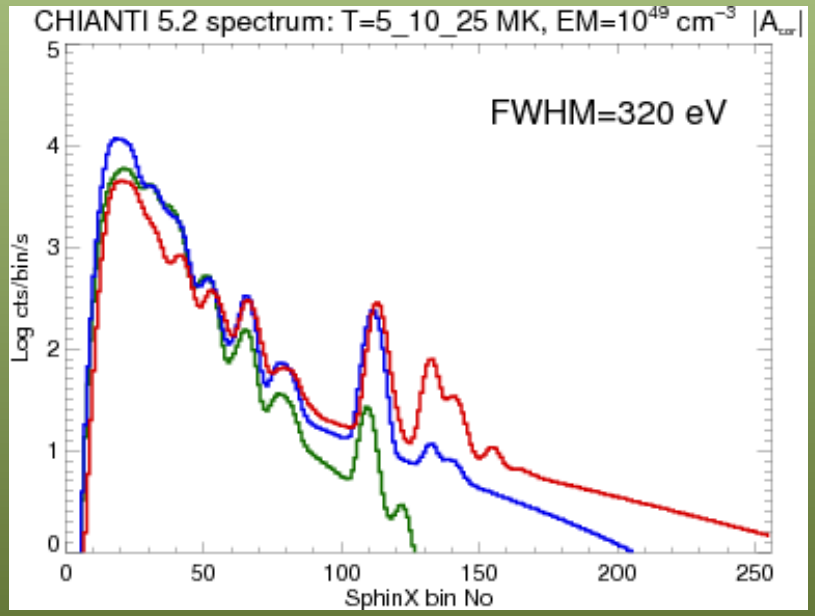
ENERGY

BINS

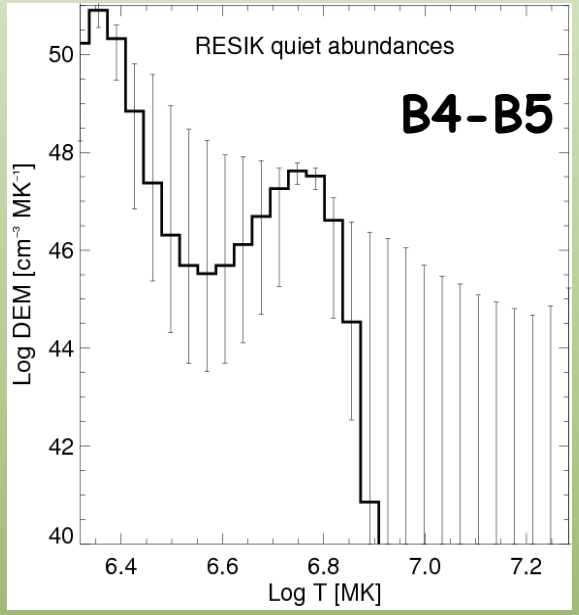


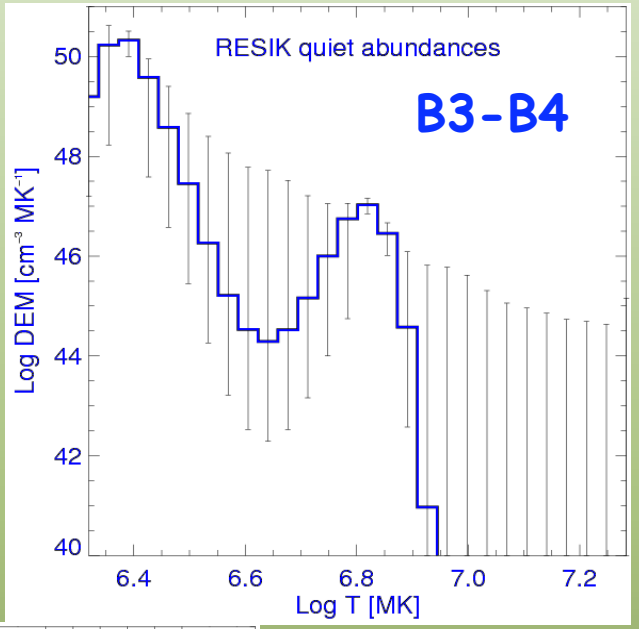


