## SphinX user guide (version 1.1, January 2010) Szymon Gburek

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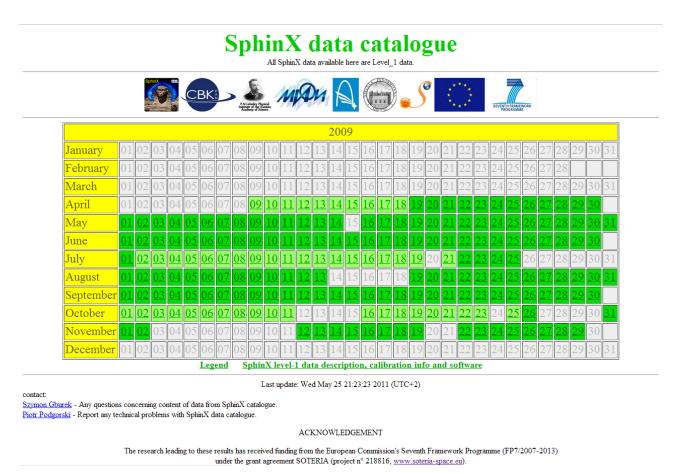
## CONTENT

1 SphinX Level-1 Data repository	1
2 SphinX Level-1 Data format	3
2.1 SphinX Event Lists	3
2.1.1 Event Extension	4
2.1.2 Exposure Extension	4
2.1.3 Good Time Interval Extension	5
2.2 Time System in SphinX FITS	5
2.3 Detector Response Matrix and Energy Boundary Information	5
2.3.1 Matrix Extension	6
2.3.2 Ebounds Extension	6
3 Application of SphinX Calibration Information in Spectral Analysis	6
4 SphinX Data Processing	7
4.1 Tools for SphinX data Processing	7
4.2 SphinX level-1 Data Processing in IDL	8
APPENDIX 1 - SphinX FITS headers format 1	0

## 1 SphinX Level-1 Data repository

SphinX level-1 data are event lists stored in Flexible Image Transport System (FITS) files. Well documented and standardized Office of Guest Investigator Programs (OGIP) FITS format was used for preparation of SphinX level-1 FITS files. A description of the OGIP FITS format for event lists can be found for example on the WWW site

http://heasarc.gsfc.nasa.gov/docs/heasarc/ofwg/docs/events/ogip\_94\_003/ogip\_94\_003.html



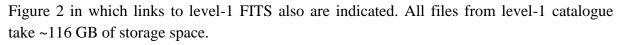
http://156.17.94.1/sphinx\_l1\_catalogue/SphinX\_cat\_main.html

Figure 1. SphinX level-1 main catalogue page.

This page (shown in Figure 1) is organized in a form of calendar in which day fields are link to SphinX daily summary pages. Days for which there are SphinX data available are indicated in green. In light green are days in which the satellite was not eclipsed by the Earth on its orbit. Below the calendar in the page there are links to Legend document with description of the catalogue sites. The second link *SphinX level-1 data description, calibration info and software* below the calendar on the right hand side opens a WWW site which contains

- guide for SphinX users,
- description of SphinX level-1 data and their format,
- IDL routines for processing SphinX level-1 data,
- example IDL programs.
- example level-1 FITS files and the calibration FITS with SphinX DRM
- description of SphinX flag system

Each day link opens a daily summary page with visualization of SphinX data and links for downloading level-1 data files. Top portion of a selected daily summary page is shown in



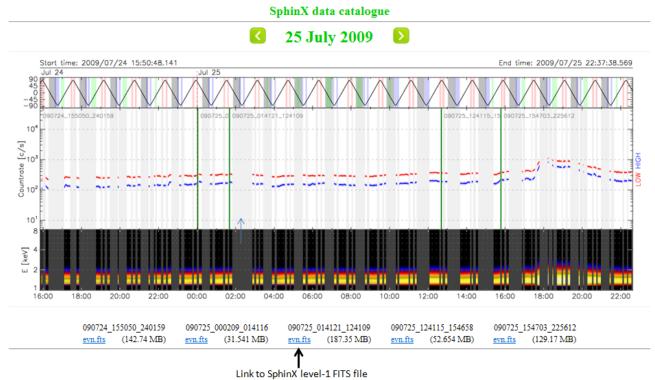


Figure 2. Top portion of a selected daily page from SphinX level-1 catalogue.

## 2 SphinX Level-1 Data format

## 2.1 SphinX Event Lists

Naming convention for SphinX level-1 event FITS files is

## SPHINX\_YYMMDD\_HHMMSS\_HHMMSS1\_evn\_D1\_L1.fts,

where *YYMMDD* is the date on which measurements in the file were recorded, *HHMMSS* is the start time and *HHMMSS1* the end time of measurements in the file. If measurements extend beyond midnight to the next day then *HH* in *HHMMSS1* is greater than 24. For example in the file *SPHINX\_090718\_230838\_260933\_evn\_D1\_L1.fts* are data recorded by SphinX from 2009-07-18 23:08:38 Universal Time Coordinated (UTC) to 2009-07-19 02:09:33 UTC.

Each particular SphinX FITS event file contains

- primary header,
- EVENT extension,
- EXPOSURE extension,
- GTI standard good time interval extension intervals of noninterrupted instrument operation.

