



# Update on interplanetary electron events observed with remote-sensing and in-situ instruments on Solar Orbiter

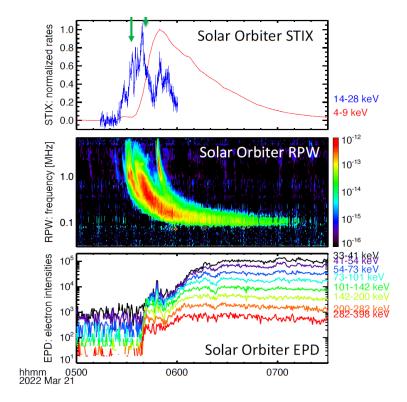
Alexander Warmuth

on behalf of the joint STIX-EPD-RPW-EUI working group

# Solar electron events (SEEs)

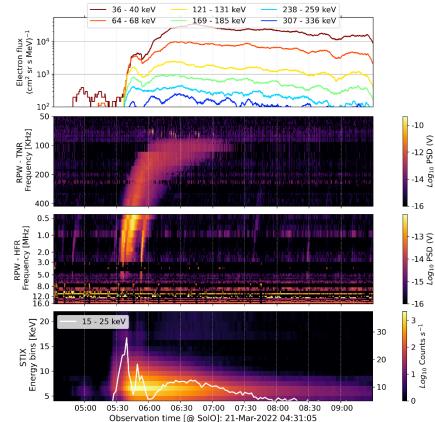
- remote-sensing (HXR, radio) and in-situ observations of energetic electrons
- strong evidence for acceleration of impulsive electron events in solar flares: correlations in timing and spectra, associated type III bursts, composition
- if same accelerator: gain insight into particle acceleration and propagation by comparing remote-sensing and in-situ observations
- however: some discrepancies between





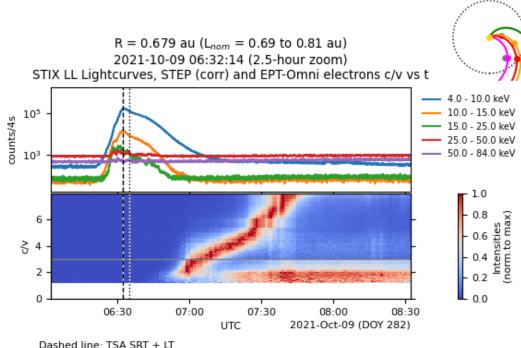
# Solar Orbiter: all essential assets on a single platform, sampling inner heliosphere

- in-situ particles: EPD
  - timing (including injection times), fluxes, spectra, composition, and anisotropy
- X-rays: STIX
  - timing, fluxes, spectra, and source location
- radio: RPW
  - presence and timing of type III bursts
- EUV: EUI
  - timing, position, morphology of EUV flare



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#### Associating in-situ electron events with solar flares: timing



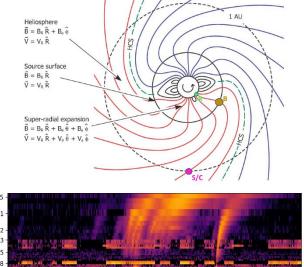
Dotted line: VDA SRT + LT

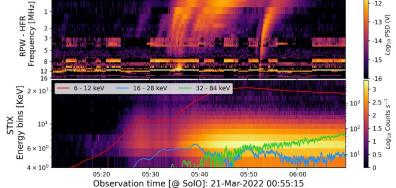
- infer injection time from EPD:
  - Time-shift analysis (TSA: needs to assume path length
  - Velocity disperion analysis (VDA; yields path length)
- compare with STIX HXR lightcurves

Associating in-situ electron events with solar flares: flare location and type III bursts

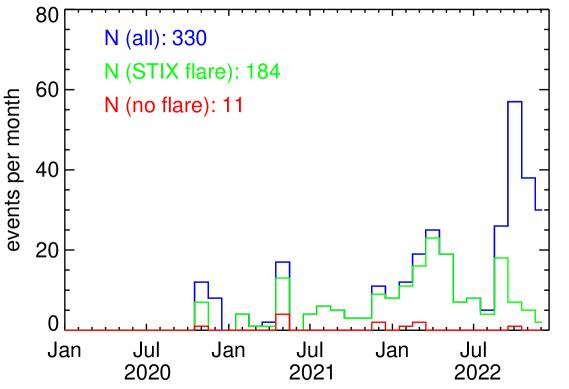
- evaluate STIX flare location in context of magnetic connectivity:
  - comparison to footpoints of magnetic field lines connecting to SolO using IRAP magnetic connectivity tool

- compare with type III burst timing:
  - get type III burst timing from RPW: how does it compare to STIX peaks and injection times?