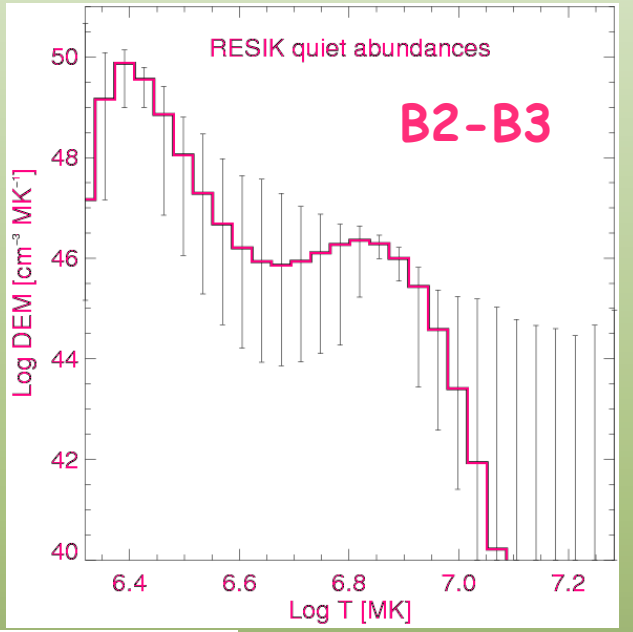


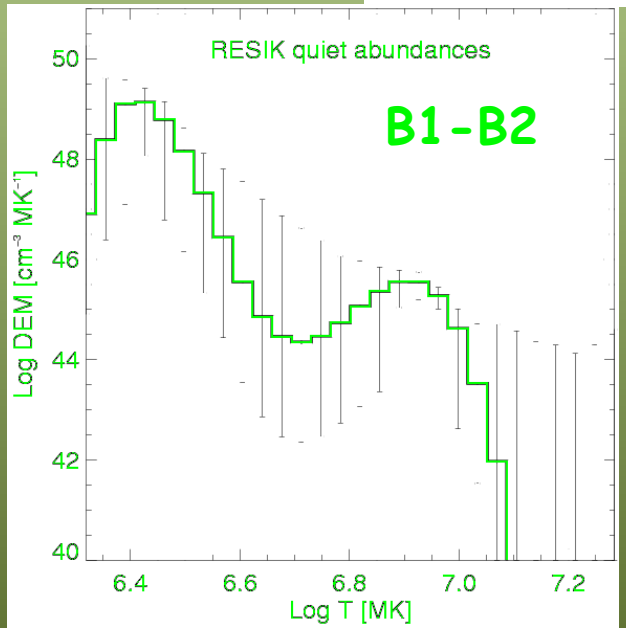


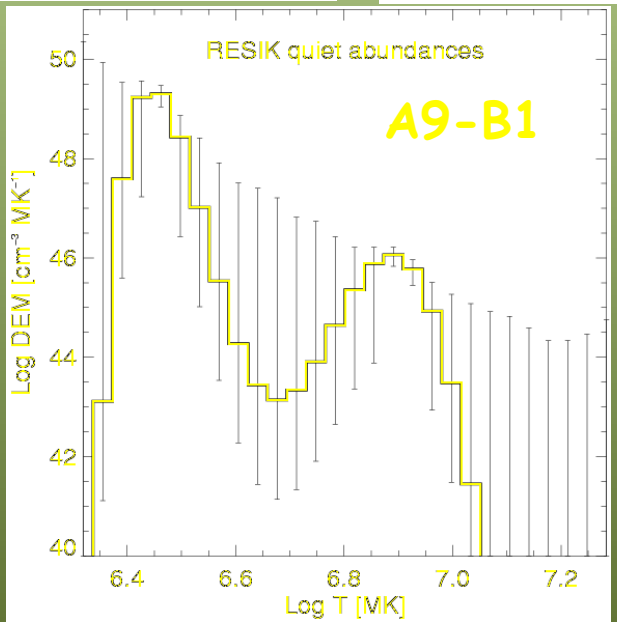
DEM for different activity levels (based on RESIK)