#### 2.1.1 Event Extension

The binary table of EVENT extension in SphinX FITS files consists of the following columns

TIME – time of event registration (in seconds from reference time)
PHA – channel in which an event was registered (0-255 for SphinX flight data)
PI – pulse invariant channel for an event
ENERGY – event energy
FLAG – an event origin and type description
NRM – SphinX specific column – gives a link to level-0 data (not important for users)

The reference time for conversion of TIME column to UTC is defined in the EVENT extension header. The conversion procedure to UTC is described in Time system in SphinX FITS section of this document below.

Not only events caused by hits of solar X-ray photons into detector crystals are tabulated in the SphinX FITS. The events coming from energetic particle hits, events recorded during satellite nights or electronic glitches are present in the SphinX event FITS files too. Therefore a FLAG is provided for each event in order to identify event origin. For good solar events the flag value is set to zero (FLAG = 0).

#### **2.1.2 Exposure Extension**

The binary table of EXPOSURE extension in SphinX FITS files consists of the following columns

TSTART TSTOP FRACEXP

SphinX registered events contiguously in a short time intervals. Between these intervals there are data gaps caused for instance by electronics resets, onboard processing or interruptions due to sending data to the telemetry buffer.

The columns TSTART and TSTOP of the EXPOSURE binary table give times between which, events were registered without interruptions. The FRACEXP column of the EXPOSURE binary table contains the correction factor which is necessary to calculate exposures. Actual exposures for events recorded between TSTART and TSTOP are defined as follows

## EXPOSURES = (TSTOP – TSTART)\*FRACEXP

TSTART and TSTOP times are in seconds from a reference time given in the EXPOSURE extension header (it is the same time as the reference time in EVENT extension header).

The reference time for conversion of TSTART and TSTOP columns to UTC is defined in the EXPOSURE extension header. The conversion procedure to UTC is described in Time system in SphinX fits section of this document below.

## 2.1.3 Good Time Interval Extension

The third extension in SphinX event FITS is Good Time interval extension – GTI extension. Its binary table has only two columns

TSTART TSTOP

TSTART and TSTOP keyword values indicate times between which the instrument registered events continuously. If no additional event filtration was performed by user TSTART and TSTOP are exactly the same as in EXPOSURE extension.

TSTART and TSTOP are in seconds from a reference time given in the GTI extension header (it is the same time as the reference time in EVENT and EXPOSURE extension header).

## 2.2 Time System in SphinX FITS

Reference time for any time columns in SphinX event FITS is start of a day on which the measurement in the file were recorded. It is specified in the TIMEZERO and MJDREF keywords in EVENT, EXPOSURE and GTI extension headers. TIMEZERO gives the number of days that elapsed from 1979-01-01T00:00:00.000 till the reference time and MJDREF keyword value gives a modified Julian date of 1979-01-01T00:00:00.000 (MJD = JD - 2400000,5). Thus modified Julian date TIME\_REF\_MJD of the reference time is

## TIME\_REF\_MJD = MJDREF + TIMEZERO

## 2.3 Detector Response Matrix and Energy Boundary Information

Data from X-ray tests of SphinX in Palermo XACT facility (2007) and Bessy II synchrotron in Berlin (2008) were finally processed in 2010 in order to prepare all necessary calibration information for SphinX data analysis. The calibration information is stored in a single OGIP FITS file – SphinX response FITS file. This FITS is named SPHINX\_RSP\_256\_nom\_D1.fts. A description of the SphinX response FITS file format can be found on the site

ftp://legacy.gsfc.nasa.gov/caldb/docs/memos/cal\_gen\_92\_002/cal\_gen\_92\_002.pdf

The SphinX response FITS file contains:

- primary header,
- MATRIX extension,
- EBOUNDS extension.

#### **2.3.1 Matrix Extension**

The binary table of MATRIX extension in SphinX response FITS file contains the Detector Response Matrix DRM and consists of the following columns

ENERG_LO	- the lower energy bound of the energy bin,	
ENERG_HI	- the upper energy bound of the energy bin,	
N_GRP	- contains the number of 'channel subsets' for the energy	
F_CHAN	– contains the channel number of the	
	start of each "channel subset" for the energy bin,	
N_CHAN	<ul> <li>– contains the number of channels</li> </ul>	
	within each "channel subset" for the energy bin,	
MATRIX	- containing all the response values for each 'channel	
	subset' for the energy bin.	

There is no grouping of channels in SphinX DRM thus N\_GRP, F\_CHAN, N\_CHAN are scalars of values 1, 1, 256 respectively. These columns are kept in the response FITS to be consistent with OGIP format requirements. Columns ENERG\_LO and ENERG\_HI are in units of keV.

#### **2.3.2 Ebounds Extension**

The second extension in SphinX response FITS is EBOUNDS extension with the following columns:

CHANNEL	– SphinX channel number
E_MIN	– nominal energy of lower boundary of the detector channel
E_MAX	– nominal energy of upper boundary of the detector channel

where columns E\_MIN and E\_MAX are in units of keV.