#### Event occurence

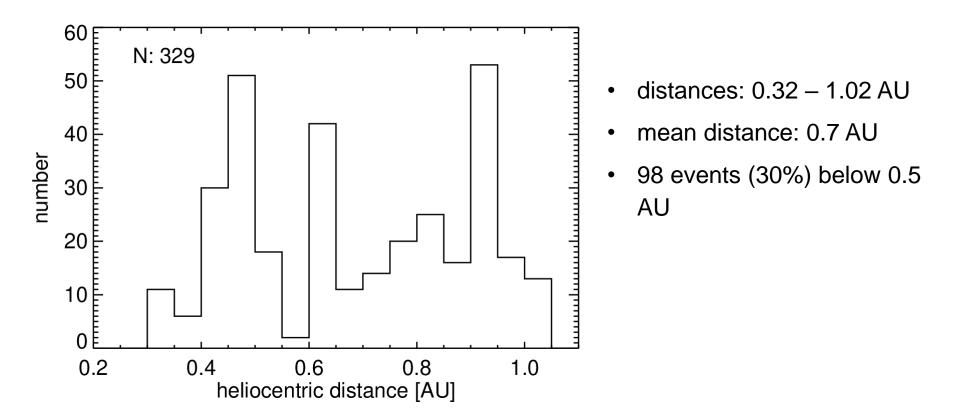


Nov 2020 – Dec 2022:

• EPD:

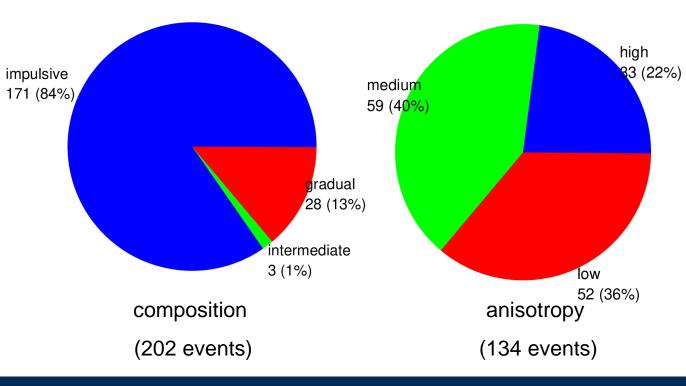
- 330 event entries
- 307 with TSA injection
- 83 with VDA injection
- 184+ EPD events with potentially associated STIX flare
- 11+ STIX non-

#### Heliocentric distances



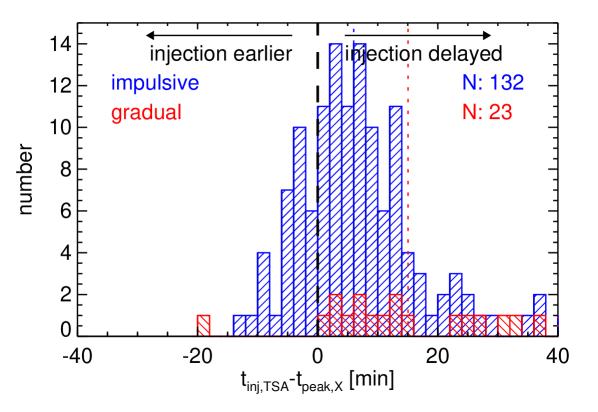
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#### Composition and anisotropy



- majority of events has impulsive composition
- ~2/3 of events have medium or high anisotropy

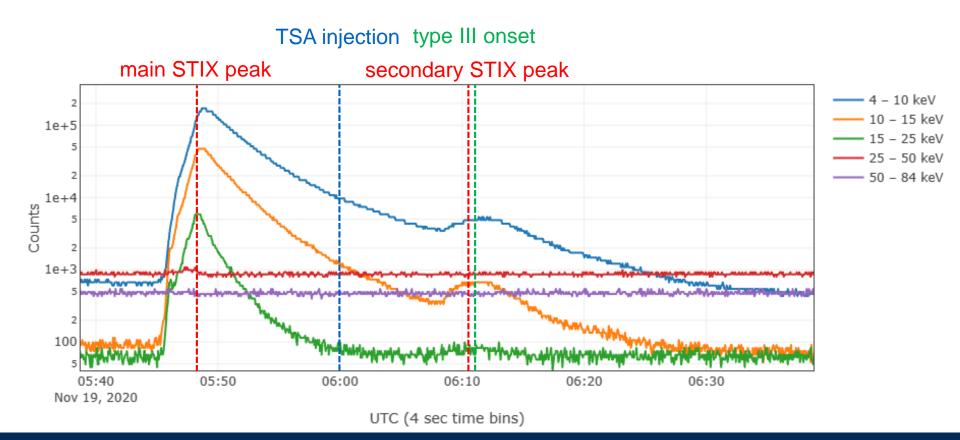
#### Time difference TSA injection – main nonthermal STIX peak