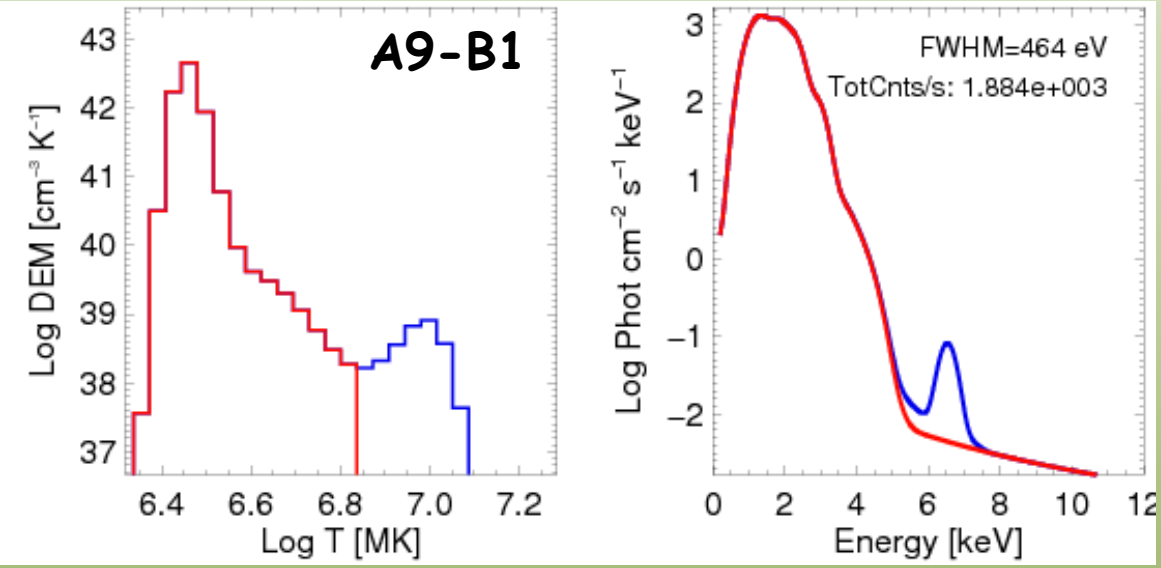


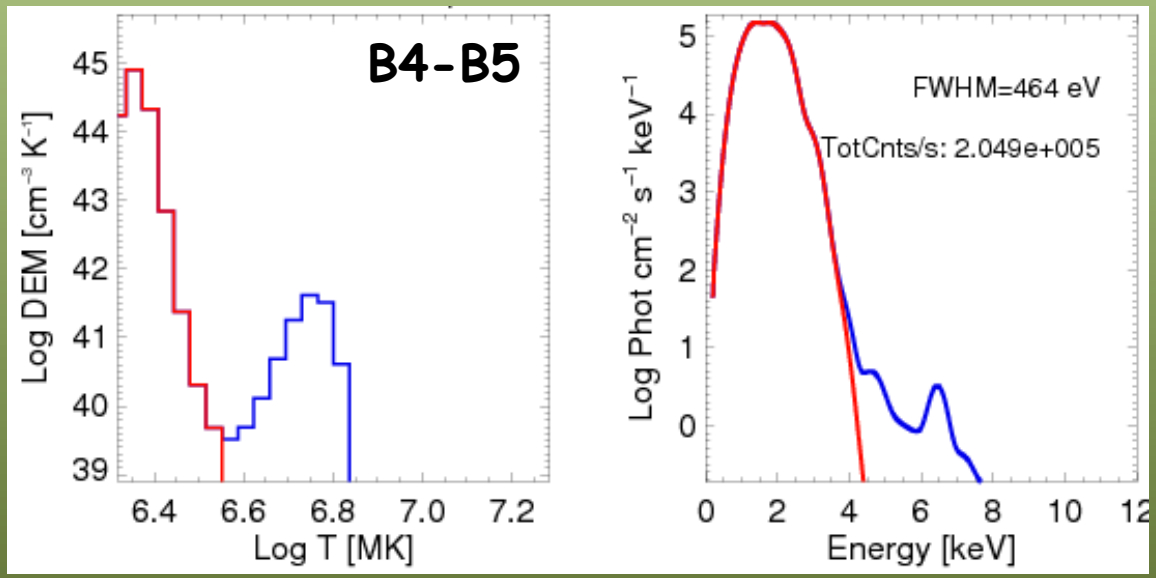






Dependence on level of activity; coronal abu.