## **3 Application of SphinX Calibration Information in Spectral Analysis**

Nominal SphinX DRM, stored in the MATRIX extension of SphinX response FITS, is of the size of  $256 \times 281$ . SphinX DRM is used in spectral analysis according to the following equation

$$O = DRM \# (M * DE),$$

where *M* is a model spectrum vector in units photons/s/cm<sup>2</sup>/keV defined for energy bin centers that are given as (ENERG\_HI + ENERG\_LO)/2. Energy bin width vector on the model side of the equation  $DE = \text{ENERG}_{HI} - \text{ENERG}_{LO}$ . *O* is the spectrum model folded with the instrument response in units of counts/s as it would be observed in SphinX channels. The # sign stays here for matrix multiplication. The energy edges for *O* vector channels are defined by E\_MIN and E\_MAX in EBOUNDS extension.

Users may work with different energy binnings from these defined in nominal SphinX response FITS by columns pairs (ENERG\_HI, ENERG\_LO) and (E\_MIN, E\_MAX) for both sides of the equation above. If it is the case, users also have to rebin DRM accordingly to their needs and energy edges specification. This can be done automatically in most of spectral analysis packages.

The vector O is next compared to real SphinX measurement D in order to find out if the model M fits to data. Normalized for degree of freedom chi-square distribution  $\chi_{norm}^2$  can be used to test goodness of fit

$$\chi_{\rm norm}^2 = \frac{1}{\nu} \sum \frac{(D-O)^2}{\sigma^2},$$

where v is the number of degrees of freedom given by N - n - 1, where N is the number of observations, and n is the number of the fitted model parameters. The errors  $\sigma$  can be calculated from the expression

$$\sigma^2 = D + I_{err}^2,$$

where the first term describes squared Poisson error and  $I_{err}$  in the second term is the instrumental error which is less than 1% of D.

## 4 SphinX Data Processing

#### 4.1 Tools for SphinX data Processing

Processing of SphinX event list consists of the following steps:

- 1) Reading data
- 2) Filtering out the selected for analysis event lists by energy bins and type. The type of an event can be determined by FLAG column. Description of SphinX flags is given in each event list extension header in SphinX FITS file and on the www site of SphinX level-1 data catalogue.
- 3) Preparing higher level data products (such as spectra lightcurves) from filtered event lists.
- 4) Performing data analysis.

Filtration and preparation of spectra/lightcurves can be done using, for example, XSELECT (FTOOLS) or EVTSELECT from XMM Newton Science Analysis System (SAS) environment.

Spectral analysis can be performed using spectral analysis packages such as XSPEC, OSPEX or by user software directly in IDL, C++ or other language. SphinX team uses OSPEX in IDL described on the site

http://hesperia.gsfc.nasa.gov/ssw/packages/spex/doc/ospex\_explanation.htm

The tools for processing and analysis of SphinX data are available for downloading and install from

## http://heasarc.gsfc.nasa.gov/docs/software.html http://xmm.esa.int/sas/

It is also possible to perform SphinX data analysis using web interface on which on-line access to the newest version of FTOOLS is provided

http://heasarc.gsfc.nasa.gov/hera/

## 4.2 SphinX level-1 Data Processing in IDL

SphinX data analysis can be performed in any programistic language. Preferred environment is IDL with installed SolarSoft package. The SolarSoft installation and downloads are available on

## http:/www.lmsal.com/solarsoft/

Recommended IDL routine for reading SphinX level-1 event list FITS and calibration file is *mrdfits* which is in SolarSoft package. To read SphinX event list FITS the following call of *mrdfits* can be used

## *data* = *mrdfits*(*filename*, *id*, *hdr*, *status*=*status*).

Here *filename* is the SphinX event list file name, *data* is an IDL structure with FITS binary extension table and *hdr* is FITS extension header. *Status* keyword is greater or equal to zero if reading was successful and *id* is FITS extension number. Header data units are ordered by *id* in SphinX event FITS always in the same way. When reading SphinX event FITS using the mrdfits routine one obtains the following IDL variables:

id = 0 - hdr = primary header,	data = 0,
id = 1 - hdr = events header, $data$ =	event list IDL structure,
id = 2 - hdr = exposure header,	<i>data</i> = exposure IDL structure,
id = 3 - hdr = GTI header,	data = GTI IDL structure.

For reading SphinX response FITS *filename* must point to SPHINX\_RSP\_256\_nom\_D1.fts file. In this case one gets:

id = 0 - hdr = primary header,	data = 0,
id = 1 - hdr = DRM header,	<i>data</i> = DRM MATRIX IDL structure,
id = 2 - hdr = EBOUNDS header,	<i>data</i> = EBOUNDS IDL structure.

The SphinX FITS files should be in working directory or full path to them should be specified in *filename* variable.

Filtration of events by flags according to the user needs can be done with *SphinX\_select.pro* IDL routine. Preparation of lightcurves and spectra, with which most users want to work, is possible using IDL routines *SphinX\_lightcurve.pro* and spectra *SphinX\_spectrum.pro*. All these routines can be downloaded from SphinX level-1 data catalogue

## http://156.17.94.1/SphinX\_11\_catalogue/SphinX\_cat\_main.html

under the link *SphinX level-1 data description, calibration info and software.* In the same place there are short sample IDL programs *SphinX\_test\_lc.pro, SphinX\_test\_spectrum.pro* which read SphinX level-1 event FITS, filter the events and produce lightcurves and spectra. In *SphinX\_test\_spectrum.pro* it is also shown how to read calibration information. Execution of these programs in IDL produces plots like the ones shown in Figure 3 and Figure 4.