- median delays:
  - impulsive: 6 min
  - gradual: 15 min

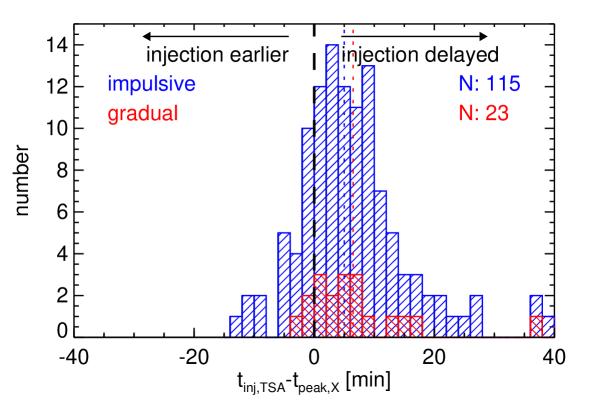
→ significantly better temporal association for impulsive events

#### Main and secondary nonthermal STIX peaks



Leibniz-Institut für Astrophysik Potsdam (AIP)

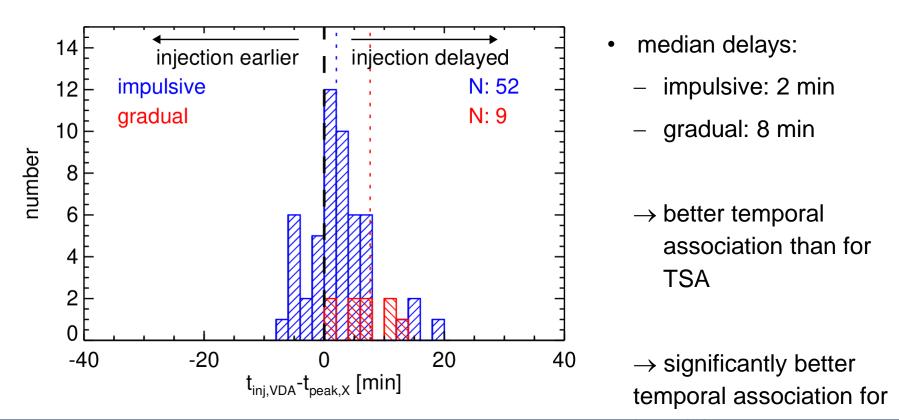
#### Time difference TSA injection – STIX peak closest to injection



- secondary STIX peaks used in ~50% of event
- median delays:
  - impulsive: 5 min
  - gradual: 7 min

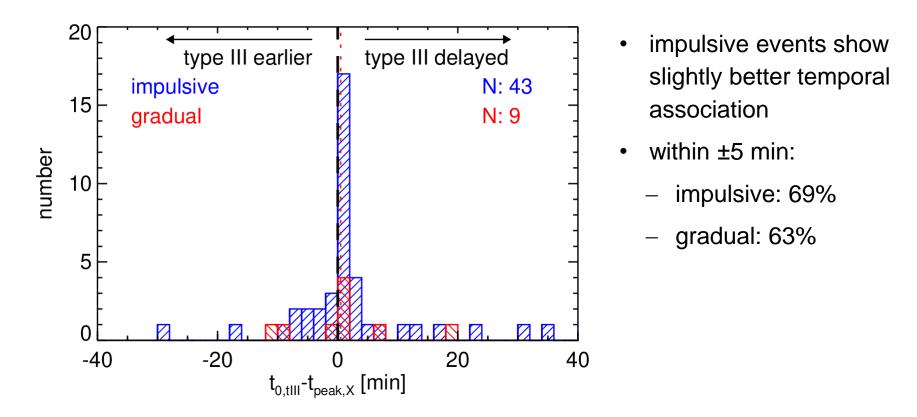
→ similar correspondence for impulsive and gradual events (due to minor STIX peaks in decay phase of