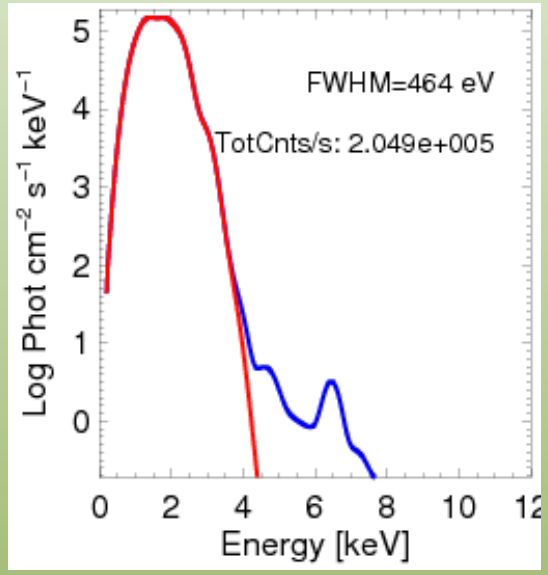
D1

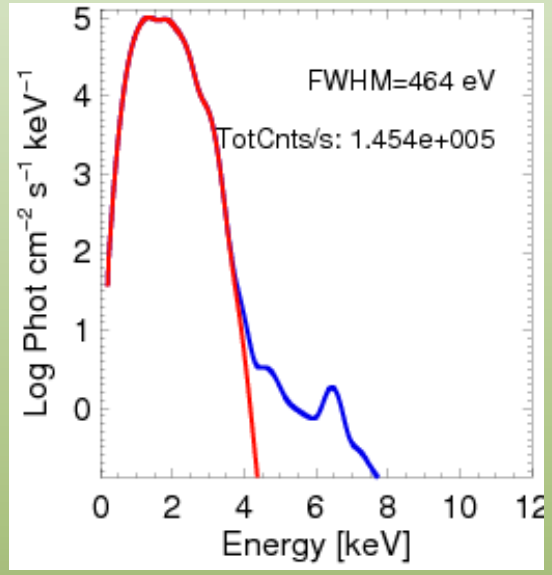
Coronal abu.

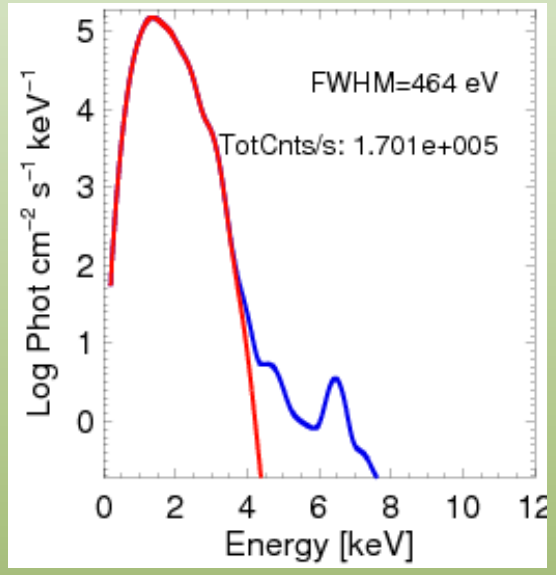
Phot. abu.

Quiet abu.