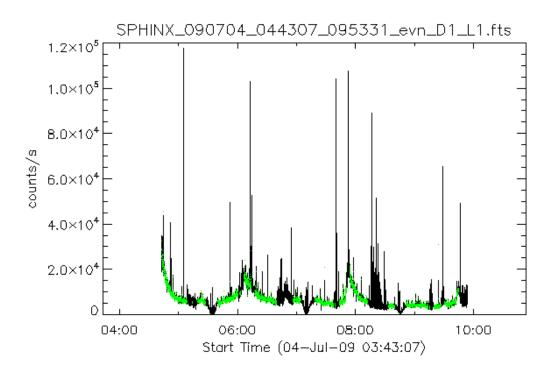


Figure 3. Example of SphinX lightcurves. Raw lightcurve obtained from events in SphinX FITS file SPHINX\_090704\_044307\_095331\_evn\_D1\_L1.fts is shown in black. A lightcurve binned only from good solar events (X-ray photons) is shown in green.

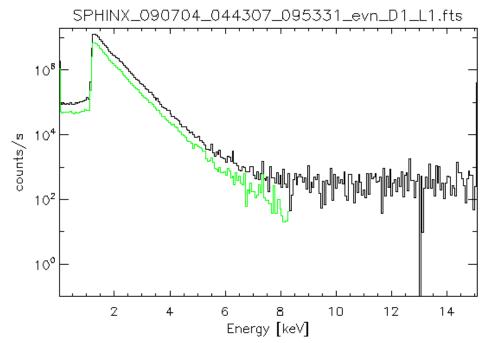


Figure 4. Example of SphinX spectrum. Raw spectrum obtained from events in SphinX FITS file SPHINX\_090704\_044307\_095331\_evn\_D1\_L1.fts is shown in black. A spectrum binned only from good solar events (X-ray photons) is shown in green.

## **APPENDIX 1 - SphinX FITS headers format**

#### **Event FITS Primary Header**

т SIMPLE = BITPIX = 16 NAXIS = 0 EXTEND = т DATE = '2011-03-17T20:15:13' ORIGIN = 'PAS SRC SPD Wroclaw' SATELITE= 'CORONAS-Photon' OBSERVER= 'Unknown ' TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' OBJECT = 'Sun ' DATE OBS= '2009-07-04T04:43:07.547' DATE END= '2009-07-04T09:53:31.358' TIMESYS = '1979-01-01T00:00:00' TIMEUNIT= 'd AUTHOR = 'sphinx\_fits2.pro' 193.472441441 RA = DEC = -6.54803262526  $RA_NOM =$ 193.472441441 DEC\_NOM = -6.54803262526 EQUINOX = 2000.00 RADECSYS= 'FK5 . .

/Written by IDL: Fri May 02 16:43:31 2008 /Number of bits per data pixel /channel /Number of data axes /File contains extensions /File creation date (YYYY-MM-DDThh:mm:ss UTC) /Origin of the file /Satellite name /Usually the name of the user who generated /Name of the Telescope or Mission /Name of the instrument /Object being observed /nominal U.T. date when integration of this /nominal U.T. date when integration of this /Reference time in YYYY MM DD hh:mm:ss /Unit for TIMEZERO, TSTARTI and TSTOPI /Program name that produced this file /Source right ascension in degrees /Source declination in degrees /r.a. nominal pointing in in degrees /dec. nominal pointing in in degrees /Equinox of celestial coordinate system /Coordinate frame used for equinox

TIMVERSN= 'OGIP XENDIAN = 'BIG ' VERSION = '1.0 ' END

#### EVENTS Extension Header

**XTENSION= 'BINTABLE'** BITPIX = 8 NAXIS = 2 NAXIS1 = 24 NAXIS2 = 5511812 PCOUNT = 0 GCOUNT = 1 TFIELDS = 6 COMMENT COMMENT \*\*\* End of mandatory fields \*\*\* COMMENT DATE = '2011-03-17T20:15:13' ORIGIN = 'PAS SRC SPD Wroclaw' SATELITE= 'CORONAS-Photon' OBSERVER= 'Unknown ' AUTHOR = 'sphinx\_fits2.pro' TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' OBJECT = 'Sun ' DATE\_OBS= '2009-07-04T04:43:07.547' DATE\_END= '2009-07-04T09:53:31.358' COMMENT -----EQUINOX = 2000.00 RADECSYS= 'FK5 ' RA = 193.472441441 DEC = -6.54803262526 RA\_NOM = 193.472441441 DEC\_NOM = -6.54803262526 COMMENT COMMENT \*\*\* Column names \*\*\* COMMENT TTYPE1 = 'TIME ' TTYPE2 = 'PHA ' TTYPE3 = 'PI ' TTYPE4 = 'ENERGY ' TTYPE5 = 'FLAG ' TTYPE6 = 'NRM ' COMMENT COMMENT \*\*\* Column formats \*\*\* COMMENT

/93-003' /OGIP memo number where the convention used /Byte order /File format version number