#### Time difference VDA injection – STIX peak closest to injection

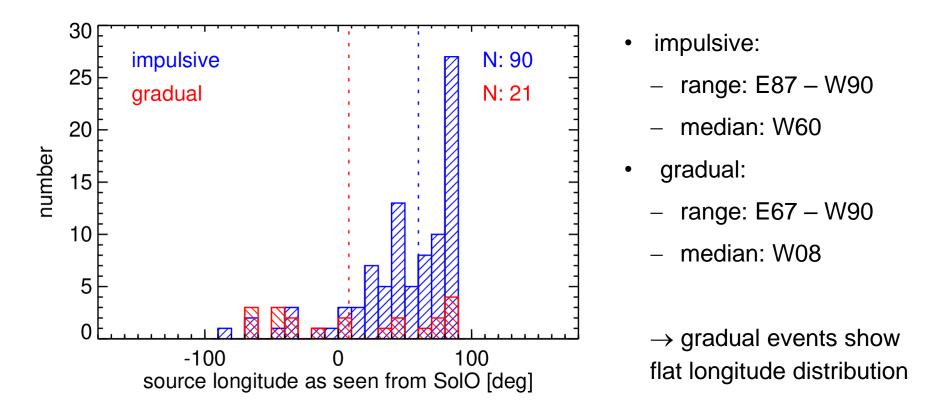


mpulsive vents Leibniz-Institut für Astrophysik Potsdam (AIP)

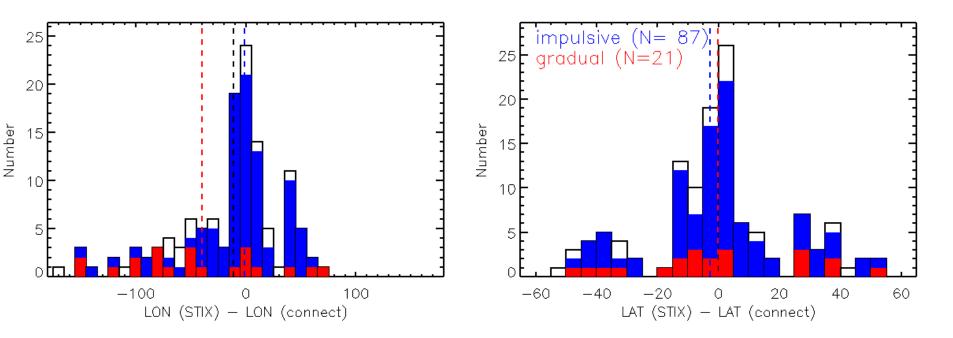
#### Time difference type III onset – nonthermal STIX peak



# Flare longitude as seen from Solar Orbiter

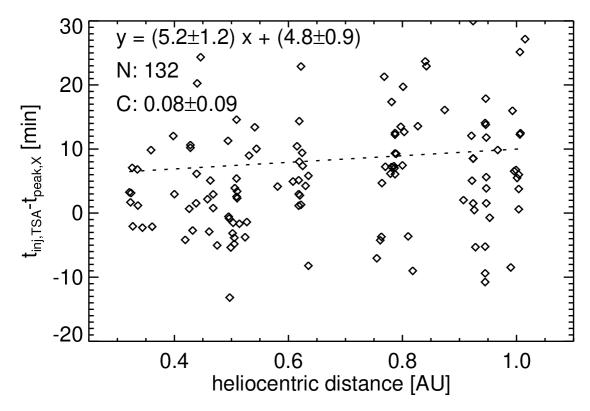


#### STIX flare location compared to footpoint of connecting field line



 $\rightarrow$  impulsive events show much better correspondence

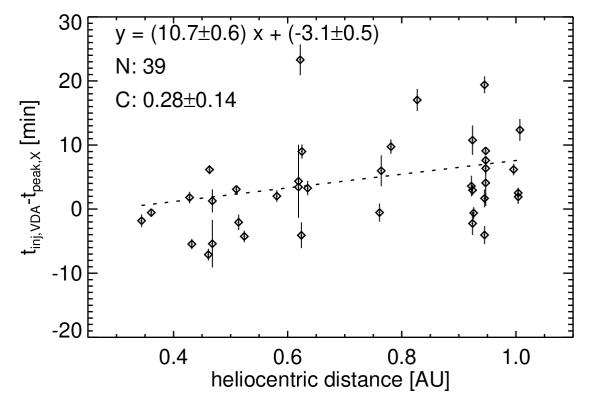
# Time difference TSA injection – nonthermal STIX peak vs. distance: impulsive events



 no correlation with distance

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Time difference VDA injection – nonthermal STIX peak vs. distance: impulsive events with high & medium anisotropy



 weak trend of time delay increasing with distance

# Outlook



- complete event list until end of 2022
- establish associations more firmly based on type III bursts and connectivity
- add CME information (will be provided by Metis team)
- publish first statistical results and joint event list
- perform spectral analysis with STIX and EPD, study relation as function of distance
- include in-situ electron observations from other locations in the heliosphere