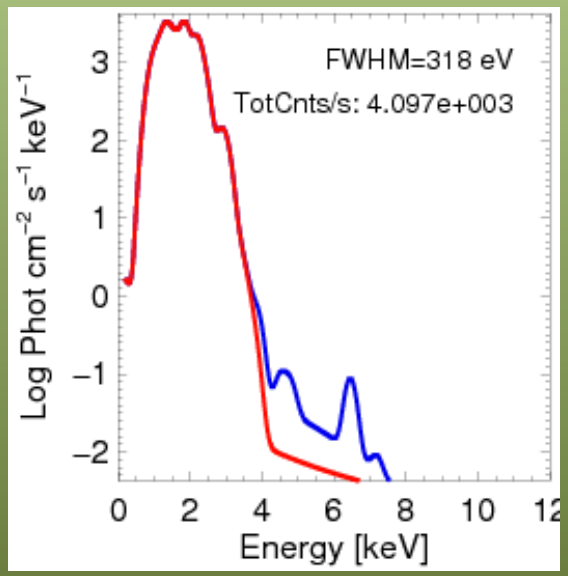
D1

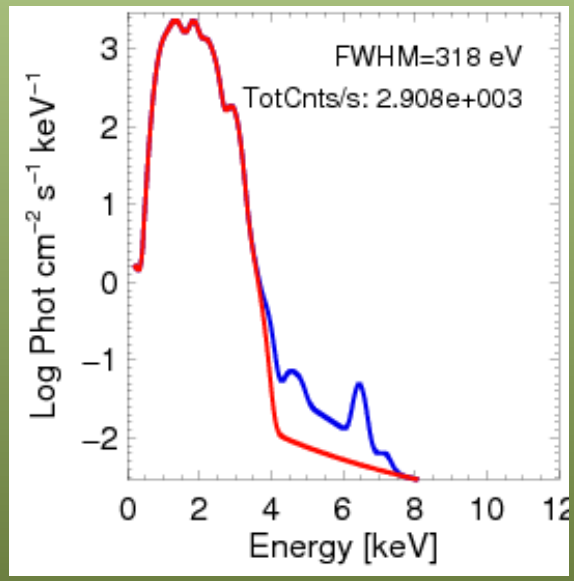


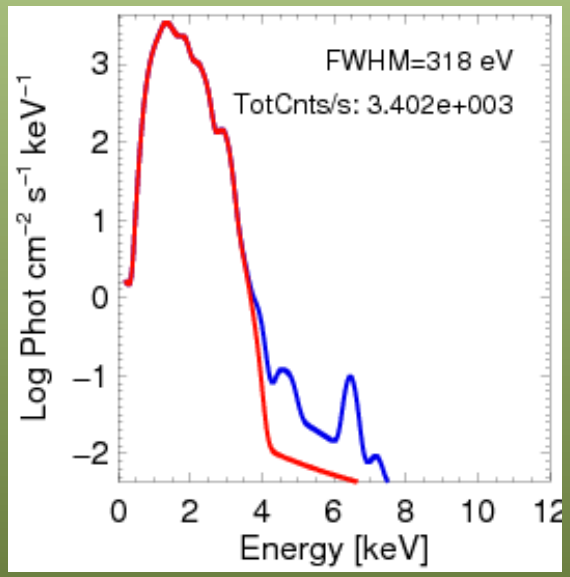




D2

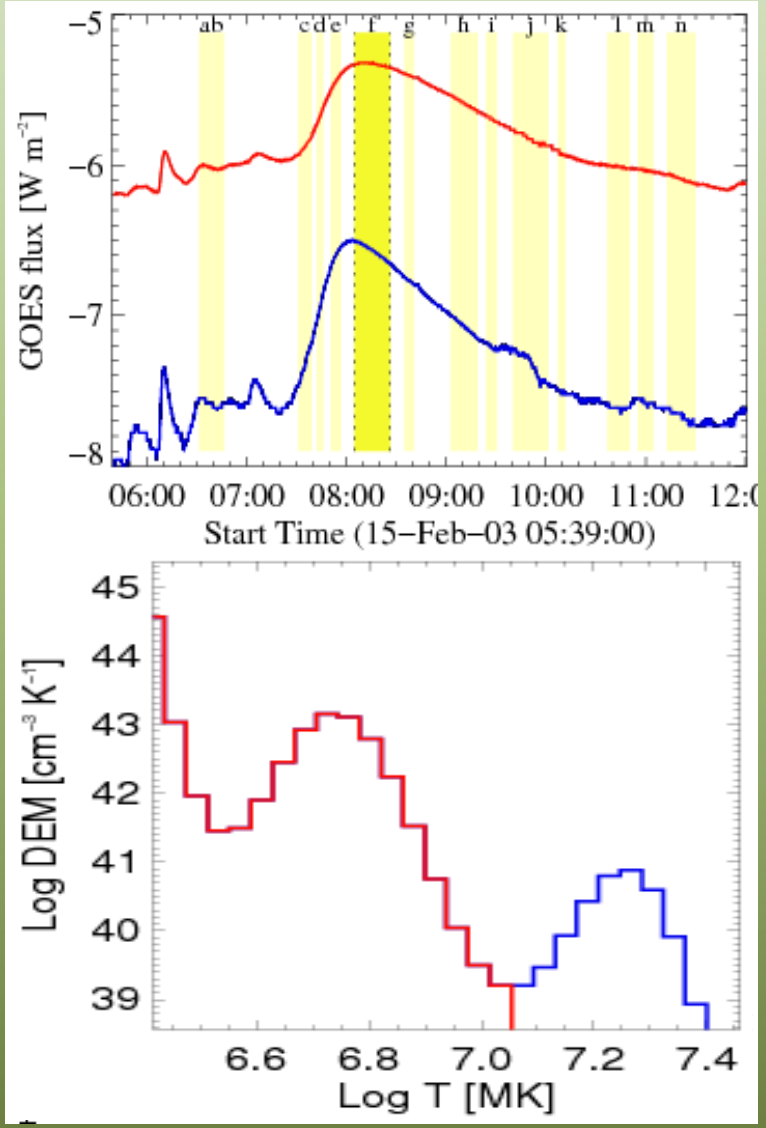


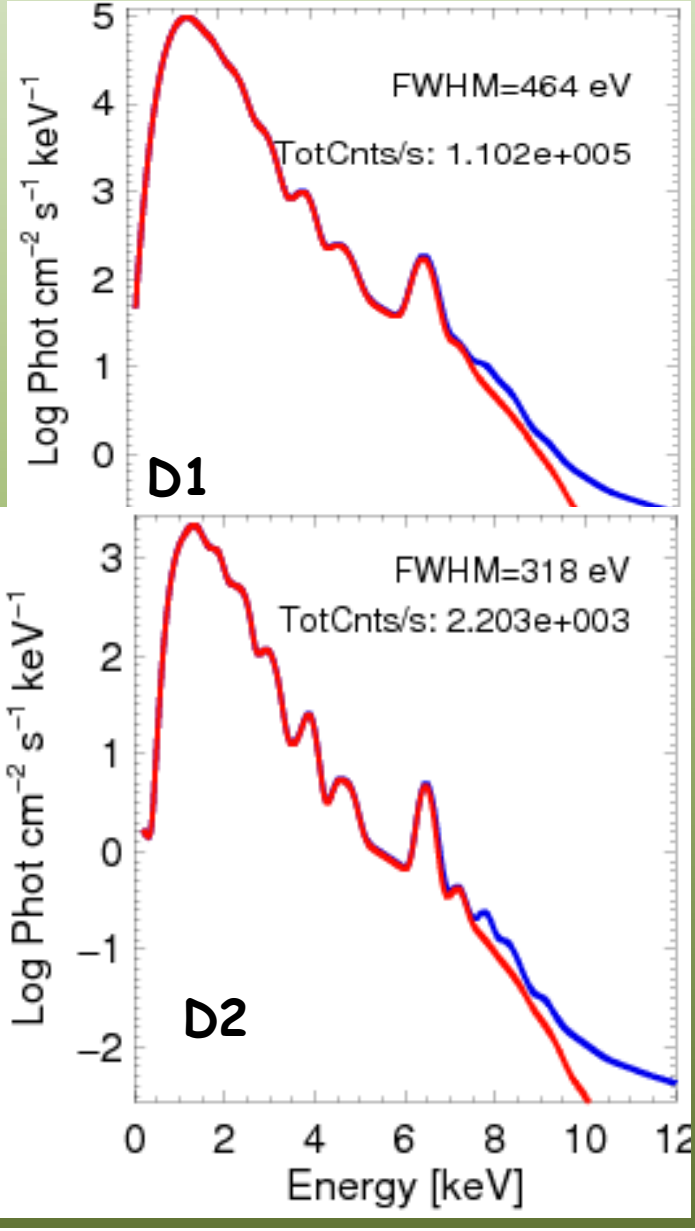




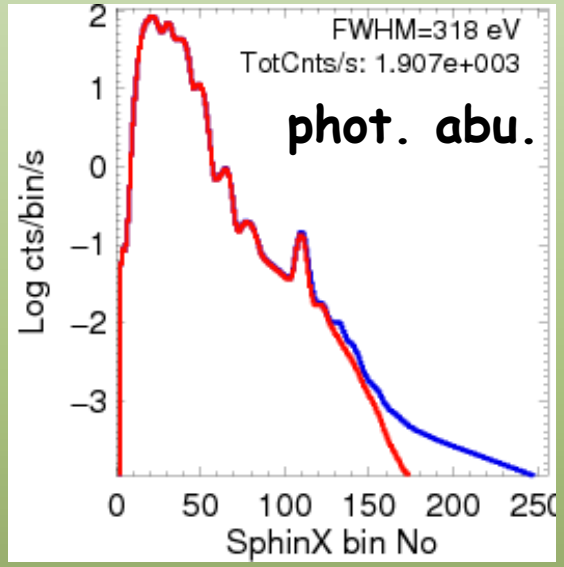
B4-B5