/Binary table written by MWRFITS v1.11 /Required value /Required value /Number of bytes per row /Number of rows /Normally 0 (no varying arrays) /Required value /Number of columns in table

/File creation date (YYYY-MM-DDThh:mm:ss UTC)
/Origin of the file
/Satellite name
/Usually the name of the user who generated
/Name of program that produced this file
/Name of the Telescope or Mission
/Name of the instrument
/Object being observed
/UTC start time of observation
/UTC end time of observation

/Equinox of celestial coordinate system /Coordinate frame used for equinox /[deg] RA of target /[deg] Dec of target /[deg] RA of nominal boresight /[deg] Dec of nominal boresight

| | | | | TFORM1 = 'D / TFORM2 = 'I / TFORM3 = 'I 1 TFORM4 = 'E 1 TFORM5 = 'J / TFORM6 = 'JCOMMENT COMMENT \*\*\* Column units \*\*\* COMMENT TUNIT1 = 's 1 TUNIT2 = 'chan ' 1 TUNIT3 = 'chan ' 1 TUNIT4 = 'keV ' 1 TUNIT5 = ' ' 1 TUNIT6 = ' 1 COMMENT ----- CHANNEL /ENERGY conversion info -----COMMENT from BESSY calibration it follows that nominally COMMENT ENERGY = 0.059217000\*PHA + 0.0050194827 COMMENT PI= fix(1. + 255. \* ENERGY/15.105547) COMMENT ----- SPHINX FLAG VALUES AND MEANING -----COMMENT BAD\_DGI\_TIME = 2L^0 - Bad SphinX times for the entire COMMENT IRESET =  $2L^{1}$  - Instrument reset COMMENT HOT = 2L^2 - Increased detector temperature COMMENT NN\_BAD = 2L^3 - upset of unknown origin COMMENT SHADOW = 2L^4 - instrument in shadow COMMENT X\_NIGHT = 2L^5 - X night COMMENT OPT\_NIGHT = 2L^6 - Optical night COMMENT RB = 2L^7 - Radiation belts COMMENT NRB = 2L^8 - RB - the north oval COMMENT SRB = 2L^9 - RB - the south oval COMMENT SAA = 2L^10 - SAA COMMENT PARTICLE = 2L^11 - Most likely particle event COMMENT HIPHA = 2L^12 - Event with exceptionally high PHA COMMENT NONGTI = 2L^13 - Event outside GTI COMMENT BAD EVN TIME = 2L^14 - Bad Event time COMMENT More detail flag description is in sphinx\_flag\_routines.pro COMMENT ----- TIME SYSTEM -----TIMESYS = '1979-01-01T00:00:00' /Reference time MJDREF in YYYY MM DD hh:mm:ss TIME\_UNI= 1 1 . . TIMEUNIT= 'd /Unit for TIMEZERO, TSTARTI and TSTOPI TIMEREF = 'LOCAL ' /Reference frame for the times LOCAL=satellite MJDREF = 43874.000 /TIMESYS in MJD (d) TIMEZERO= 11142.0000000 /Start day of the first bin rel to TIMESYS TSTARTI = 11142 /Integer portion of start time rel to TIMESYS TSTARTF = 0.1966151286069 /Fractional portion of start time TSTOPI = 11142 /Integer portion of stop time rel to TIMESYS TSTOPF = 0.4121684955098 /Fractional portion of stop time TASSIGN = 'SATELLITE' /Place of time assignment TIERRELA= 5.000000E-007 /Relative time error [s]

TIERABSO= 0.00100000 ONTIME = 1170.3904103795466 TELAPSE = 18623.8109003658464 EXPOSURE= 1144.83609813 LIVETIME= 1144.83609813 CLOCKCOR= 1 COMMENT absTime[i] = mjd2any(MJDREF + TIMEZERO) + TIME[i] COMMENT ------E MIN = 0.00000 E\_MAX = 15.1352 EUNIT = 'keV ' DETCHANS= 256 CHANTYPE= 'PI ' COMMENT -----VERSION = '1.0 ' EXTNAME = 'EVENTS ' HDUCLASS= 'OGIP ' HDUCLAS1= 'EVENTS ' HDUCLAS2= 'ALL ' HDUVERS = '1.2 ' TIMVERSN= 'OGIP COMMENT ------GROUPING= 0 DETNAM = 'D1 ' GEOAREA = 0.196350000000 FILTER = 'AIMyAIBe' RESPFILE= ' COMMENT COMMENT \*\*\* SphinX specyfic keywords \*\*\* COMMENT AMPNAME = 'A OBS ID = 'DS1 ' END

#### **EXPOSURE** Extension Header

**XTENSION= 'BINTABLE'** BITPIX = 8 NAXIS = 2 NAXIS1 = 24 NAXIS2 = 21910 PCOUNT = 0 GCOUNT = 1 TFIELDS = 3 COMMENT COMMENT \*\*\* End of mandatory fields \*\*\* COMMENT DATE = '2011-03-17T20:15:13'

/Absolute time error [s] baryc corr. etc ??? /sum of all Good Time Intervals - GTIs /Elapsed time in seconds between TSTART and TSTO /Integration time, corrected for deadtime/gaps /Livetime /Clock Correction to UT

/lower energy boudary /higher energy boundary /Energy units /Total number of detector channels available /Channels assigned by detector electronics