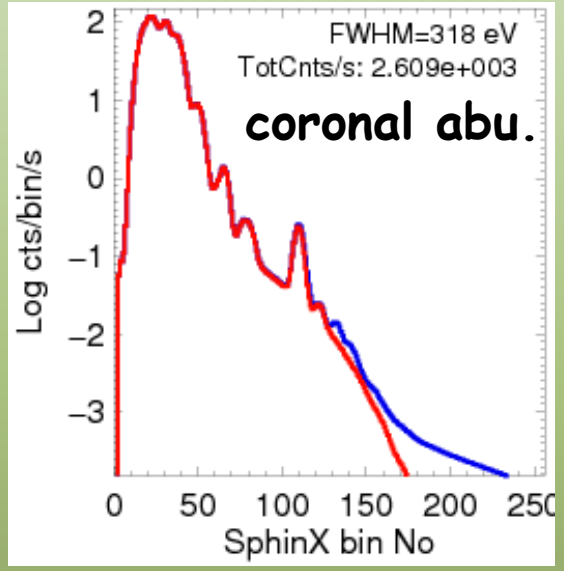
15 Feb. 2003 flare (C 4.5); flare abu.

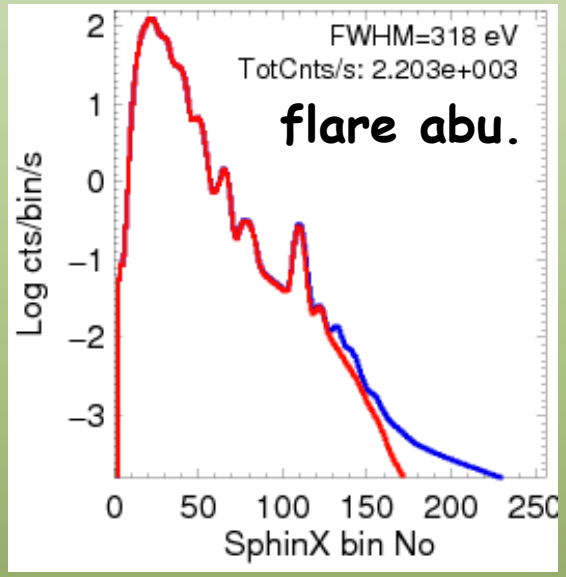




Flare case; Abundance dependence; D2







Conclusions

- For present activity level, expected count rates are ~100 cts/s in D1
- For activity levels ~B2-3, D1 becomes saturated→ switch to D2
- It will be possible to distinguish changes of abundance pattern with activity level
- For flares >B5, 0.01s spectra sampling time appears possible

☺ Thank You ☺

What we observe?

$$F_{obs} = F_{calc}(T, \lambda) \times \varphi(T) \times \eta(\lambda)$$