/File format version /Extension Name /File conforms to OGIP /GSFC convention /Extension contains Events /extension contains all events detected /File conforms to this version of OGIP /93-003' /OGIP memo number where the convention used

/No grouping of data has been defined /Detector name /Detector area [cm^2] /Filter used /SRM matrix

/Detector amplifier name /Observation id code

/Binary table written by MWRFITS v1.11 /Required value /Required value /Number of bytes per row /Number of rows /Normally 0 (no varying arrays) /Required value /Number of columns in table

/File creation date (YYYY-MM-DDThh:mm:ss UTC)

ORIGIN = 'PAS SRC SPD Wroclaw' SATELITE= 'CORONAS-Photon' OBSERVER= 'Unknown ' AUTHOR = 'SpX2OGIP.pro' TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' OBJECT = 'Sun ' DATE\_OBS= '2009-07-04T04:43:07.547' DATE END= '2009-07-04T09:53:31.358' COMMENT ------EQUINOX = 2000.00 RADECSYS= 'FK5 ' RA = 193.472441441 DEC = -6.54803262526 193.472441441 RA\_NOM = DEC\_NOM = -6.54803262526 COMMENT COMMENT \*\*\* Column names \*\*\* COMMENT TTYPE1 = 'TSTART ' TTYPE2 = 'TSTOP ' TTYPE3 = 'FRACEXP ' COMMENT COMMENT \*\*\* Column formats \*\*\* COMMENT TFORM1 = 'D ' TFORM2 = 'D ' . TFORM3 = 'D COMMENT \*\*\* Column units \*\*\* TUNIT1 = 's ' TUNIT2 = 's TUNIT3 = 'FRACTION' COMMENT ----- TIME SYSTEM ------TIMESYS = '1979-01-01T00:00:00' TIME UNI= 1 TIMEUNIT= 'd TIMEREF = 'LOCAL ' MJDREF = 43874.000 TIMEZERO= 11142.0000000 TSTARTI = 11142 TSTARTF = 0.1966151286069 TSTOPI = 11142 TSTOPF = 0.4121684955098 TASSIGN = 'SATELLITE' TIERRELA= 5.000000E-007 TIERABSO= 0.00100000 ONTIME = 1170.3904103795466 TELAPSE = 18623.8109003658464 EXPOSURE= 1144.83609813

/Origin of the file /Satellite name /Usually the name of the user who generated /Name of program that produced this file /Name of the Telescope or Mission /Name of the instrument /Object being observed /UTC start time of observation /UTC end time of observation

/Equinox of celestial coordinate system /Coordinate frame used for equinox /[deg] RA of target /[deg] Dec of target /[deg] RA of nominal boresight /[deg] Dec of nominal boresight

/

1

1

1

/ /Unit for TIMEZERO, TSTARTI and TSTOPI /Reference frame for the times LOCAL=satellite /TIMESYS in MJD (d) /Start day of the first bin rel to TIMESYS /Integer portion of start time rel to TIMESYS /Fractional portion of start time /Integer portion of stop time rel to TIMESYS /Fractional portion of stop time /Place of time assignment /Relative time error [s] /Absolute time error [s] baryc corr. etc ??? /sum of all Good Time Intervals - GTIs /Elapsed time in seconds between TSTART and TSTOP /Integration time, corrected for deadtime/gaps

/Reference time MJDREF in YYYY MM DD hh:mm:ss fo

LIVETIME= 1144.83609813 CLOCKCOR= 1 COMMENT absTime[i] = mjd2any(MJDREF + TIMEZERO) + TIME[i] COMMENT -----COMMENT FRACEXP was calculated as a ratio OCR COMMENT ICR was calculated by solving the following equation COMMENT OCR = ICR\*exp(-dt\*ICR) with constant dt = 5.24d-6 [s] COMMENT ------EUNIT = 'keV ' VERSION = '1.0 ' EXTNAME = 'EXPOSURE' HDUCLASS= 'OGIP ' HDUCLAS1= 'EXPOSURE' HDUVERS = '1.2 ' TIMVERSN= 'OGIP COMMENT -----GROUPING= 0 . DETNAM = 'D1 GEOAREA = 0.196350000000 FILTER = 'AIMyAIBe' COMMENT COMMENT \*\*\* SphinX specyfic keywords \*\*\* COMMENT AMPNAME = 'A OBS\_ID = 'DS1 ' END

#### **GTI** Extension Header

XTENSION= 'BINTABLE' BITPIX = 8 NAXIS = 2 NAXIS1 = 16 NAXIS2 = 21910 PCOUNT = 0 GCOUNT = 1 TFIELDS = 2 COMMENT COMMENT \*\*\* End of mandatory fields \*\*\* COMMENT DATE = '2011-03-17T20:15:13' ORIGIN = 'PAS SRC SPD Wroclaw' SATELITE= 'CORONAS-Photon' OBSERVER= 'Unknown ' AUTHOR = 'SpX2OGIP.pro' TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' OBJECT = 'Sun '

/Livetime /Clock Correction to UT /ICR, where /Energy units /File format version /Extension Name /File conforms to OGIP /GSFC convention /Extension contains exposures /File conforms to this version of OGIP /93-003' /OGIP memo number where the convention used

/No grouping of data has been defined /Detector name /Detector area [cm^2] /Filter used

/Detector amplifier name /Observation id code

/Binary table written by MWRFITS v1.11 /Required value /Required value /Number of bytes per row /Number of rows /Normally 0 (no varying arrays) /Required value /Number of columns in table

/File creation date (YYYY-MM-DDThh:mm:ss UTC) /Origin of the file /Satellite name /Usually the name of the user who generated /Name of program that produced this file /Name of the Telescope or Mission /Name of the instrument /Object being observed

DATE\_OBS= '2009-07-04T04:43:07.547' DATE\_END= '2009-07-04T09:53:31.358' COMMENT -----EQUINOX = 2000.00 RADECSYS= 'FK5 ' RA = 193.472441441 DEC = -6.54803262526 RA\_NOM = 193.472441441 DEC NOM = -6.54803262526 COMMENT COMMENT \*\*\* Column names \*\*\* COMMENT TTYPE1 = 'TSTART ' TTYPE2 = 'TSTOP ' COMMENT COMMENT \*\*\* Column formats \*\*\* COMMENT TFORM1 = 'D . . TFORM2 = 'D ' COMMENT \*\*\* Column units \*\*\* TUNIT1 = 's TUNIT2 = 's COMMENT ----- TIME SYSTEM ------TIMESYS = '1979-01-01T00:00:00' TIME UNI= 1 TIMEUNIT= 'd ' TIMEREF = 'LOCAL ' MJDREF = 43874.000 TIMEZERO= 11142.0000000 TSTARTI = 11142 TSTARTF = 0.1966151286069 TSTOPI = 11142 TSTOPF = 0.4121684955098 TASSIGN = 'SATELLITE' TIERRELA= 5.000000E-007 TIERABSO= 0.00100000 ONTIME = 1170.3904103795466 TELAPSE = 18623.8109003658464 EXPOSURE= 1144.83609813 LIVETIME= 1144.83609813 CLOCKCOR= 1 COMMENT absTime[i] = mjd2any(MJDREF + TIMEZERO) + TIME[i] COMMENT -----EUNIT = 'keV ' VERSION = '1.0 ' EXTNAME = 'GTI ' HDUCLASS= 'OGIP ' HDUCLAS1= 'GTI ' HDUCLAS2= 'STANDARD'

/UTC start time of observation /UTC end time of observation /Equinox of celestial coordinate system /Coordinate frame used for equinox /[deg] RA of target /[deg] Dec of target /[deg] RA of nominal boresight /[deg] Dec of nominal boresight 1 1 /Reference time MJDREF in YYYY MM DD hh:mm:ss fo /Unit for TIMEZERO, TSTARTI and TSTOPI /Reference frame for the times LOCAL=satellite /TIMESYS in MJD (d) /Start day of the first bin rel to TIMESYS /Integer portion of start time rel to TIMESYS /Fractional portion of start time /Integer portion of stop time rel to TIMESYS /Fractional portion of /Place of time assignment /Relative time error [s] /Absolute time error [s] baryc corr. etc ??? /sum of all Good Time Intervals - GTIs /Elapsed time in seconds between TSTART and TSTOP /Integration time, corrected for deadtime/gaps /Livetime /Clock Correction to UT

/Energy units /File format version /Extension Name /File conforms to OGIP /GSFC convention /Extension contains good time intervals /Standard GTIs HDUVERS = '1.2 . TIMVERSN= 'OGIP COMMENT -----GROUPING= 0 DETNAM = 'D1 ' GEOAREA = 0.196350000000 FILTER = 'AIMyAIBe' COMMENT COMMENT \*\*\* SphinX specyfic keywords \*\*\* COMMENT AMPNAME = 'A , OBS ID = 'DS1 ' END

#### **Response FITS Primary Header**

Т

SIMPLE =

/File conforms to this version of OGIP /93-003' /OGIP memo number where the convention used

/No grouping of data has been defined /Detector name /Detector area [cm^2] /Filter used

/Detector amplifier name /Observation id code

/ file does conform to FITS standard / number of bits per data pixel

BITPIX = 16 NAXIS = 0 / number of data axes EXTEND = Т / FITS dataset may contain extensions COMMENT FITS (Flexible Image Transport System) format defined in Astronomy and COMMENT Astrophysics Supplement Series v44/p363, v44 /p371, v73 /p359, v73 /p365. COMMENT Contact the NASA Science Office of Standards and Technology for the COMMENT FITS Definition document #100 and other FITS information. DATE = '03/03/11' / FITS file creation date (dd/mm/yy) CONTENT = 'CONV RESPONSE MATRIX' RMFVERSN= 'R6 . / Redistribution version number ORIGIN = 'PAS SRC SPD Wroclaw' /Origin of the file SATELITE= 'CORONAS-Photon' /Satellite name TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' /Name of the instrument OBJECT = 'Sun /Object being observed ORIGIN = 'SRC PAS SPD' CREATOR = 'SG' / Creator END

#### MATRIX Extension Header

XTENSION= 'BINTABLE'		
BITPIX =	8	
NAXIS =	2	
NAXIS1 =	2074	
NAXIS2 =	281	
PCOUNT =	0	
GCOUNT =	1	
TFIELDS =	6	
TTYPE1 = 'ENERG_LO'		

/ (convolved) RMF and EBOUNDS extension /Name of the Telescope or Mission / Organization which created this file

/ binary table extension / 8-bit bytes / 2-dimensional binary table / width of table in bytes / number of rows in table 1 / one data group (required keyword) / number of fields in each row

/ label for field 1

TFORM1 = 'D TUNIT1 = 'keV ' TTYPE2 = 'ENERG\_HI' TFORM2 = 'D TUNIT2 = 'keV . TTYPE3 = 'N\_GRP ' TFORM3 = 'I . TTYPE4 = 'F\_CHAN ' TFORM4 = 'JTTYPE5 = 'N\_CHAN ' TFORM5 = 'JTTYPE6 = 'MATRIX ' TFORM6 = '256D ' EXTNAME = 'MATRIX ' CHANTYPE= 'PI ' DETCHANS= 256 LO\_THRES= 0.00000E+00 EFFAREA = 1.00000E+00 RMFVERSN= 'R6 SATELITE= 'CORONAS-Photon' TELESCOP= 'SPHINX ' INSTRUME= 'SPHINX ' OBJECT = 'Sun DETNAM = 'D1 ' FILTER = 'AIMyAIBe' CCLS0001= 'CPF CCNM0001= 'MATRIX ' CBD10001= 'ENERGY(0-15)keV' CBD20001= 'CHAN (0- 255)' CDTP0001= 'DATA ' CVSD0001= '03/03/11 CVST0001= '00:00:00' HDUCLASS= 'OGIP ' HDUCLAS1= 'RESPONSE' HDUCLAS2= 'RSP MATRIX' HDUCLAS3= 'FULL ' HDUVERS1= '1.0.0 ' HDUVERS2= '1.1.0 ' CDES0001= 'SphinX DRM' END

#### EBOUNDS Extension Header

XTENSION= 'BINTABLE'			
BITPIX =	8		
NAXIS =	2		
NAXIS1 =	12		
NAXIS2 =	256		

/ data format of the field: 4-byte REAL / physical unit of field / label for field 2 / data format of the field: 4-byte REAL / physical unit of field / label for field 3 / data format of the field: 2-byte INTEGER / label for field 4 / data format of the field / label for field 5 / data format of the field / label for field 6 / data format of the field / name of this binary table extension / Whether channels corrected / Type of channel PHA/PI / Lower threshold for stored matrix / Area scaling factor / Redistribution version number /Satellite name /Name of the Telescope or Mission /Name of the instrument /Object being observed / / Standard / OGIP class of calibration file / OGIP codename for this type of cal file / Dataset parameter boundary / Dataset parameter boundary / OGIP type of dataset (DATA, TASK, etc)

/ General Comments to identify the dataset
/ binary table extension
/ 8-bit bytes
/ 2-dimensional binary table

/ Dataset validity start time (UTC)

/ Organisation which devised format

/ Version of the HDUCLAS1 format in use

/ Version of the HDUCLAS1 format in use

/ Extension contains inst response

/ The type of data stored

/ Nature of stored matrix

/ Dataset validity start time (UTC, of day CVSD)

- / width of table in bytes
- / number of rows in table

PCOUNT = 0 GCOUNT = 1 TFIELDS = 3 TTYPE1 = 'CHANNEL ' TFORM1 = 'J TTYPE2 = 'E\_MIN ' TFORM2 = 'E . TUNIT2 = 'keV ' TTYPE3 = 'E MAX ' TFORM3 = 'E TUNIT3 = 'keV . EXTNAME = 'EBOUNDS ' CHANTYPE= 'PI ' DETCHANS= 256 EFFAREA = 1.00000E+00 RMFVERSN= 'R6 . . TELESCOP= 'SphinX ' INSTRUME= 'SphinX ' DETNAM = 'D1 FILTER = 'AIMyAIBe' CLASS = 'Patterns 0-12' HDUCLASS= 'OGIP ' HDUCLAS1= 'RESPONSE' HDUCLAS2= 'EBOUNDS ' HDUVERS1= '1.0.0 ' HDUVERS2= '1.1.0 ' CCLS0001= 'CPF ' CCNM0001= 'MATRIX ' CBD10001= 'ENERGY(0-15)keV' CDTP0001= 'DATA ' CVSD0001= '01/01/11 CVST0001= '00:00:00' CDES0001= 'Sphinx DRM' END

/ size of special data area / one data group (required keyword) / number of fields in each row / label for field 1 / data format of the field: 2-byte INTEGER / label for field 2 / data format of the field: 4-byte REAL / physical unit of field / label for field 3 / data format of the field: 4-byte REAL / physical unit of field / name of this binary table extension / Whether channels corrected / Total number of detector channels / Area scaling factor / Redistribution version number / / / / Standard / X-ray Pattern Subset / Organisation which devised format / Extension contains inst response / The type of data stored / Version of the HDUCLAS1 format in use / Version of the HDUCLAS1 format in use / OGIP class of calibration file / OGIP codename for this type of cal file / Dataset parameter boundary / OGIP type of dataset (DATA, TASK, etc) / Dataset validity start time (UTC) / Dataset validity start time (UTC, of day CVSD) / General Comments to identify the dataset